

#### Plan of the talk

Numerical Weather Prediction (NWP)
 and Data Assimilation (DA)

Observations (in-situ and remote sensing)

Error Covariances and Ensemble DA

A posteriori diagnostics
 (observation-minus-background departures)



## Numerical Weather Prediction and Data Assimilation



## The two main ingredients of weather forecasting

What will be the weather tomorrow?

Bjerknes (1904):

In order to do a good forecast, we need to:

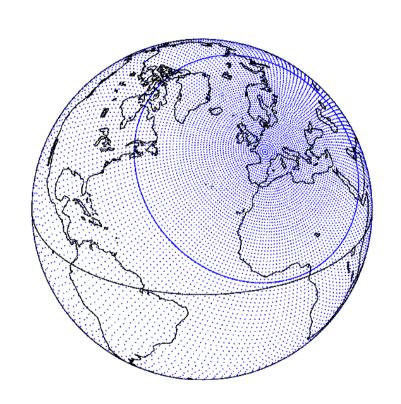
- know the atmospheric evolution laws (~ modeling);
- know the atmospheric state at initial time (~ data assimilation).

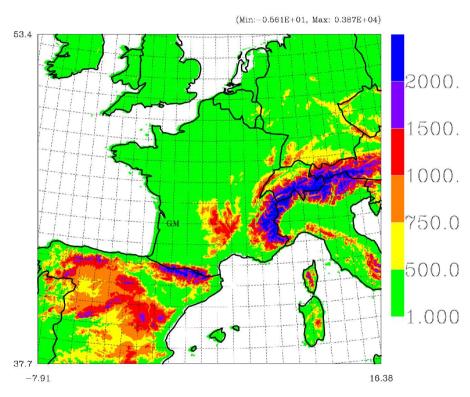


## Numerical Weather Prediction at Météo-France (in collaboration with e.g. ECMWF)

Global model (Arpège) : DX ~ 7-40 km

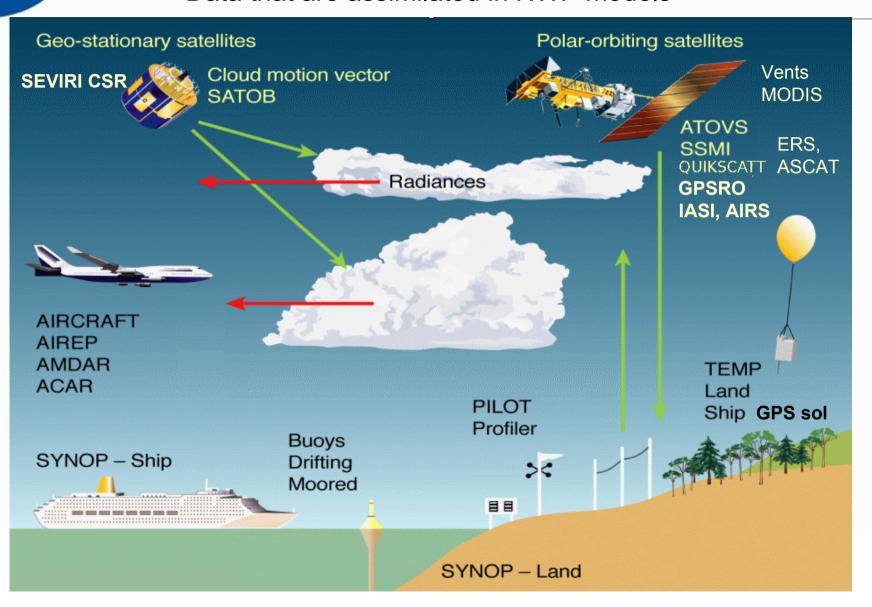
Arome:  $DX \sim 1.3 \text{ km}$ 



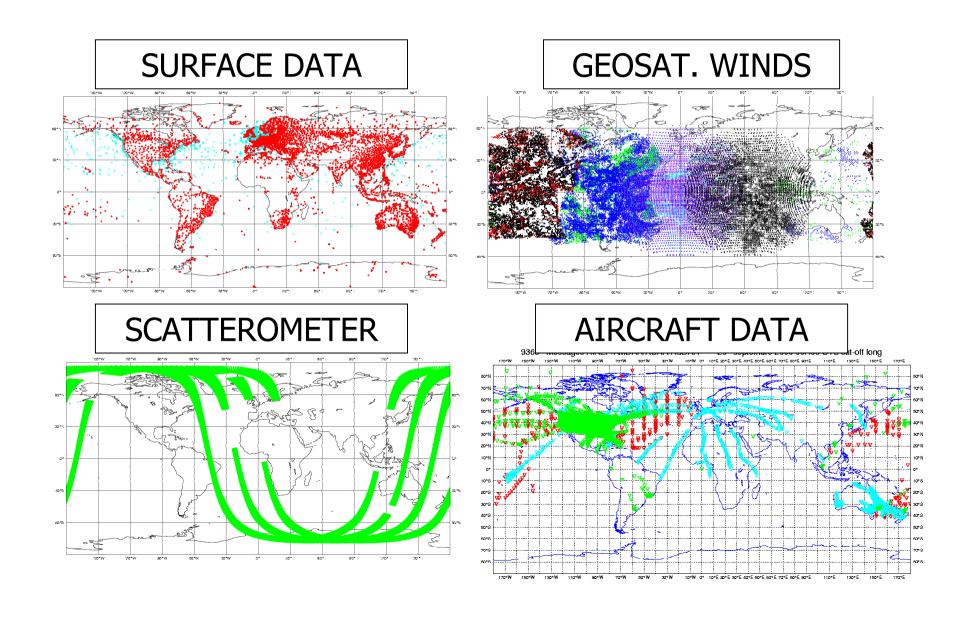


Equations of hydrodynamics and physical parametrizations (radiation, convection,...) to predict the evolution of temperature, wind, humidity, ...

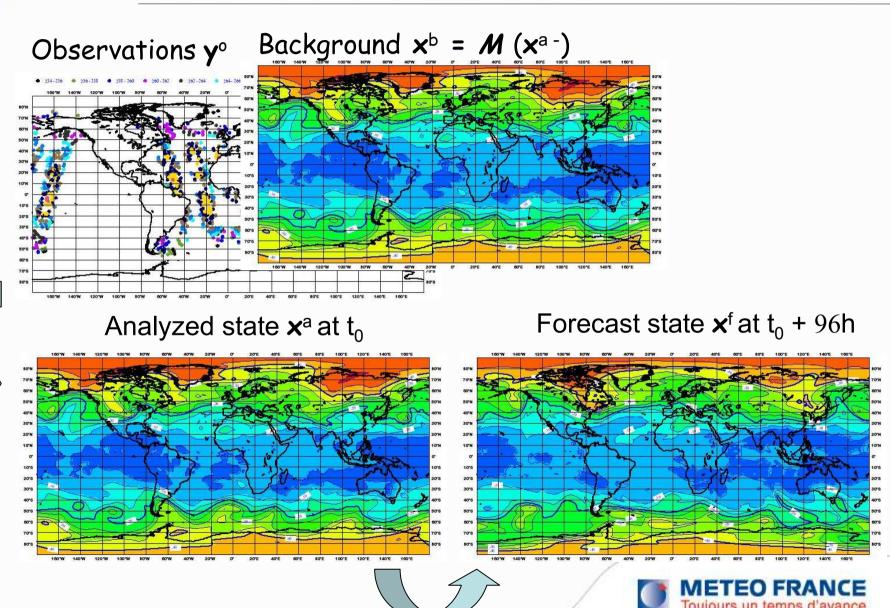
#### Data that are assimilated in NWP models



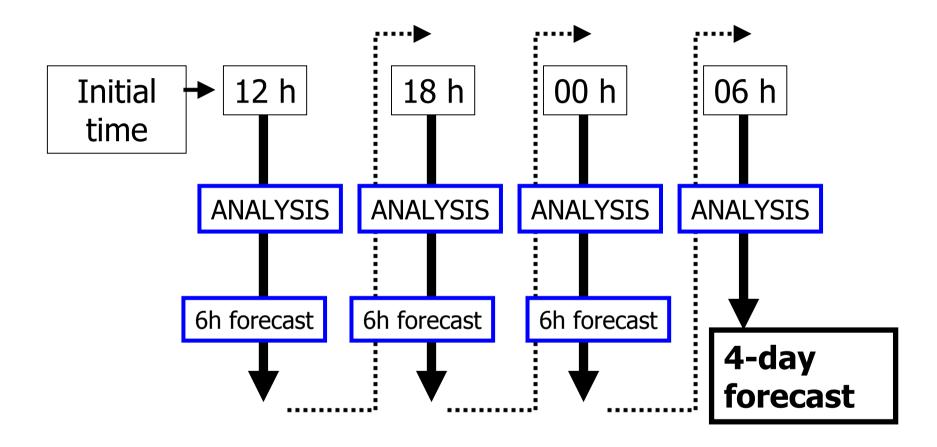
## Spatial coverage and density of observations



## Data assimilation for NWP: illustration



#### The data assimilation cycle



Memory of DA system is updated ∼ continuously

#### Linear estimation of model state (1)

- BLUE analysis equation : x<sup>a</sup> = (I-KH) x<sup>b</sup> + K y<sup>o</sup>
- H = observation operator = projection from model to observation space (e.g. spatial interpolation, radiative transfer, NWP model).
- K = observation weights :

$$K = BH^{T} (HBH^{T} + R)^{-1}$$
  
 $H K = (I + R (HBH^{T})^{-1})^{-1}$ 

- → ratio between background error covariances (matrix B) and observation error covariances (matrix R).
- ⇒ Accounts for relative accuracy of observations, and for spatial structures of background errors.



#### Linear estimation of model state (2)

Analysis increment equation :

$$x^a - x^b = K (y^o - H x^b)$$
  
 $\delta x = K \delta y$ 

Single-observation case (with uniform variances):

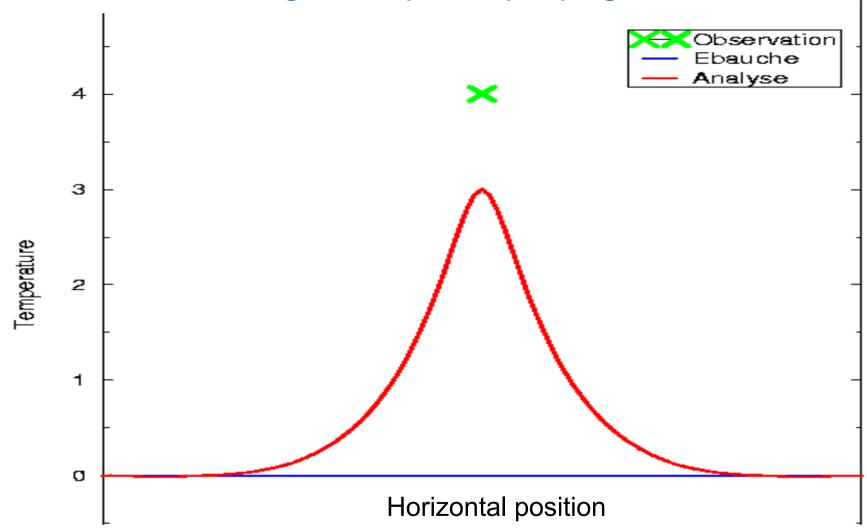
$$\delta x(j) = cor^{b}(i,j) 1/(1+(\sigma^{o}/\sigma^{b})^{2}) \delta y(i)$$

- ⇒ Filtering of observed information, as a function of obs/bkd error variance ratios.
- ⇒ Spatial propagation of observed information, as a function of background error correlations.



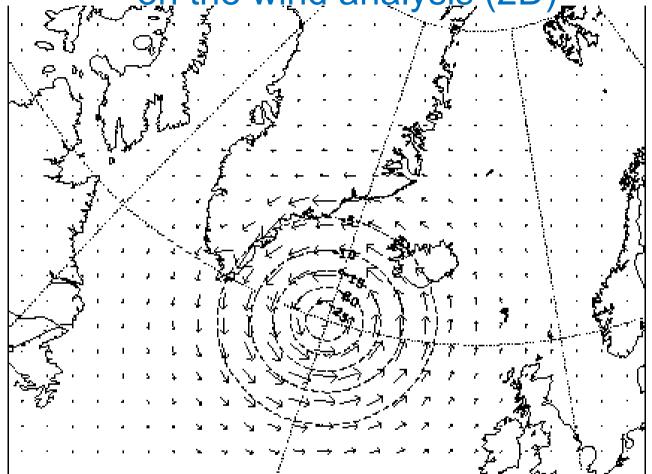


## Impact of one observation (1D): filtering and spatial propagation



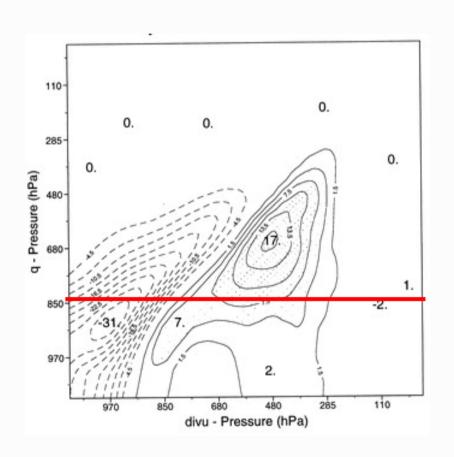
⇒ relative accuracies of observations and background, and characteristic spatial scales of bkd errors are accounted for.

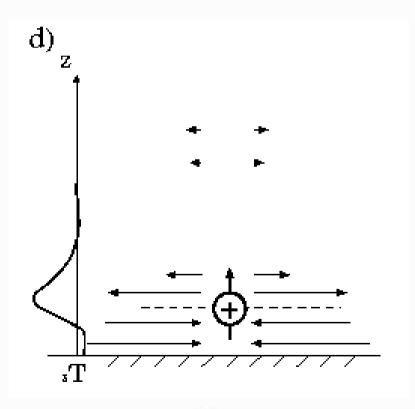
Impact of one surface pressure observation on the wind analysis (2D)\_\_\_\_\_



⇒ multivariate couplings (ex: pressure/wind) are also accounted for.

#### Divergence/humidity couplings





(Berre 2000, Montmerle et al 2006)

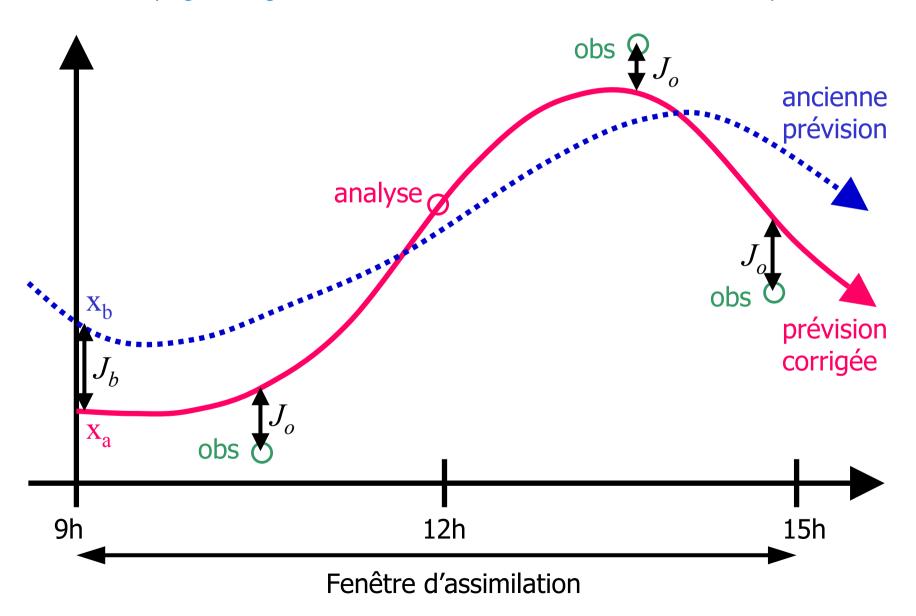
#### Linear estimation of model state (3)

- Size of B is huge: square of model size  $\sim (10^8)^2 \sim 10^{16}$ .
- ⇒ error covariances need to be estimated, simplified and modeled.
- Matrices too large to be inverted, but equivalent to minimize distance J(x<sup>a</sup>) to x<sup>b</sup> and y<sup>o</sup> (4D-Var) without explicit matrix inversions (e.g. Talagrand and Courtier 1987).
- Non linear features accounted for in calculation of departures between yo and H(xb), and in iterative applications of 4D-Var.



#### Principle of 4D-VAR assimilation

(e.g. Talagrand and Courtier 1987, Rabier et al 2000)



#### Implementation of 4D-Var

Analysis increment (BLUE equation) :

$$\delta x = x^a - x^b = K (y^o - H x^b) = K \delta y$$

but **K** is difficult to handle explicitly in a real size system.

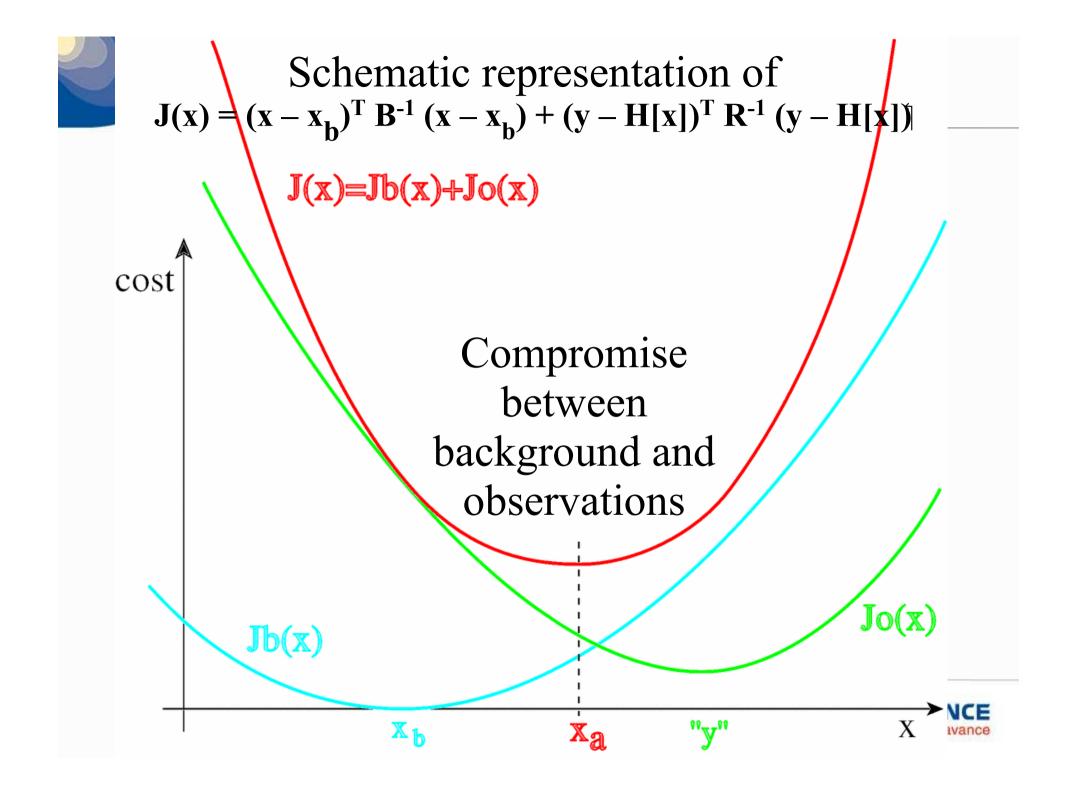
Variational formulation :

cost function : 
$$J(\delta \mathbf{x}) = \delta \mathbf{x}^T \mathbf{B}^{-1} \delta \mathbf{x} + (\mathbf{d} - \mathbf{H} \delta \mathbf{x})^T \mathbf{R}^{-1} (\mathbf{d} - \mathbf{H} \delta \mathbf{x})$$

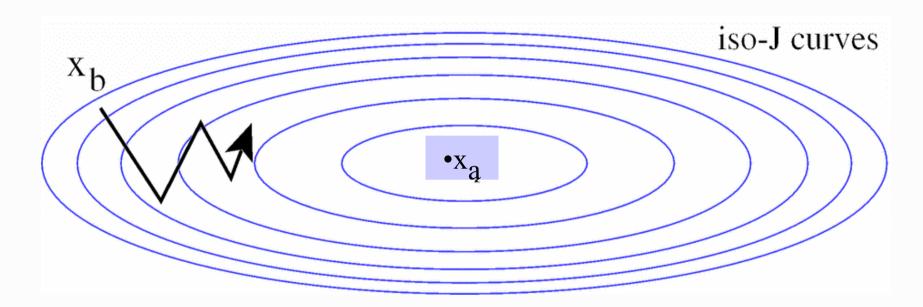
minimised when gradient  $J'(\delta \mathbf{x})=0$  (equivalent to BLUE).

- Computation of J': development and use of adjoint operators (transpose).
- Generalized observation operator H: includes NWP model M.
- Cost reduction : analysis increment δx can be computed at low resolution (Courtier, Thépaut et Hollingsworth, 1994)





#### Importance of preconditioning



- Some gradient directions have much larger amplitudes than others
  problem of "narrow valley" linked to the metric of x.
- Use a change of variable such as J becomes nearly "circular": much faster convergence.



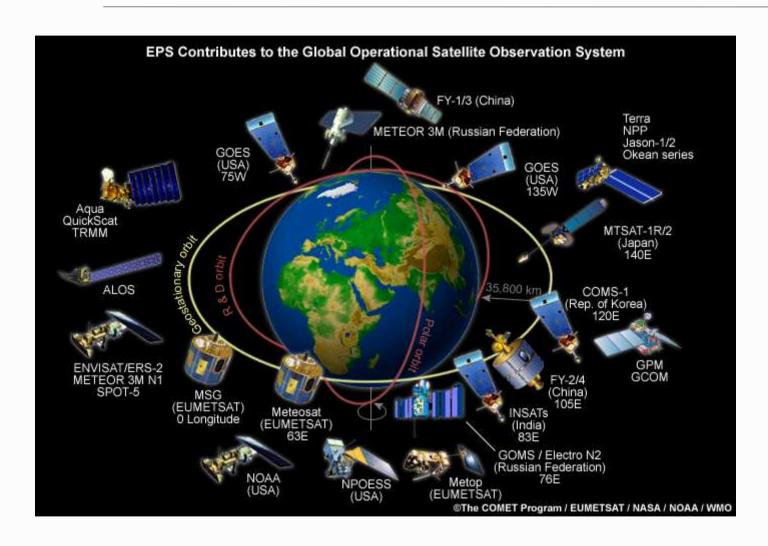
# In-situ observations and remote sensing data



## Observation networks in meteorology: in situ measurements



### Observation networks in meteorology: satellite data

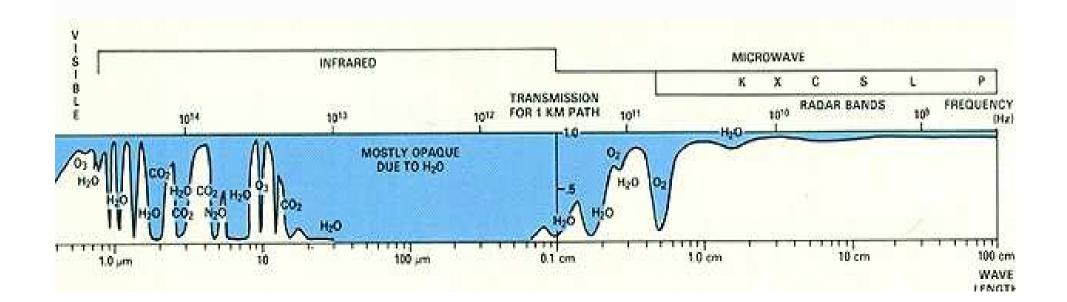


Constellation of polar orbiting or geostationary satellites



#### What is measured by satellite sensors?

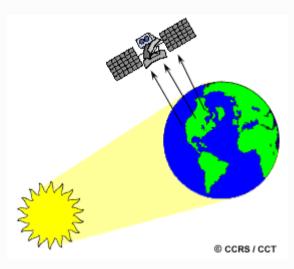
- Sensors do not measure directly atmospheric temperature and humidity, but electromagnetic radiation : brightness temperature or radiance.
- □ Depending on wave length (or frequency), information on gas concentration or physical properties (temperature or pressure or humidity) of atmosphere.
- Observations in atmospheric windows → information on surface.



#### What is measured by satellite sensors?

#### **Passive measures**

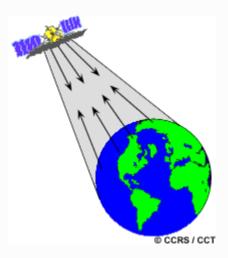
(no energy emitted from instrument)



Measures natural radiation emitted by Earth/Atmosphere from Sun origin

#### **Active measures**

(energy emitted from instrument)

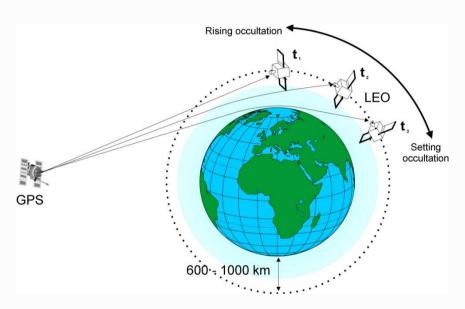


Radiation emitted by satellite and then reflected or diffused by Earth/Atmosphere

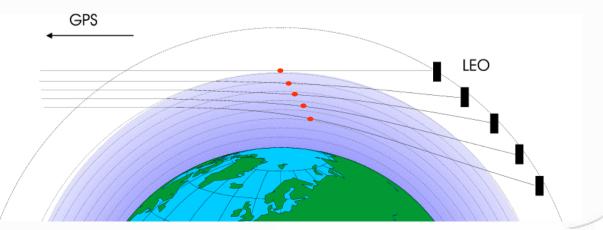


#### Example of active remote sensing

#### **GPS radio occultation:**



- Low-Earth Orbit satellites receive a signal from a GPS satellite.
- The signal passes through the atmosphere and gets refracted along the way.
- The magnitude of the refraction depends on temperature, moisture and pressure.
- The relative position of GPS and LEO changes over time => vertical scanning of the atmosphere.





#### GPS stations of Météo France: Toulouse and Guipavas





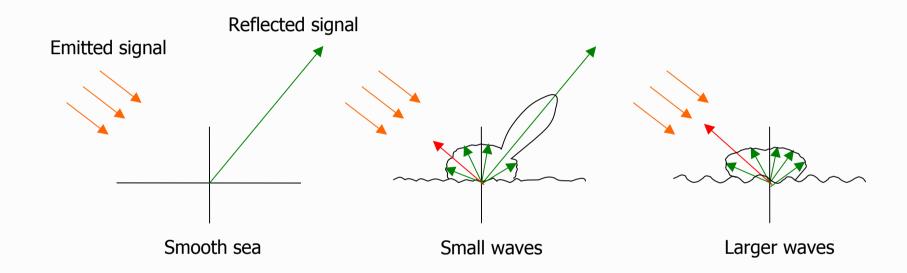
- Propagation of GPS signal is slowed by atmosphere (dry air and water vapour)
- More than 500 GPS stations over Europe provide an estimation of Zenith Total Delay (ZTD) in real time to weather centres.
  - All weather instrument
  - High temporal resolution



#### Scatterometers

They send out a microwave signal towards a sea target.

The fraction of energy returned to the satellite depends on wind speed and direction.



=> Measurements of near surface wind over the ocean, through backscattering of microwave signal reflected by waves.



#### Passive remote sensing

Only natural sources of radiation (sun, earth...) are involved, and the sensor is a simple receiver, « passive ».

Atmosphere in Parallel Plan, no diffusion, specular surface

 $T(p,\upsilon) = \varepsilon(p,\upsilon)Ts\tau + (1-\varepsilon(p,\upsilon))\tau T(\upsilon,\downarrow) + T(\upsilon,\uparrow)$ 

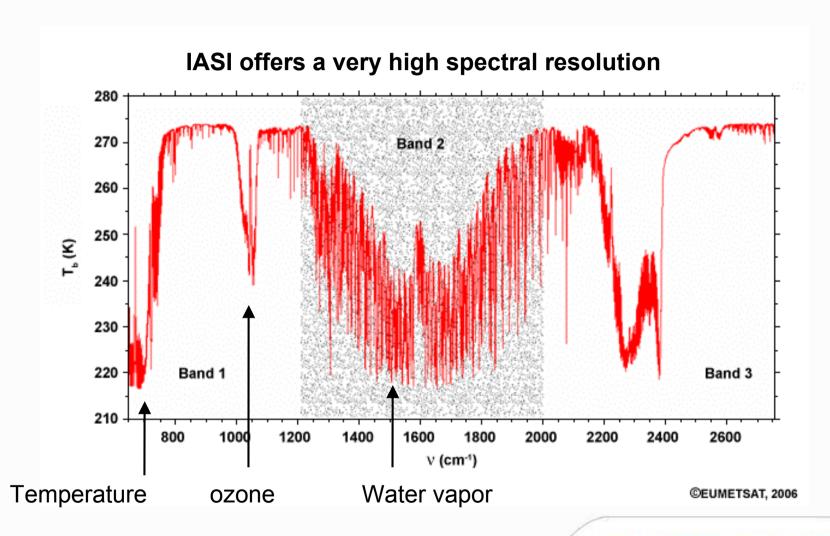
Energy source (1) Radiation 1 **Top of Atmosphere** Signal attenuated by atmosphere Surface (emissivity, temperature)

**Emissivity** 

Model outputs for RT: T, Q forecast or radiosondes or reanalyses

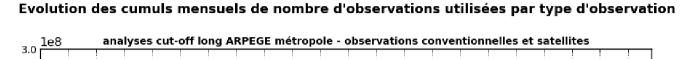


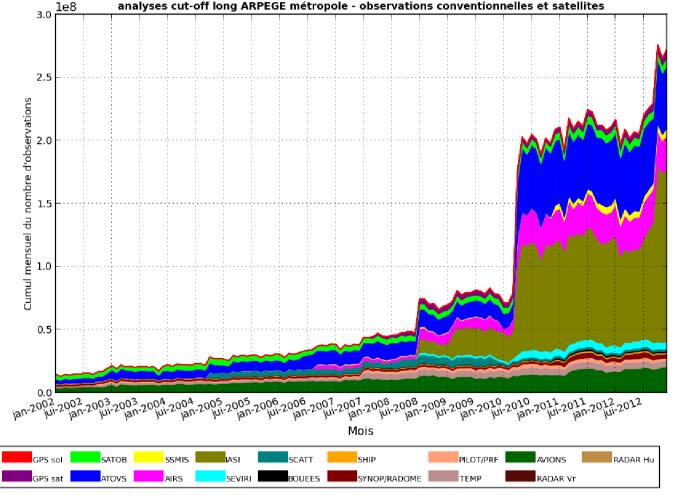
## IASI, infra-red interferometer developed by CNES and EUMETSAT





## Number of observations used in ARPEGE (global DA at Météo-France)







#### Radar network in France

• 24 radars (17 Doppler C-Band, every 15 minutes).

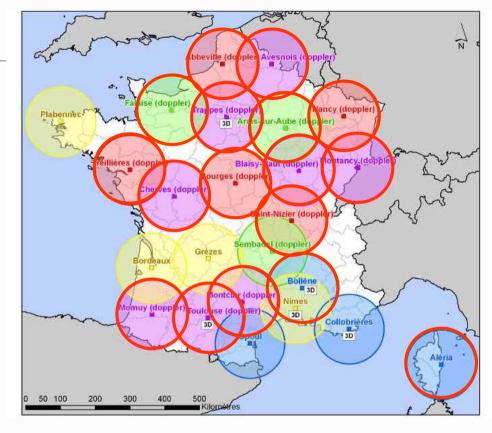
Doppler Radar

#### Observations

reflectivities Z (related to precipitation),

radial winds Vr (doppler effect : modified frequency of signal, when the target is moving => wind observation),

archived at 1km resolution.



# 10 km 100 km

## Observations assimilated as profiles in the model

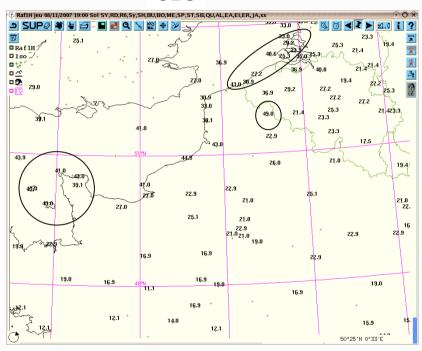
Pixel altitude is computed using a constant refractivity index along the path (effective radius approximation)

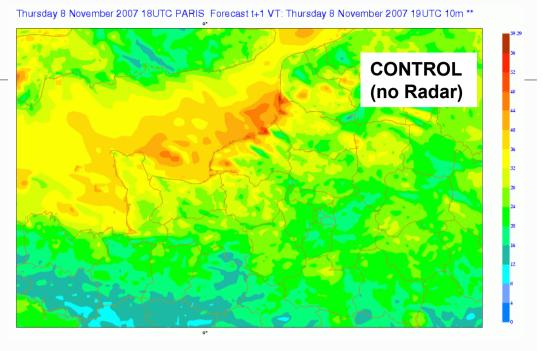


## Assimilation of radar radial winds

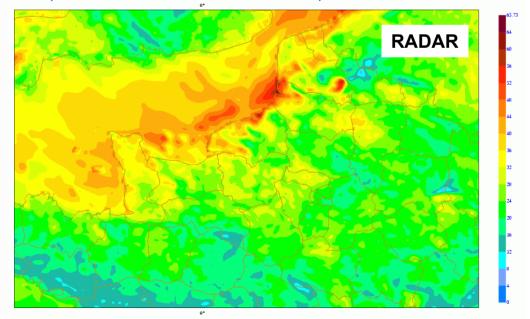
#### Wind gust at 10 m (kt) Forecast +1h (19 UTC)

#### OBS



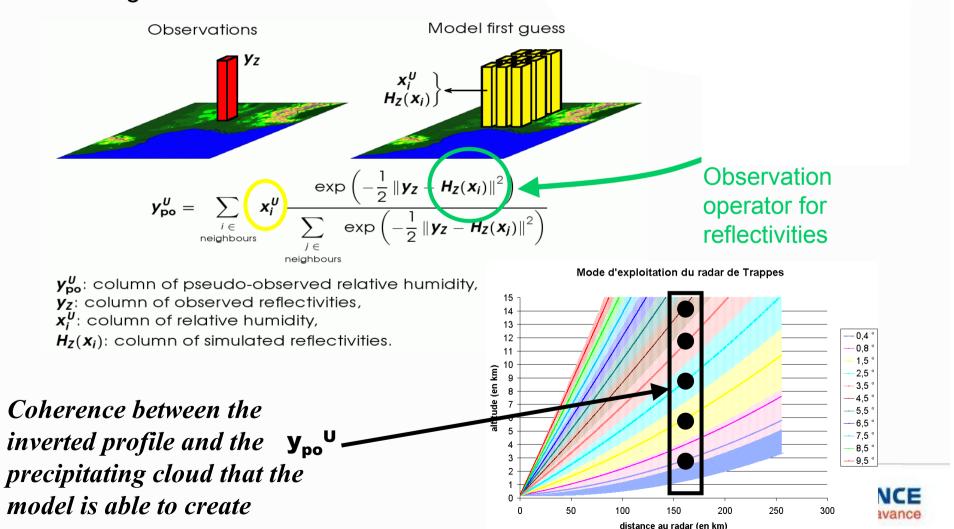


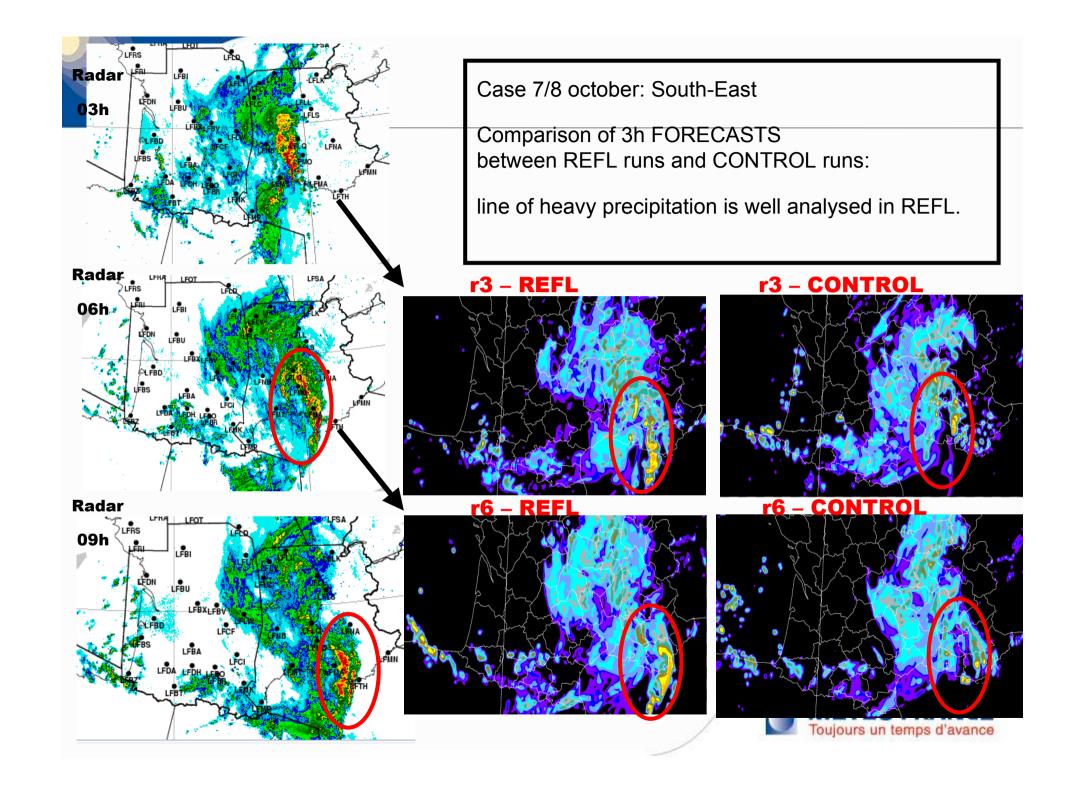




#### Inversion method of reflectivity profiles

## Caumont, 2006: use model profiles in the neighborhood of observations





#### 3. Error Covariances and

**Ensemble Data Assimilation** 



#### Observation weights and Error covariances

BLUE analysis equation :

$$x^a = (I-KH) x^b + K y^o$$

K = observation weights :

$$K = BH^{T} (HBH^{T} + R)^{-1}$$

⇒ Need to estimate B and R, before specifying them.



#### How can we estimate error covariances?

- The true atmospheric state is never exactly known.
- Use observation-minus-background departures :

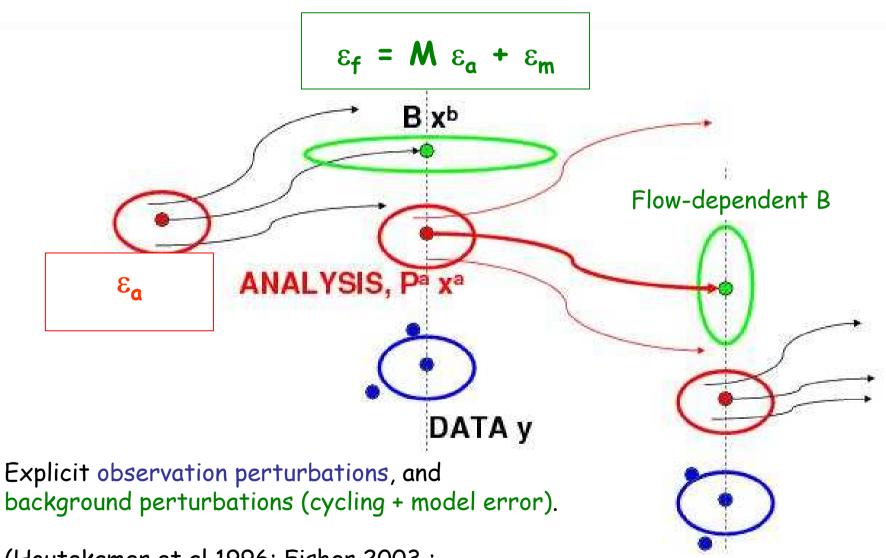
$$y_o - H x_b \sim (y_o - H x_t) + (H x_t - H x_b)$$
  
  $\sim e_o - H e_b$ 

to estimate some average features (e.g. variances, correlations) of R and B, using assumptions on spatial structures of errors.

 Use ensemble to simulate the error evolution and to estimate complex background error structures.



# Ensemble assimilation (EnDA = EnVar, EnKF, ...): simulation of the error evolution



(Houtekamer et al 1996; Fisher 2003; Ehrendorfer 2006; Berre et al 2006)

## Analysis error equation

Analysis state (BLUE, K = 4D-Var gain matrix):

$$x_a = (I-KH) x_b + K y_o$$

True state:

$$X_{\dagger} = (I-KH) X_{\dagger} + K H X_{\dagger}$$

Analysis error:

$$e_a = X_a - X_t$$

i.e.

$$e_a = (I-KH) e_b + K e_o$$



## Analysis perturbation equation

Perturbed analysis:

$$x'_{a} = (I-KH) x'_{b} + K y'_{o}$$

Unperturbed analysis:

$$x_a = (I-KH) x_b + K y_o$$

• Analysis perturbation :

$$\varepsilon_a = x'_a - x_a$$

$$\varepsilon_a = (I-KH) \varepsilon_b + K \varepsilon_a$$

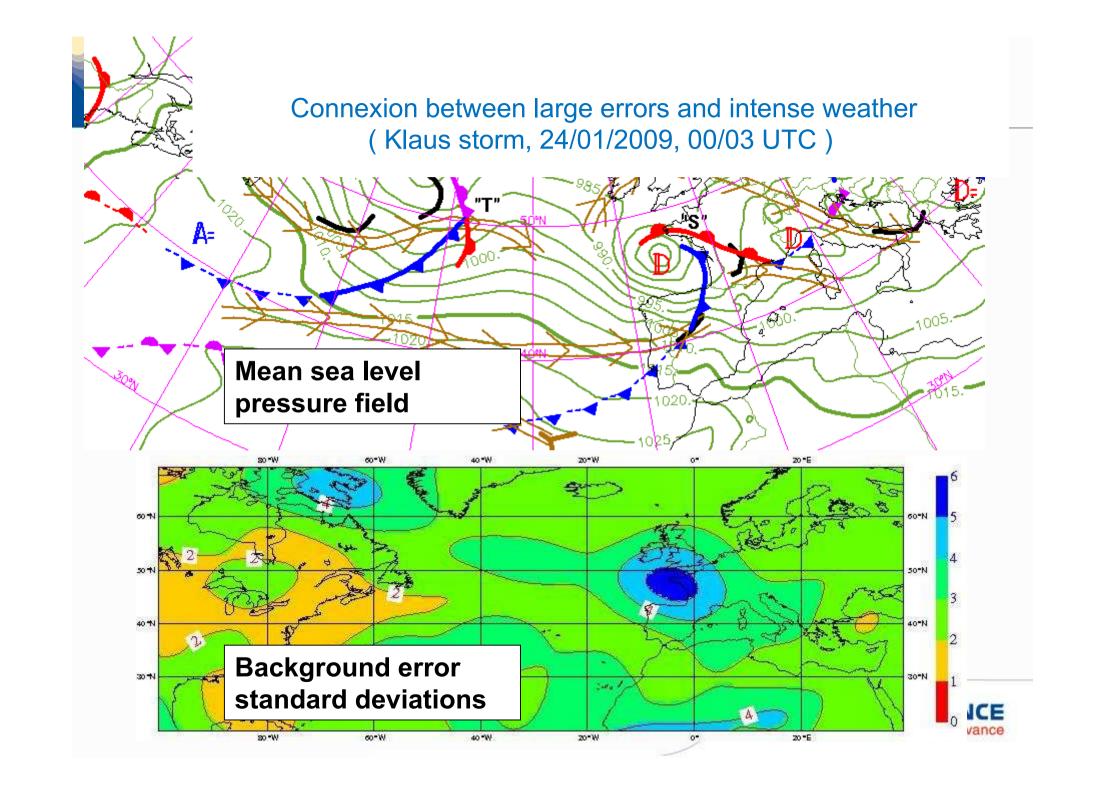
i.e.

=> Estimate 4D-Var errors by using perturbed inputs.

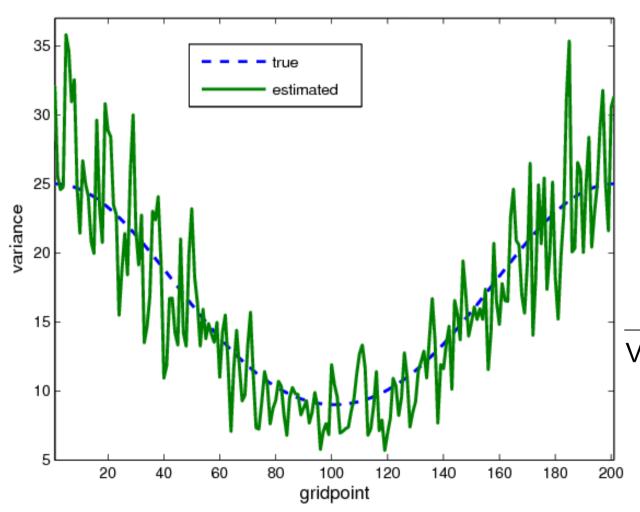
# Estimation of background error variances from ensemble spread

$$Var(e_b) = 1/(N-1) \sum_{n} (x'_b(n) - x'_b(mean))^2$$





# Spatial structure of sampling noise for variances (Raynaud et al 2009, Berre and Desroziers 2010)



$$\varepsilon_{b} = \mathbf{B}^{1/2} \, \eta$$
 $\eta \sim \mathcal{N}(0, \mathbf{I})$ 

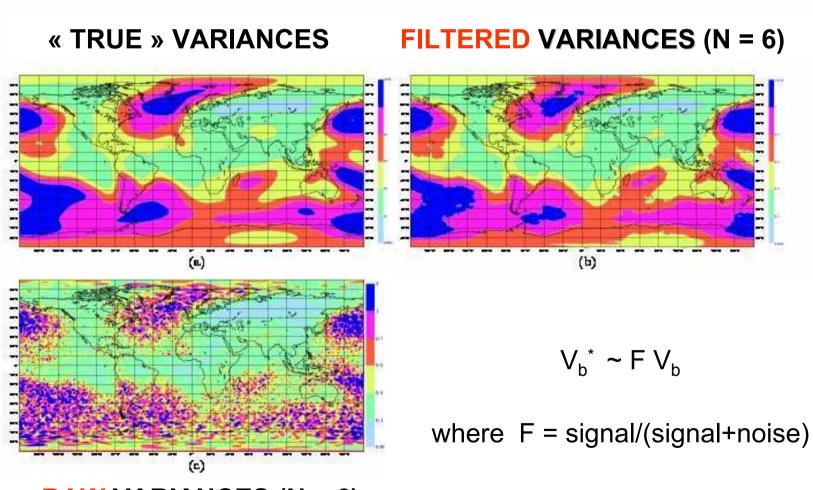
N = 50 members  
L(
$$\varepsilon_b$$
) = 200 km

$$V^{e} (V^{e})^{T} = 2/(N-1) \mathbf{B}^{*} \circ \mathbf{B}^{*}$$

⇒ Employ filtering in order to extract large scale signal, and remove small scale sampling noise.

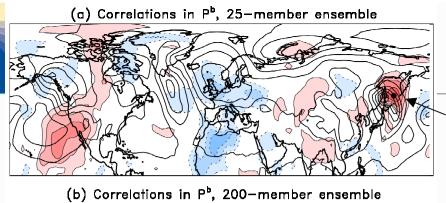


# "OPTIMIZED" SPATIAL FILTERING OF THE VARIANCE FIELD

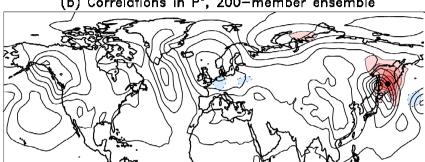


RAW VARIANCES (N = 6) (Berre et al 20

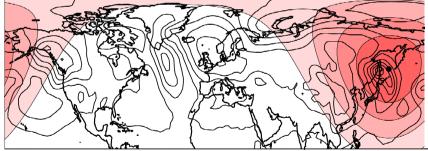
(Berre et al 2007,2010, Raynaud et al 2008,2009)



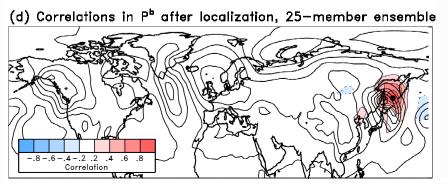
obs location



(c) Gaspari & Cohn correlation function



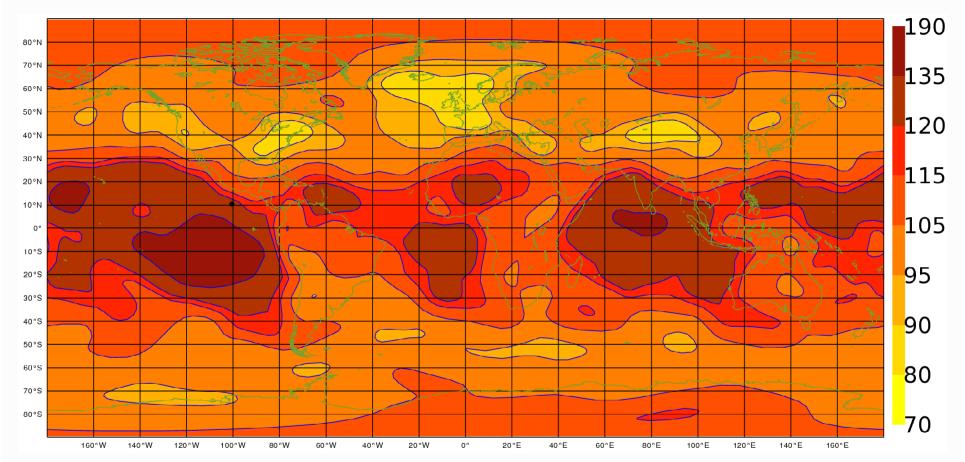
Schur filtering of long-distance correlations



from Hamill, Chapter 6 of "Predictability of Weather and Climate"



# Flow-dependent background error correlations using EnDA and wavelets



Wavelet-implied horizontal length-scales (in km), for wind near 500 hPa, averaged over a 4-day period.

(Varella et al 2013)

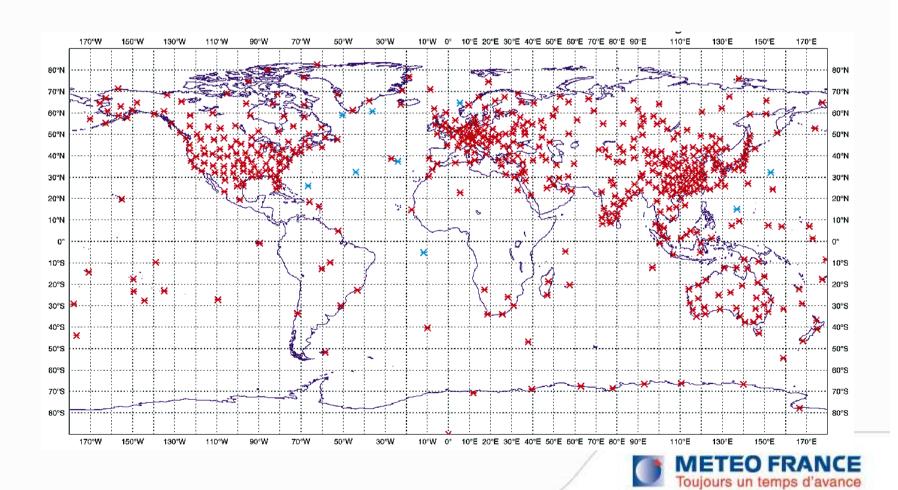


# 4. A posteriori diagnostics

(observation-minus-background departures)



## RADIOSONDE OBSERVATIONS



### Covariances of innovations

• Innovation = observation-minus-background :

$$y_o - H x_b = y_o - H x_t + H x_t - H x_b$$
  
=  $e_o - H e_b$ 

Innovation covariances:

$$E[(y_o-Hx_b)(y_o-Hx_b)^T] = R + HBH^T$$

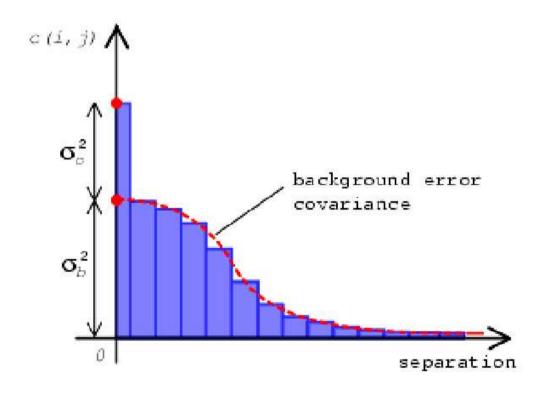
assuming that  $E[(e_o)(He_b)^T]=0$ .

(e.g. Hollingsworth and Lönnberg 1986)





## Hollingsworth and Lönnberg method



(From Bouttier and Courtier, ECMWF)

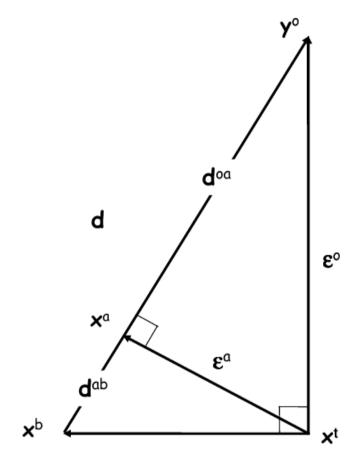


## Innovation method: properties

- Provides estimates in observation space only.
- A good quality data dense network is needed.
- Assumption that observation errors are « white ».
- An objective source of information on B and R.



### Diagnostics in observation space



(Desroziers et al, 2005)

10/36

$$d = y^{\circ} - \mathcal{H}(x^{\circ})$$

$$\mathbf{d}^{\mathrm{oa}} = \mathbf{y}^{\mathrm{o}} - \mathcal{H}(\mathbf{x}^{\mathrm{a}})$$

$$\mathbf{d}^{ab} = H(\mathbf{x}^a) - H(\mathbf{x}^b)$$

• 
$$E[d^{\circ \alpha} d^T] = R$$

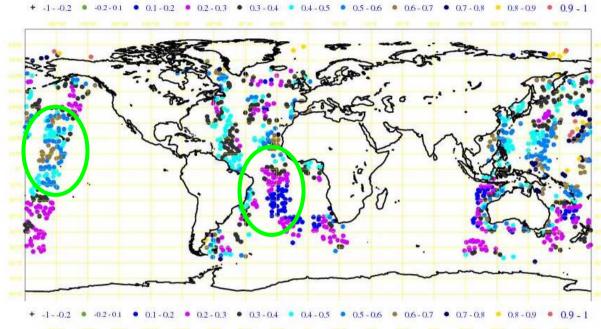
• 
$$E[d^{ab} d^T] = HBH^T$$

• 
$$\langle \varepsilon, \varepsilon' \rangle = E[\varepsilon \varepsilon'^{T}]$$



# Validation of flow-dependent estimates of errors in HIRS 7 space (28/08/2006 00h) (Berre et al 2007, 2010)

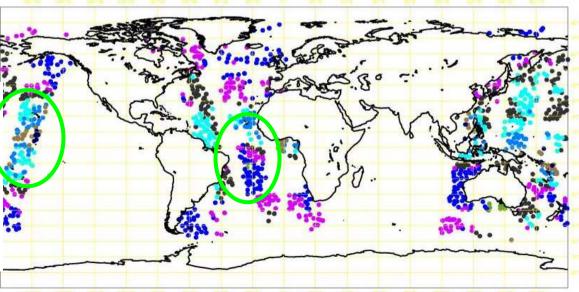
Ensemble estimate of error std-devs



« Observed » error std-devs  $cov(H dx, dy) \sim H B H^T$ 

(Desroziers et al 2005)

=> model error estimation.



# Use of innovations to estimate model error covariances Q=cov(e<sub>m</sub>)

Forecast error equation :

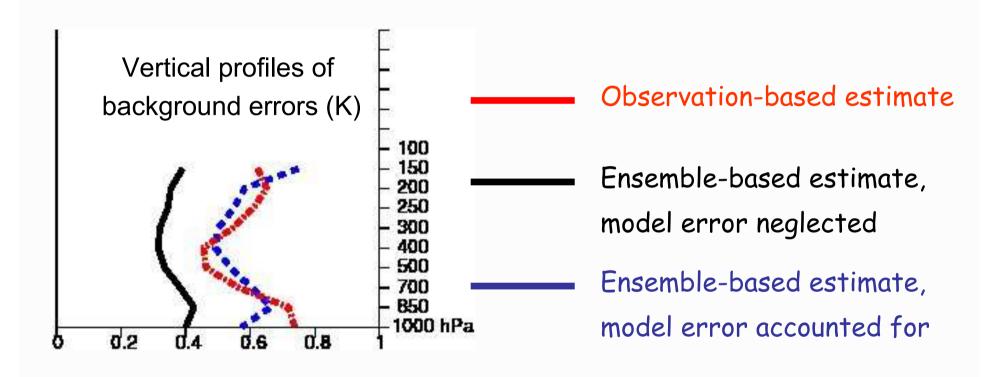
$$e_f = M e_a + e_m$$

$$B = MAM^T + Q \qquad \text{(if } e_m \text{ uncorrelated with } e_a\text{)}$$

- Use ensemble assimilation (before adding model perturbations) to estimate evolved analysis error covariances ( MAM<sup>T</sup> ).
- Use innovation diagnostics to estimate « B » (or at least HBH<sup>T</sup>)
   (forecast error covariances ).
- Estimate Q by comparing B and MAM<sup>T</sup> (e.g. Daley 1992).
- Represent model error by inflating forecast perturbations in accordance with Q estimate.



# Model error in M.F. ensemble 4D-Var (Raynaud et al 2012, QJRMS)





### Conclusions

- Data Assimilation (DA) is vital for weather forecasting (NWP).
- Observations are very diverse in type, density and quality.
- 4D-Var for temporal and non linear aspects.
- Ensemble DA methods for error simulation and covariance estimation.
- Sampling noise issues and filtering techniques.
- A posteriori diagnostics for validation of error covariances, and for estimation of model errors.



### Some references

- Desroziers, G., Berre, L., Chapnik, B. and Poli, P. (2005), Diagnosis of observation, background and analysis-error statistics in observation space. Q.J.R. Meteorol. Soc., 131: 3385-3396.
- Fisher, M., 2003: Background error covariance modeling. Proc. ECMWF Seminar on "Recent Developments in Data Assimilation for Atmosphere and Ocean", 8-12 Sept 2003, Reading, U.K., 45-63.
- Hollingsworth, A. and Lönnberg, P., 1986: The statistical structure of short-range forecast errors as determined from radiosonde data. Part I: The wind field. Tellus, 38A, 111-136
- Houtekamer, P. L., Louis Lefaivre, Jacques Derome, Harold Ritchie, Herschel L. Mitchell, 1996: A System Simulation Approach to Ensemble Prediction. Mon. Wea. Rev., 124, 1225– 1242.
- Houtekamer, P. L., Herschel L. Mitchell, Xingxiu Deng, 2009: Model Error Representation in an Operational Ensemble Kalman Filter. Mon. Wea. Rev., 137, 2126-2143.
- Rabier et al 2000: The ECMWF operational implementation of four-dimensional variational assimilation. Part I: Experimental results with simplified physics. Q. J. R. Meteorol. Soc., 126, 1143-1170.
- Talagrand, O. and P. Courtier, 1987: Variational assimilation of meteorological observations with the adjoint vorticity equation. I: Theory. Quart. J. Roy. Meteor. Soc., 113, 1311-1328.
- Berre, L., Ştefănescu, S., Belo Pereira, M.. The representation of the analysis effect in three error simulation techniques. Tellus A, 58A, pp 196-209.

# Thank you

for your attention

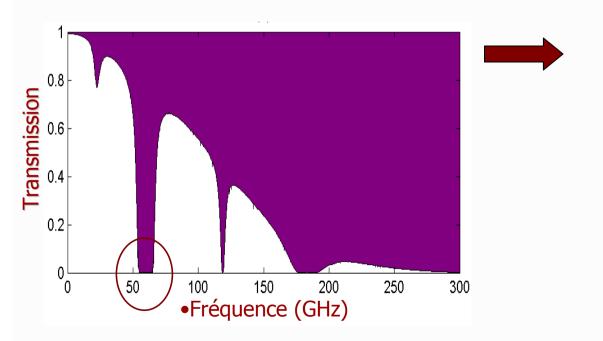


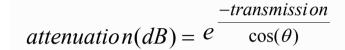
#### What is measured by satellite sensors?

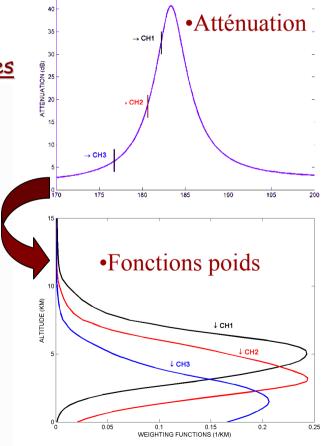
#### Soundings of atmosphere?

- In micro-waves: absorption par by water vapor, oxygen
- Largeur des bandes d'absorption: Pression (altitude) (< 60km): les bandes d'absorption plus larges quand la pression augmente

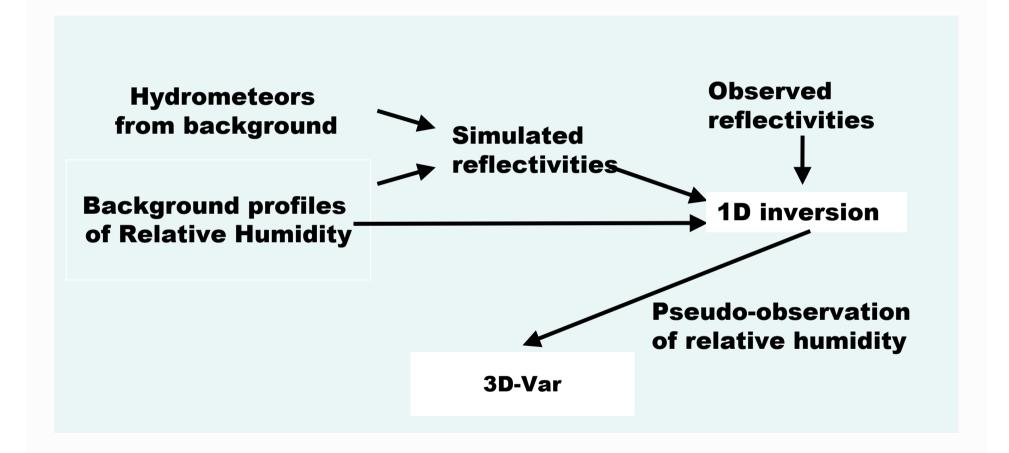
Les mesures loin (proches) d'une bande d'absorption: information sur les basses (hautes) couches atmosphériques







# Assimilation of reflectivities in AROME : Method 1D + 3D-Var : general algorithm





#### Spatial filtering of raw ensemble variances

Expansion of the raw variance field V<sub>raw</sub>:

$$V_{raw} = V_{signal} + V_{noise}$$

with  $V_{signal}$  assumed uncorrelated with true signal  $V_{signal}$ 

Filtering V<sub>raw</sub> through linear regression formalism :

$$V_{signal} \sim V_{filtered} = F V_{raw}$$

$$= cov(V_{signal}, V_{raw}) / var(V_{raw}) V_{raw}$$

$$= 1/(1 + var(V_{signal}) / var(V_{noise})) V_{raw}$$

Estimation of signal and noise variances (in spectral space):

$$var(V_{noise}) = 2/(N-1) B* ° B*$$
  
 $var(V_{signal}) = var(V_{raw}) - var(V_{noise})$ 

=> F = low-pass spectral filter, equivalent to local spatial averaging.



# Modelling of background error covariances

- Size of B is far too large.
- Can't be computed explicitly (nor stored in memory).
- $\Rightarrow$  Model B as product of sparse operators.



## B as product of sparse operators

$$B^{1/2} = L S C_{11}^{1/2}$$

L: ~ cross-covariances (~sparse regressions),

5 : diagonal matrix of standard deviations.

 $C_{\rm u}$ : sparse model of auto-correlations (e.g. diagonal matrix in spectral space).

$$B = L S C_u S L^T$$

### Covariances of residuals

• Analysis increment :  $H \delta x = HK (y_o - Hx_b)$ with  $HK = HBH^T (HBH^T + R)^{-1}$ 

• Covariances between  $H\delta x$  and omb:

$$E[(H \delta x)(y_{o}-Hx_{b})^{T}] = HK E[(y_{o}-Hx_{b})(y_{o}-Hx_{b})^{T}]$$

$$\sim HK (HB_{t}H^{T}+R_{t})$$

$$\sim HBH^{T} (HBH^{T}+R)^{-1} (HB_{t}H^{T}+R_{t})$$

$$\sim HB_{t}H^{T}$$

either assuming K ~ optimal, or, for averaged  $\sigma_b$ , assuming that structures in B,R are much different. (Desroziers et al 2005)

### Model error representations

- Additive inflation (temporally uncorrelated):
   random draws from estimated model error covariances.
- Multiplicative inflation (temporally correlated):
   mult. amplification of forecast perturbations.
- Multi-model ensembles (difficult to maintain?):
   use different models to reflect model uncertainties.
- Stochastic physics: perturbations with amplitudes proportional to physical tendencies.
- SKEB: backscattering of small scale energy dissipated by horizontal diffusion.
- ⇒ Comparison by Houtekamer et al 2009 : inflation is the most « efficient » approach.