Introduction to the principles and methods of data assimilation in the geosciences

Lecture notes
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Introduction to data assimilation
Synopsis

In the first part of the lecture notes, the statistical tools that represent the foundations of data assimilation are introduced. These tools are closely related to those of estimation theory and to those of optimal control. The basic concept of statistical interpolation will be introduced and the pursued goals will be clarified. Under certain conditions, statistical interpolation becomes the so-called optimal interpolation, that entirely relies on the method known as BLUE (that stands for Best Linear Unbiased Estimator). The BLUE analysis turns out to often be equivalent to a variational problem which consists in minimising a least squares functional.

Time is easily incorporated into the optimal interpolation approach, using cycles of optimal interpolation, which yields the method known as 3D-Var. A more sophisticated sequential extension in time is known as the Kalman filter. Under certain conditions, a variational principle equivalent to the Kalman filter can be derived. It is meant to minimise a functional defined over a four-dimensional space (3 for space and 1 for time). This method is known as 4D-Var.

In the second part of the lecture notes, key probabilistic and ensemble data assimilation techniques are presented. The hope is that the uncertainty quantification be better achieved with these methods, while, even though the adjoint of the evolution model is not used, non-linearity could be better taken into account. We will explain what the probabilistic Bayesian standpoint to data assimilation is and introduce Bayes’ rule. We will sketch the basics of the particle filter, which is essentially a Monte-Carlo implementation of Bayes’ rule. Finally we will introduce the ensemble Kalman filter (EnKF) which is the practical and efficient adaptation of the Kalman filter for high-dimension applications, such as in the geosciences. To make it viable, one needs a few clever tricks whose explanation together with simple implementations are given.
Introduction to data assimilation
Textbooks and lecture notes

One of the fathers of modern data assimilation was Roger Daley, who wrote the first textbook on the subject (Daley, 1993). Crystal-clear lecture notes and review articles that have been written by researchers with significant contributions to the subject offer a nice first reading on the topic of data assimilation (Talagrand, 1997; Bouttier, 1997). Furthermore, detailed lecture notes are available on the web (in particular Todling, 1999). A textbook on data assimilation and predictability in meteorology has been written by one of the main researchers in the field, Eugenia Kalnay (Kalnay, 2003), and is filled with nice insights, ideas and explanations. For a clarification on the concept of errors in the field of data assimilation, Cohn (1997) is the recommended review article. For an exposition of conventions and notations in the field, as well as a review of the methods, one is referred to Ide et al. (1999). Other highly recommended but more specific and more recent textbooks are Evensen (2009); Lahoz et al. (2010); Blayo et al. (2015); Fletcher (2017). Others recommended ones but a more mathematical emphasis are Reich and Cotter (2015); Law et al. (2015).

I highly recommend the recent textbook on data assimilation by Asch et al. (2016). I am co-author of the book and I am proud to mention that these notes inspired several chapters of the book. The first part of the book is for beginners (and covers the first three chapters of these notes); the second part covers more advanced topic (including chapter 5 of these notes) while the last part gives many illustrated examples of the use of data assimilation in many fields.

Finally, the review articles Carrassi et al. (2018); Janjić et al. (2018) offer a modern overview on data assimilation, while being shorter than the textbooks.
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Part I

Methods of data assimilation
Chapter 1

Statistical interpolation

1.1 Introduction

How does one forecast the state of complex physical systems such as the atmosphere?

Let us take the concrete example of the temperature in Paris. Suppose daily temperature has been measured over the city in the past few days. One wishes to forecast the daily temperature in the forthcoming days. A simple procedure consists in extrapolating past and already monitored temperatures to the coming days, assuming that the historical temperature curve is smooth enough. Or one can rely on a statistical and historical database that gives the average temperature for a given day of the year. This was the practice before the era of numerical computing because the procedure is computationally cheap.

Another approach consists in using the knowledge one has on the dynamics of the system by simulating the evolution of the temperature in the days to come. This is nowadays enabled by considerable computing power. This requires to initialise the numerical model with the latest recorded temperatures.

However, these methods appear to have serious limitations and soon enough the forecasted temperature curve diverges from the true temperature curve (Figure 1.1), because of the fundamentally poor predictability of the meteorological system.

An optimal procedure would be to combine the most exhaustive theoretical and observational knowledge. In our example, this means initialising the dynamics of the system using all the past and present observations and not only the latest.

Data assimilation is defined as the set of statistical techniques that allows to improve the knowledge of the past, present or future system states, jointly using experimental data and the theoretical (a priori) knowledge on the system.

Statistical interpolation is one of the most simple techniques offering a solution to this problem. Even though the technique is elementary (basically equivalent to a linear regression), its implementation on complex high-dimensional systems is not straightforward.

1.1.1 Representation of the physical system

The state of the system is represented by a vector $\mathbf{x}$ composed of scalar entries. For instance, each one of these entries can stand for the value of temperature of an air parcel.
at precise coordinates (latitude, longitude, altitude) at a given well defined date.

In principle, a comprehensive description of the system state supposes that \( x \) is a continuous vector field of components \( x_\alpha(\overrightarrow{OM}, \tau) \) rather than a finite-dimensional vector field. For each location and time \((\overrightarrow{OM}, \tau)\), \( x \) takes value in a vector space, the components of which are indexed by \( \alpha \) and can represent temperature, wind speed, moisture, pressure, gas concentrations, etc.

With a view to numerical computations, it is necessary to represent this continuous field in a limited-memory computer and hence to discretise it. Formally, one needs an operator \( \Pi \) that projects the infinite dimensional space of the vector field to the finite-dimensional space of its numerical representation. The image of \( x \) through \( \Pi \) is denoted \( x^t \), with \( x^t \in \mathbb{R}^{N_x} \). The \( t \) index refers to \textit{truth}, the \textit{true} state of the system. It is obviously an abuse of language since \( x^t \) is only a projection of the true continuous state. For instance, \( x^t \) could be the average of field \( x \) within a cell of a grid used for the numerical representation.

However, it is in practice impossible to know the exact true value of the entries of \( x^t \), as much as it is impossible to know those of \( x \). Therefore, our goal is to estimate this vector. The result of the estimation procedure for \( x^t \) is a vector of the same dimension \( N_x \), denoted \( x^a \), where \( a \) refers to \textit{analysis}. If the outcome of an earlier or present time partial analysis is used as the starting point of a new analysis, this vector is denoted \( x^b \in \mathbb{R}^{N_x} \) where \( b \) stands for \textit{background}. It represents some prior knowledge meant to be used in the subsequent analysis.

An estimation system needs to feed on observations of the real system in order to reduce the errors committed in the estimation of the system state. These observations can be performed in any location and time. They can even be averages of the state variables when the instrument resolution is coarse. In realistic systems, these locations are not necessarily collocation points used in the numerical discretisation of the system computer representation. By interpolating, extrapolating and filtering, one should be able to provide a data set of size \( N_y \) which is related by a map to all or part of the entries of \( x^t \); \( y \in \mathbb{R}^{N_y} \).
will denote the vector of observations.

1.1.2 The observational system

Instruments to monitor the atmosphere

An ambitious data assimilation system requires a lot of observational data. Here, we give an overview of the global observational system used by the numerical weather prediction operational centres.

Fundamentally, there are two types of observations. The first set gathers the conventional observations. Today, they are far less abundant than the remote sensing observations, especially those delivered by space-born instruments (on satellites). However, the conventional observations are less redundant and, in general, more precise. Besides, they offer a direct, often in situ, sampling of the physical system as opposed to remote sensing observations, such as radiances. In 2014, 99% of the data processed by the European Centre for Medium-Range Weather Forecast (ECMWF) originated from satellites but only 91.5% of the observations that are actually used in the analysis were space-born.

These conventional observations include:

- the measurements from synoptical stations (SYNOP network),
- measurements from commercial aircraft on regular flights,
- measurements from the ocean buoys and measurements from the commercial ships,
- measurements from the sounding balloons (radiosounding),
- LIDAR measurements (active sounding by laser ray from the ground),
- radar measurements, that enables to locate water content, precipitations.

The second set of observations correspond to the satellite measurements that emerged in the 1970's. Since the end of the 1970's until a few years from now, the total number of measurements acquired at each round of observation had increased exponentially. In 2014, the ECMWF processed $70 \times 10^6$ scalar data per day, but only $3.6 \times 10^6$ scalar data were actually assimilated.

From the data assimilation standpoint, satellite has enabled the use of a considerable amount of data. It has also enabled a much better coverage of the globe, and especially in the southern hemisphere mostly covered by oceans. Significant progress has been recorded when the monitoring of the southern hemisphere by satellite was substantially improved.

1.1.3 Error modelling

Modelling the observation process and the observational error

In practice, the knowledge of $y$ only gives partial and flawed information about the state vector $x_t$. There is a chain of transformations between $y$ and $x_t$ needed to relate them. Some of these transformations are accurate; others are approximate and erroneous. Let us have a look at this chain.

We have already mentioned the fact that only part of the state vector can be probed by observation. For instance, in oceanography, satellite sounders provide data on the sea
surface height, and hence provide information on the sea motion at the surface, but it is
much more difficult to get marine data in the depths. The lack of observations or observ-
bility is not a source of error per se; it is merely the signature of an underdetermination
of the system state (roughly $N_y \ll N_x$). For the ECMWF meteorological model (the IFS)
and for a data assimilation cycle, vector $x^t$ is of size $N_x = 2 \times 10^9$ (T1279L137 since July
2013) whereas the observation vector $y$ is of dimension $N_y = 2 \times 10^7$!

To complicate the processing of observations, there is not always an immediate relation-
ship between the state variables of $x^t$ and the observations of $y$. For instance, radiances
measured by satellite instruments depend on the total vertical atmospheric column. The
entries of $y$ are related to the system state only through linear combinations or nonlinear
functions of the variables of $x^t$. Hence, there is a quite involved map that relates $x^t$ to $y$.

The underdetermination is not the only source of mismatch between the observations
and the system state. These errors can be classified into instrumental errors and
representativeness errors.

The most obvious usually well identified and characterised, source of error is the instrumen-
tal error induced by the measurement process. It affects the value of $y$ additively or/and multiplicatively. Knowing the true and continuous state of the system $x$, one wishes to build the vector of the observations meant to represent $y$. We shall denote this vector $h[x]$, whose outcome is the construction of the map $h$. In the absence of instrumen-
tal error, one has $y = h[x]$. But in the presence of (additive) instrumental error $e^t \in \mathbb{R}^{N_y}$, we should write instead

$$y = h[x] + e^t. \quad (1.1)$$

However, in practice, one can only use $x^t$, the numerical counterpart of $x$. It is therefore useful to build the vector of observations that can be compared to $x$, using the true discrete state of the system. This somehow different observation map can be written $H[x^t]$. The map $H$ is known as the observation operator. It includes several transformations such as projections, but also interpolations made necessary by the information lost in the projection by $\Pi$. This is formally written as

$$y = h[x] + e^t + e^\mu,$$

where

$$e^r \triangleq h[x] - H[x^t] = h[x] - H[\Pi x] = (h - H \circ \Pi)[x], \quad (1.3)$$

is the representativeness, or representation error. The symbol $\triangleq$ signals a defini-
tion. This construct is summarised by the schematic in Fig. 1.2. It can be condensed into

$$y = H[x^t] + e^o,$$

with $e^o \triangleq e^r + e^\mu$. Here, we have assumed that $h$ and $H$ are well known. But an additional source of error could be an imperfect knowledge of the map $H$. In that case, the error in the modelling of $H$ is called model error.

**Error statistics**

We first assume that the observation error has no bias. This means that $\mathbb{E}[e^o] = 0$, where $\mathbb{E}$ is the expectation operator. If, on the contrary, there is a bias, *i.e.* $b \triangleq \mathbb{E}[e^o] \neq 0$, it is
often possible to diagnose it and to subtract its value from \( e^o \), so as to make the corrected error unbiased, \( i.e. \mathbb{E}[e^o - b] = 0 \). Let us introduce \( R \), the observation error covariance matrix defined by

\[
[R]_{ij} = \mathbb{E}[[e^o]^i, [e^o]^j].
\]

(1.5)

This is a symmetric matrix of dimension \( \mathbb{R}^{N_y \times N_y} \). Moreover, we assumed it to be positive definite (\( i.e. e^\top R e > 0 \) for all \( e \in \mathbb{R}^{N_y} \) different from \( 0 \)), which implies that it is invertible.

In the following, we shall assume that \( R \) is known. Practically, it is important to either know it or to have a good estimation of it; the quality of the analysis depends on it. Quite often \( R \) is diagonal or assumed so, which means that the observations are statistically independent from one another. In practice, correlations are nevertheless possible: temporal and spatial correlation for the satellite measurements, etc.

**Background and analysis errors**

The background error is defined by

\[
e^b = x^b - x^i.
\]

(1.6)

This is a measure of the discrepancy between an a priori estimate (before the analysis) and the (unknown) truth. This leads to the definition of the background error covariance matrix \( B \):

\[
[B]_{ij} = \mathbb{E}[[e^b]^i, [e^b]^j].
\]

(1.7)

This is a symmetric matrix of dimension \( \mathbb{R}^{N_x \times N_x} \). Moreover, it is assumed to be positive definite, hence invertible. It is assumed that \( e^b \) has no bias, \( i.e. \mathbb{E}[e^b] = 0 \), or that the bias has been subtracted.

The analysis error is defined by

\[
e^a = x^a - x^i.
\]

(1.8)
It defines the gap between the outcome $x^a$ of the analysis process and the (unknown) truth $x^t$. It leads to the definition of the analysis error covariance matrix $P^a$ as

$$[P^a]_{ij} = E [(e^a)_i, (e^a)_j].$$

which is a symmetric matrix of dimension $\mathbb{R}^{N_x \times N_x}$.

Furthermore, we shall assume that $e^a$ and $e^b$ are uncorrelated. There is no fundamental reason why the observation of a system, seen as a stochastic process, be correlated to a prior knowledge of the same system. This seems reasonable even in a realistic context. However, if the background has been obtained from observations of the same type, possibly with the same instruments, this assumption could be breached.

**Model error**

In the field of geophysical data assimilation, an evolution model $\Phi$ relate $x(\tau + 1)$ to $x(\tau)$, system states at times $\tau + 1$ and $\tau$:

$$x(\tau + 1) = \Phi [x(\tau)].$$

In meteorology, this model could be the integral operator, or *resolvent*, of the so-called *primitive* equations (momentum equations, air mass conservation equation, water mass conservation equation, state equation for the air (perfect gas), conservation of energy (first law of thermodynamics)). In atmospheric chemistry data assimilation, it could correspond to the resolvent of the transport and fate equations of chemical gaseous, particulate and aerosol species.

At first, we assume that this transition operator $\Phi$ is known. It is a faithful representation of the dynamics of the real system. But, again, a numerical representation implies a discretisation: $x^t = \Pi x$. The propagation Eq. (1.10) has to be projected using $\Pi$, which yields

$$x^t(\tau + 1) = \Pi x(\tau + 1) = \Pi \Phi [x(\tau)]$$

$$\triangleq M [x^t(\tau)] + e^m,$$

where the *model error* $e^m$ is given by

$$e^m \triangleq \Pi \Phi [x(\tau)] - M [x^t(\tau)] = \Pi \Phi [x(\tau)] - M [\Pi x(\tau)] = (\Pi \circ \Phi - M \circ \Pi) [x(\tau)].$$

This kind of model error belongs to the class of the *representativeness (or representation) errors*.

However, because the system that is modelled is complex and because of the necessary approximations in the derivation and implementation of the model, the numerical model $\Phi$ has to be imperfect, let alone $M$. Therefore, the model error $e^m$ is actually the sum of a representativeness error as already identified, and of an error that characterises the imperfection of the model. This latter error is distinct from the modelling error in the observation operator although they are similar in nature.

**1.1.4 The estimation problem**

The goal is to study a physical system described by the vector state $x^t$. One assumes that our best estimation of the system state is $x^b$. This background is likely to result from an
earlier data assimilation analysis or from an earlier statistical analysis. In the absence of any other source of information, it sums up to our best estimation of the system state. Observations $y$ that are performed on the system brings in new information through the $H$ operator. Moreover, ideally, one assumes to know the statistics of the observation error $e^o$ up to second-order moments ($\mathbb{E}[e^o], R$). We also assume that we know the statistics for the background error $e^b$ up to second-order moments ($\mathbb{E}[e^b], B$).

When new observations are available, we are in a position to improve our estimation and yield $x^a$, exploiting the information contained in these new observations as well as the background $x^b$. We are also interested in evaluating the error $e^a$ committed in the analysis, or its statistics.

There are of course many ways to define an analysis. Yet, we are interested in the best possible analysis, i.e. an analysis that would minimise the magnitude of the analysis error, for instance by minimising the trace of the analysis error covariance matrix $\text{Tr}(P^a)$.

### 1.2 Statistical interpolation

#### 1.2.1 An Ansatz for the estimator

**The linear hypothesis**

Here, we shall assume that the observation operator $H$ is linear. It can be seen as a matrix in $\mathbb{R}^{N_y \times N_x}$. It will be denoted $H$, which indicates that the operator is linear, or that it is the tangent linear operator\(^1\) of the nonlinear operator $H : \mathbb{H} = H'$.

Let us define what a tangent linear operator is as it plays an important role. If the observation operator is the map: $x \mapsto y = H[x]$, the related tangent linear operator is defined by the expansion of the operator around $x$; for $i = 1, \ldots, N_y$

$$\delta y_i = \sum_{j=1}^{N_x} \left. \frac{\partial H_i}{\partial x_j} \right|_x \delta x_j. \quad (1.13)$$

The tangent linear operator is the linear operator whose action is defined by the matrix $H$, of entries

$$[H]_{ij} = \left. \frac{\partial H_i}{\partial x_j} \right|_x. \quad (1.14)$$

If $H$ is linear, i.e. $H[x] = Hx$, then $H' = H$; the tangent linear identifies with the original observation operator. Consequently, the tangent linear operator $H' = H$ depends on $x$ if and only if $H$ is nonlinear.

A simple but nevertheless non-trivial Ansatz for the estimate $x^a$ consists in choosing for the analysis a vector of the form

$$x^a = Lx^b + Ky, \quad (1.15)$$

where $L$ is a matrix of dimension $N_x \times N_x$ and $K$ is a matrix of dimension $N_x \times N_y$. Hence $x^a$ is a linear combination of the available information ($L$ and $K$ are linear operators).

---

\(^1\)also called the *Jacobian* matrix.
Given the observation equation \( y = Hx^t + e^o \), the error attached to this combination can be estimated as follows.

\[
x^a - x^t = L \left( x^b - x^t + x^t \right) + K \left( Hx^t + e^o \right) - x^t
\]

\[
e^a = Le^b + Ke^o + \left( L + KH - I \right)x^t.
\] (1.16)

First, one wishes that the errors be unbiased. Recalling that we assumed that the observation and background errors are unbiased \( \mathbb{E}[e^o] = 0 \) and \( \mathbb{E}[e^b] = 0 \), one can infer from the previous calculation that \( \mathbb{E}[e^a] = \left( L + KH - I \right) \mathbb{E}[x^t] \). Therefore, we require that

\[
L = I - KH,
\] (1.17)

which is a sufficient condition (though not necessary).

As a result, we obtain a more constrained Ansatz which is now linear and unbiased:

\[
x^a = (I - KH)x^b + Ky,
\]

\[
x^a = x^b + K \left( y - Hx^b \right).
\] (1.18)

\( K \) is a linear map, i.e. a matrix, from \( \mathbb{R}^{Ny} \) to \( \mathbb{R}^{Nx} \). It is called the gain. The vector \( y - Hx^b \) in \( \mathbb{R}^{Ny} \) is called the innovation vector. This innovation is an expression formula for the additional information brought in by the observations compared to the background. The error covariance matrix that can be obtained from the innovation vectors is what is usually called the information matrix in information theory.

Since \( K \) is linear, the analysis is merely a linear interpolation. Actually, it should more adequately be called a linear regression. This interpolation has an historical naming since the first analyses used in meteorology (Cressman) were genuine mathematical interpolations (i.e. \( y = H[x^a] \)).

With this linear estimator, the estimation problem now amounts to finding a “satisfying gain” \( K \).

The posterior error

Assume for a moment that we have determined an optimal gain matrix \( K \). Then, what would be the error covariance matrix \( P^a \)? In order to compute it, Eq. (1.18) is written using the error vectors that have already been introduced

\[
e^a = e^b + K(e^o - He^b)
\] (1.19)

so that

\[
P^a = \mathbb{E}\left[ (e^a)(e^a)^\top \right] = \mathbb{E}\left[ (e^b + K(e^o - He^b))(e^b + K(e^o - He^b))^\top \right]
\]

\[
= \mathbb{E}\left[ (Le^b + Ke^o)(Le^b + Ke^o)^\top \right] = \mathbb{E}\left[ Le^b(e^b)^\top L^\top \right] + \mathbb{E}\left[ Ke^o(e^o)^\top K^\top \right]
\]

\[
= LBL^\top + KRK^\top,
\] (1.20)

where we used the decorrelation of \( e^o \) with \( e^b \) and the linearity of \( K \). To summarise, we have

\[
P^a = (I - KH)B(I - KH)^\top + KRK^\top.
\] (1.21)
1.2.2 Optimal estimation: the BLUE analysis

Continuing on, we seek to minimise the error committed in the analysis, as measured by the scalar quantity \( \text{Tr}(P^a) \). Since we look for an optimal gain \( K^* \), we study the variation of \( \text{Tr}(P^a) \) with respect to a variation \( \delta K \) of \( K \) (i.e. a generic infinitesimal variation in the entries of \( K \)):

\[
\delta(\text{Tr}(P^a)) = \text{Tr}\left((-\delta KH)BL^T + LB(-\delta KH)^T + \delta KRK^T + KR\delta K^T\right)
\]

\[
= \text{Tr}\left((-LB^TH^T - LBH^T + KR^T + KR)(\delta K)^T\right)
\]

\[
= 2\text{Tr}\left((-LBH^T + KR)(\delta K)^T\right). \quad (1.22)
\]

We have used \( \text{Tr}(A) = \text{Tr}(A^T) \) and the fact that \( B \) and \( R \) are both symmetric. At optimality, one infers that

\[
-(I - K^*H)BH^T + K^*R = 0
\]

from which we obtain

\[
K^* = BH^T(R + HBH^T)^{-1}. \quad (1.23)
\]

The estimates of \( x^a \) and \( P^a \), which are the outcomes of this analysis, are called a BLUE analysis for Best Linear Unbiased Estimator, since it is: linear (primary hypothesis through \( L \) and \( K \)), without bias (first step of the derivation: \( L = I - KH \)) and optimal (second step of the derivation).

1.2.3 Properties

Sherman-Morrison-Woodbury formula

Formula Eq. (1.23) is the most frequently used form of the optimal gain. However, the same optimal gain can be written differently

\[
K^* = (B^{-1} + H^TR^{-1}H)^{-1}H^TR^{-1}. \quad (1.24)
\]

The equivalence between Eq. (1.23) and Eq. (1.24) is an immediate consequence of the Sherman-Morrison-Woodbury identity. It is proven through

\[
BH^T(HBH^T + R)^{-1}
\]

\[
= (B^{-1} + H^TR^{-1}H)^{-1}(B^{-1} + H^TR^{-1}H)BH^T(HBH^T + R)^{-1}
\]

\[
= (B^{-1} + H^TR^{-1}H)^{-1}(H^T + H^TR^{-1}HBH^T)(HBH^T + R)^{-1}
\]

\[
= (B^{-1} + H^TR^{-1}H)^{-1}H^TR^{-1}(R + HBH^T)(HBH^T + R)^{-1}
\]

\[
= (B^{-1} + H^TR^{-1}H)^{-1}H^TR^{-1}. \quad (1.25)
\]

This equivalence turns out to be useful, both from a theoretical but also a practical standpoint. For instance, the observation space is quite often much smaller than the dimension of the state space, so that the inversion of the matrix \( B^{-1} + H^TR^{-1}H \) is much more costly than the inversion of the matrix \( R + HBH^T \). Therefore, it is often useful to resort to (1.23).
Optimal analysis error

Choosing the optimal gain Eq. (1.23) for the error estimation, the posterior error covariance matrix Eq. (1.21) simplifies to

\[ P^a = (I - K^* H)B. \]  

(1.26)

Indeed, we have

\[
  P^a = (I - KH)B(I - KH)^\top + KRK^\top \\
  = (I - KH)B + \left[ KR - (I - KH)BH^\top \right] K^\top. \tag{1.27}
\]

The expression in brackets in the right-hand side is zero when the gain is optimal \((K = K^*)\) so that we obtain the desired result. This justifies the name \textit{gain matrix}; the operator \(I - K^* H\) measures the shrinkage of the innovation vector into

\[
y - Hx^a = (I - HK^*)(y - Hx^b), \tag{1.28}
\]

which is called the \textit{analysis residue}.

Two classical expressions for \(P^a\) can be obtained using Eq. (1.26), together with the two formula for the gain, Eq. (1.23) and Eq. (1.24). Indeed one has in the one hand

\[
P^a = (I - K^* H)B \\
= (I - BH^\top (R + HBH^\top)^{-1} H)B \\
= B - BH^\top (R + HBH^\top)^{-1} HB, \tag{1.29}
\]

and

\[
P^a = (I - (B^{-1} + H^\top R^{-1} H)^{-1} H^\top R^{-1} H)B \\
= (B^{-1} + H^\top R^{-1} H)^{-1} \left( B^{-1} + H^\top R^{-1} H - H^\top R^{-1} H \right) B \\
= (B^{-1} + H^\top R^{-1} H)^{-1} \tag{1.30}
\]

in the other hand. This last formula establishes that \(P^a\) is invertible and hence that it is a symmetric positive definite matrix. It also shows that the inverses of the error covariance matrices (\textit{confidence}, or \textit{precision} matrices) are additive because

\[
(P^a)^{-1} = B^{-1} + H^\top R^{-1} H. \tag{1.31}
\]

The optimal gain \(K^*\) can be related to \(P^a\) through the following useful formula

\[
K^* = P^a H^\top R^{-1}. \tag{1.32}
\]

Indeed, one has

\[
P^a H^\top R^{-1} = (B^{-1} + H^\top R^{-1} H)^{-1} H^\top R^{-1} = K^*. \tag{1.33}
\]
The innovation and the analysis residue are unbiased

A consequence of the observation operator linearity and of the absence of bias of the analysis error $e^a$ is that the analysis residue $y - Hx^a$ is also unbiased. Indeed

$$y - Hx^a = Hx + e^o - Hx^a = e^o - He^a,$$

from which it is immediate to conclude $E[y - Hx^a] = 0$. Because we have assumed that the background error is unbiased, we have the same for the innovation vector:

$$y - Hx^b = Hx + e^o - Hx^b = e^o - He^b,$$

from which we obtain $E[y - Hx^b] = 0$.

Orthogonality of the analysis with the analysis error

Let us calculate the following covariance matrix even if the gain is not optimal

$$C = E[x^a (e^a)^\top].$$

We assume that the background satisfies:

$$E[x^b (e^b)^\top] = 0.$$

This means that the background and its related error are uncorrelated.

We have seen that $x^a = x^b + K(y - Hx^b) = x^b + K(-He^b + e^o)$. As a consequence

$$C = E[(x^b + K(-He^b + e^o))(I - KH)^\top] = -KHE[e^b(e^b)^\top](I - KH)^\top + KE[e^o(e^o)^\top]K^\top$$

Then, if the analysis is optimal, we obtain $KR - (I - KH)BH^\top = 0$, so that $C = 0$. Hence, the estimate $x^a$ and the analysis error $e^a$ are orthogonal. Actually orthogonality and optimality are equivalent here! That is why the analysis error is uncorrelated from the estimate resulting from the same analysis.

1.3 Variational equivalence

Let us make the same assumptions as for the BLUE derivation, i.e. that $H$ is a linear operator and denoted $H$. All the statistical hypotheses remain the same. We wish to show that the BLUE result can be be obtained using variation calculus, i.e. through the minimisation of a (multivariate) function. This is a variational approach to the problem. A sufficiently regular cost function $F$, i.e. a function from $\mathbb{R}^{N_x}$ to $\mathbb{R}$, have a generic quadratic expansion around $x_0$ of the form

$$F(x) = F(x_0) + (x - x_0)^\top \nabla F|_{x_0} + \frac{1}{2}(x - x_0)^\top \text{Hess}|_{x_0}(x - x_0) + o(||x - x_0||^2),$$

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where $\nabla F|_{x_0}$, a vector that generalises the first derivative, is called the gradient, and Hess, a matrix that generalises the second derivative, is called the Hessian. They are defined by

$$\nabla F|_x = \frac{\partial F}{\partial x_i|_x}, \quad [\text{Hess}|_x]_{ij} = \frac{\partial^2 F}{\partial x_i \partial x_j|_x}. \quad (1.40)$$

### 1.3.1 Equivalence with BLUE

We define the following function, from $\mathbb{R}^{N_x}$ into $\mathbb{R}$

$$J(x) = \frac{1}{2} (x - x^b)^\top B^{-1} (x - x^b) + \frac{1}{2} (y - Hx)^\top R^{-1} (y - Hx), \quad (1.41)$$

which is called a cost function or objective function. Since $H$ is linear, $J$ is a quadratic functional of $x$. Since $B$ is positive definite, this functional is strictly convex and as a consequence has a unique minimum. Where is it?

$$\delta J(x) = \frac{1}{2} (\delta x)^\top B^{-1} (x - x^b) + \frac{1}{2} (x - x^b)^\top B^{-1} \delta x$$

$$+ \frac{1}{2} (-H\delta x)^\top R^{-1} (y - Hx) + \frac{1}{2} (y - Hx)^\top R^{-1} (-H\delta x)$$

$$= (\delta x)^\top B^{-1} (x - x^b) - \frac{1}{2} (x - x^b)^\top H^\top R^{-1} (y - Hx)$$

$$= (\delta x)^\top \nabla J. \quad (1.42)$$

The extremum condition is $\nabla J = B^{-1}(x^* - x^b) - H^\top R^{-1}(y - Hx^*) = 0$, which can be written

$$x^* = x^b + \frac{(B^{-1} + H^\top R^{-1}H)^{-1}H^\top R^{-1}(y - Hx^b)}{\kappa^*}. \quad (1.43)$$

Therefore, $x^*$ identifies with the BLUE optimal analysis $x^a$. Let us remark that the Sherman-Morrison-Woodbury would be needed to prove the equivalence with the initial BLUE result, Eq. (1.23).

### 1.3.2 Properties of the variational approach

#### Analysis precision and the Hessian

The functional gradient is obtained from Eq. (1.42),

$$\nabla J(x) = B^{-1}(x - x^b) - H^\top R^{-1}(y - Hx). \quad (1.44)$$

In this particular case, since the cost function is quadratic, it is simple to compute the Hessian

$$\text{Hess}|_x = B^{-1} + H^\top R^{-1}H. \quad (1.45)$$

As a matter of fact, this coincides with the expression of the inverse of the analysis error covariance matrix that we have already encountered! Therefore, we have

$$P^a = \text{Hess}|_x^{-1}. \quad (1.46)$$

This is more than just another expression for $P^a$. This formula shows that the precision of the analysis is proportional to the curvature of the cost function $J$. Hence, the narrower the minimum, the better the analysis.
Nonlinear extension

Besides a new light shed on the problem, the variational formalism has two important traits. Firstly, it offers an elegant and straightforward extension of the linear problem to the nonlinear problem when $H$ is not linear anymore. By definition, this extension is impossible within the BLUE formalism without linearising the observation operator, which would be an approximation.

A more general algorithm

Secondly, the variational approach has an algorithmic advantage. Within the BLUE approach, one needs to compute the inverse of matrix $\mathbf{R} + \mathbf{HBH}^\top$ that appears in the gain $\mathbf{K}^\star$ (or at least solve a linear system thereof). Within the variational approach, the cost function $J$ is minimised, which requires to compute the product of a vector by the inverses of $\mathbf{B}$ and $\mathbf{R}$ several times. This may show a lower computational burden than the inversion of the full matrix when the number of iterations is limited.

The variational formulation of the statistical interpolation problem is the analysis and the main step of what is usually coined 3D-Var. This type of analysis has been used operationally in meteorological weather services in the 1990’s replacing the BLUE-like optimal interpolation. In the 2000’s it has been replaced in many centres by the 4D-Var for the synoptic scale, a generalisation of the 3D-Var that we shall discuss in chapter 3.

1.3.3 When the observation operator $H$ is non-linear

As we just explained, a significant advantage of the variational method is the possibility to rigorously handle the case where $H$ is a nonlinear operator. Let us look at the modification that this induces in the analysis. In this case, $J$ is defined by

$$ J(x) = \frac{1}{2} (x - x^b)^\top \mathbf{B}^{-1} (x - x^b) + \frac{1}{2} (y - H[x])^\top \mathbf{R}^{-1} (y - H[x]). \quad (1.47) $$

We introduce the tangent linear of $H$ at $x$, denoted $\mathbf{H}$ (see Eq. (1.14)). As opposed to the linear case, $\mathbf{H}$ now depends on where $H$ has been linearised, i.e. $x$. The gradient of $J$ is similar and reads

$$ \nabla J(x) = \mathbf{B}^{-1} (x - x^b) - \mathbf{H}_x \mathbf{R}^{-1} (y - H[x]). \quad (1.48) $$

Again, the precision of the analysis is nicely obtained from $\mathbf{P}^x = \text{Hess}^{-1}_{\mathbf{x}}$. This time, however, the Hessian really depends on $x$.

1.3.4 Dual formalism

Here, we assume that the observation operator is linear. The stating point is the following cost function:

$$ J(x) = \frac{1}{2} (x - x^b)^\top \mathbf{B}^{-1} (x - x^b) + \frac{1}{2} (y - \mathbf{Hx})^\top \mathbf{R}^{-1} (y - \mathbf{Hx}). \quad (1.49) $$

A vector of Lagrange multipliers $\mathbf{w}$ is introduced, of dimension that of the observation space, i.e. $\mathbf{w} \in \mathbb{R}^N$. It is used to enforce the observation equation, through the Lagrangian

$$ L(x, \epsilon, \mathbf{w}) = \frac{1}{2} (x - x^b)^\top \mathbf{B}^{-1} (x - x^b) + \frac{1}{2} \epsilon^\top \mathbf{R}^{-1} \epsilon + \mathbf{w}^\top (y - \mathbf{Hx} - \epsilon). \quad (1.50) $$

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The optimum with respect to $w$ of this functional is equivalent to the previous cost function $J(x)$. According to the minmax theorem, the minimum of the cost function $J(x)$ coincides with the maximum of the functional depending on $w$ generated by the minimum of $L$ with respect to $x$ and $\epsilon$, denoted $G(w)$. What is it? The null gradient condition at the minimum implies that

$$x^* = x^b + BH^Tw, \quad \epsilon^* = Rw. \quad (1.51)$$

This leads to the following dual cost function

$$G(w) = -L(x^*, \epsilon^*, w)$$

$$= \frac{1}{2}w^T\left(R + HBH^T\right)w - w^T\left(y - Hx^b\right). \quad (1.52)$$

The main advantage of this approach is that the optimisation of this cost function takes place in the observation space $\mathbb{R}^{N_y}$ rather than in state space $\mathbb{R}^{N_x}$. The observation space is usually of dimension much smaller than that of the state space. This formalism is known as PSAS standing for Physical Statistical space Assimilation System.

### 1.4 A simple example

Imagine you are shipwrecked at sea, a few kilometres away from the shores. Before boarding a small lifeboat, you just had time to measure your coordinates $(u, v) = (0, v^b)$ with high accuracy. The Ox axis is parallel to the shore, while the Oy axis is perpendicular. One hour later, you want to estimate your new coordinates because the lifeboat has drifted in the meantime. To this goal, you roughly guess the distance to the shore, denoted $v_0$ with a variance $\sigma_0^2$. But we want to use our knowledge of the precise coordinate of the wreckage location, one hour ago. In the meantime, and in the absence of major sea stream, the lifeboat has drifted. The probability for it to be at position $(u^b, v^b)$ follows a normal distribution of variance $\sigma_b^2$ which is a linear function of the elapsed time since the wreckage. It is assumed that there is no correlation between the observation process and the drifting process.

#### 1.4.1 Observation equation and error covariance matrix

The state vector is that of the lifeboat coordinates $x = (u, v)^T$. The observation equation is $y = Hx + \epsilon$, where the observation operator is $H = (0, 1)$, the observation vector is $(v_0)$, and the error $\epsilon$ follows a normal distribution of variance $\sigma_0^2$. The observation error covariance matrix is $R = (\sigma_0^2)$, while the background error covariance matrix is

$$B = \begin{pmatrix} \sigma_b^2 & 0 \\ 0 & \sigma_b^2 \end{pmatrix}. \quad (1.53)$$

#### 1.4.2 Optimal analysis

The linear analysis is of the form $x^a = x^b + K(y - Hx^b)$ which translates into

$$\begin{pmatrix} u^a \\ v^a \end{pmatrix} = \begin{pmatrix} 0 \\ v_b \end{pmatrix} + K \left(v_0 - (0, 1) \begin{pmatrix} 0 \\ v_b \end{pmatrix}\right). \quad (1.54)$$
1.4 A simple example

Let us compute the optimal gain $K^\star$. To that end, we use the most efficient expression of $K^\star$ when the number of observations is smaller than the number of state variables, i.e.

$$K^\star = \sigma_b^2 \left( \begin{array}{c} 0 \\ 1 \end{array} \right) \left( \sigma_o^2 + (0, 1)\sigma_b^2 \left( \begin{array}{c} 0 \\ 1 \end{array} \right) \right)^{-1}.$$  

$$= \frac{\sigma_b^2}{\sigma_o^2 + \sigma_b^2} \left( \begin{array}{c} 0 \\ 1 \end{array} \right).$$  

(1.55)

One concludes that

$$\left( \begin{array}{c} u_a \\ v_a \end{array} \right) = \left( \begin{array}{c} 0 \\ v_b + \frac{\sigma_b^2}{\sigma_o^2 + \sigma_b^2} (v_o - v_b) \end{array} \right).$$  

(1.56)

Since the observation did not bring any new information on coordinate Ox, we stick to the wreckage coordinate $u_a = 0$. However, as time goes by, $\sigma_b$ increases. Hence, $v_a$ converges to $v_o$, which remains the most trustworthy information in this limit.

1.4.3 Posterior error

From $P^a = (I - K^\star H)B$, we obtain that

$$P^a = \begin{pmatrix} \sigma_b^2 & 0 \\ 0 & \sigma_o^2 \sigma_b^2 \sigma_o^2 + \sigma_b^2 \end{pmatrix}.$$  

(1.57)

We see that, as expected, the analysis does not improve our knowledge of the coordinate parallel to the shore $u$. Its uncertainty stems from the drift, hence it is proportional to time and increases with it. Differently, the uncertainty in coordinate $v$ is shrunk by a factor $\sqrt{\frac{\sigma_o^2}{\sigma_o^2 + \sigma_b^2}}$. The finer the observation, the smaller the error. Denoting $\sigma_a^2$ the error on coordinate $v$ after the analysis, we find

$$\frac{1}{\sigma_a^2} = \frac{1}{\sigma_o^2} + \frac{1}{\sigma_b^2}.$$  

(1.58)
The interpretation of Eq. (1.30) is: \textit{the confidence (or precision) of the analysis is the sum of the confidence of the observation and of the background.}

### 1.4.4 3D-Var and PSAS

The related cost function of the equivalent variational problem is

\[
J(u, v) = \frac{1}{2\sigma^2_b} (u^2 + (v - v_b)^2) + \frac{1}{2\sigma^2_o} (v_o - v)^2. \tag{1.59}
\]

The cost function in the PSAS formalism is

\[
G(w) = \frac{1}{2} (\sigma^2_o + \sigma^2_b) w^2 - w(v_0 - v_b), \tag{1.60}
\]

from which it is easy to derive the solution

\[
w^* = \frac{v_o - v_b}{\sigma^2_o + \sigma^2_b}. \tag{1.61}
\]
Chapter 2

Sequential interpolation: The Kalman filter

In the previous chapter, we focused on the analysis part of a data assimilation scheme. We have found an optimal estimator, as well as its uncertainty, given some prior information (the background) and an observation set. A possible temporal dimension of the problem was mentioned but mostly left aside.

In meteorology, we are interested in real time data assimilation. The collection of the observations at the NWP centres, as well as the resulting analyses meant for real time forecasting, need to be cycled sequentially in time, since the objective is to perpetually track the system state. That is why a series of times $t_0, t_1, \ldots, t_k, \ldots, t_K$ is defined. They mark the analyses of the data assimilation scheme. For synoptical scale NWP systems, one typically has $t_{k+1} - t_k = 6$ hours.

Moreover, sequential data assimilation introduces a new ingredient in the problem compared to statistical interpolation: the evolution model for the system state typically defined between times $t_k$ and $t_{k+1}$. In meteorology and oceanography, it is the numerical model that simulates the real system dynamics (numerical implementation of the primitive equations). In atmospheric chemistry, this could be a chemical transport model, which transports the species and have them react.

The typical data assimilation scheme is as follows: at time $t_k$, we have at our disposal the outcome of a previous forecast, denoted $x_f^k$, the index $f$ standing for “forecast”. Therefore, $x_f^k$ is the analogue of the background $x_b$ of statistical interpolation of chapter 1. At time $t_k$, we collect a set of observations stacked into the vector $y_k$. Given $x_f^k$ and the observations $y_k$, an analysis is performed yielding the state estimate $x_a^k$. Then, we make use of the forward model to forecast the system state from time $t_k$ to time $t_{k+1}$, from $x_k$ to $x_{k+1}$. Note that we have given up on the index $t$ that referred to the truth in chapter 1. The outcome of the forecast is denoted $x_f^{k+1}$. It will serve as the background in the next cycle (see Fig. 2.1). And so on . . .
2.1 Stochastic modelling of the system

The following stochastic equations describe the evolution and observation of the physical system

\[
\begin{align*}
\mathbf{x}_k &= M_k(\mathbf{x}_{k-1}) + \mathbf{w}_k, \\
\mathbf{y}_k &= H_k(\mathbf{x}_k) + \mathbf{v}_k,
\end{align*}
\]

where \( \mathbf{x}_k \) is the true state of the system at time \( t_k \). Let us describe this stochastic system:

- the first equation is the forecast step. \( M_{k+1} \) is called the **resolvent.** It simulates the evolution of the system from \( t_k \) to \( t_{k+1} \). It may depend on time (non-autonomous process), hence the \( k \) index. \( \mathbf{w}_k \) is the model error within \( M_k \) compared to the true physical processes. It is assumed that this noise is unbiased, uncorrelated in time (white noise) and of error covariance matrix \( Q_k \) for each time \( t_k \), i.e.

\[
E[\mathbf{w}_k] = 0 \quad \text{and} \quad E[\mathbf{w}_k \mathbf{w}_l^\top] = Q_k \delta_{kl}.
\]

- the second equation corresponds to the observation equation. \( H_k \) is the observation operator at time \( t_k \). \( \mathbf{v}_k \) is the observation error, which is assumed to be unbiased, uncorrelated in time (white noise) and of error covariance matrix \( R_k \) for each time \( t_k \), i.e.

\[
E[\mathbf{v}_k] = 0 \quad \text{and} \quad E[\mathbf{v}_k \mathbf{v}_l^\top] = R_k \delta_{kl}.
\]

As an additional assumption, though quite a realistic one, there is no correlation between model error and observation error, which translates into

\[
E[\mathbf{v}_k \mathbf{w}_l^\top] = 0.
\]

From now on, it is assumed that the operators are linear. Hence \( M_k \) and \( H_k \) are noted \( \mathbf{M}_k \) and \( \mathbf{H}_k \). This hypothesis will be discussed, possibly lifted, later on.
2.1 Stochastic modelling of the system

2.1.1 Analysis step

At time $t_k$, we have a forecast $x^f_k$ which is assumed unbiased. It will serve as our background. The error covariance matrix related to $x^f_k$ is $P^f_k$. The statistical analysis that stems from these elements, is similar to statistical interpolation. Moreover, if the estimator of this analysis has the minimal possible variance, this analysis is the same as statistical interpolation. Because of the linearity and of the null bias, it must be a BLUE analysis. That is why we now use the results of chapter 1.

The $x^a_k$ analysis is of the form

$$x^a_k = x^f_k + K_k \left( y_k - H_k x^f_k \right), \quad (2.5)$$

which guarantees that it is unbiased, since the related errors, $y_k$ and $x^f_k$, are unbiased. As a consequence, one has $e^a_k = (I - K_k H_k) e^f_k + K_k v_k$, so that the analysis error covariance matrix is

$$P^a_k = (I - K_k H_k) P^f_k (I - K_k H_k) + K_k R_k K_k^T. \quad (2.6)$$

If the analysis is optimal, the optimal gain matrix $K^*_k$ at time $t_k$ satisfies:

$$- (I - K^*_k H_k) P^f_k H_k^\top + K^*_k R_k = 0. \quad (2.7)$$

Hence, it can be written

$$K^*_k = P^f_k H_k^\top \left( H_k P^f_k H_k^\top + R_k \right)^{-1}. \quad (2.8)$$

Finally, we obtain

$$P^a_k = (I - K^*_k H_k) P^f_k. \quad (2.9)$$

Within this sequential context, the gain is often called the Kalman gain, in honour of Rudolph Kalman, electrical engineer and mathematician, who invented this filter in the 1960s.

2.1.2 Forecast step

Up to this point, we have used the observations to update our prior knowledge of the system state (the background). Now, we need to forecast the system state from time $t_k$ to time $t_{k+1}$, using our (often imperfect) knowledge of the dynamics of the system.

We can build the following estimator

$$x^f_{k+1} = M_{k+1} x^a_k. \quad (2.10)$$

The linearity of $M_{k+1}$ ensures that the estimator is unbiased. The associated error is

$$e^f_{k+1} = x^f_{k+1} - x_{k+1}$$

$$= M_{k+1} (x^a_k - x_k) - (x_{k+1} - M_{k+1} x_k)$$

$$= M_{k+1} e^a_k - w_{k+1}. \quad (2.11)$$
from which is calculated the forecast error covariance matrix:

\[
P_{k+1}^f = \mathbb{E} \left[ e_{k+1}^f (e_{k+1}^f)^\top \right]
\]

\[
= \mathbb{E} \left[ (M_{k+1} e_k^w - w_{k+1}) (M_{k+1} e_k^w - w_{k+1})^\top \right]
\]

\[
= M_{k+1} \mathbb{E} \left[ (e_k^w) (e_k^w)^\top \right] M_{k+1}^\top + \mathbb{E} \left[ w_{k+1} w_{k+1}^\top \right]
\]

\[
= M_{k+1} P_k^a M_{k+1}^\top + Q_{k+1}.
\]  (2.12)

Contrary to the analysis step where the precision matrices were additive, the error covariance matrices are additive in the forecast step.

### 2.2 Summary, limiting cases and example

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<tr>
<td>• Computation of the analysis</td>
</tr>
<tr>
<td>( x_k^a = x_k^f + K_k^* (y_k - H_k x_k^f) )</td>
</tr>
<tr>
<td>• Computation of the error covariance matrix</td>
</tr>
<tr>
<td>( P_k^a = (I - K_k^* H_k) P_k^f )</td>
</tr>
<tr>
<td>(b) Forecast</td>
</tr>
<tr>
<td>• Computation of the forecast ( x_{k+1}^f = M_{k+1} x_k^a )</td>
</tr>
<tr>
<td>• Computation of the error covariance matrix</td>
</tr>
<tr>
<td>( P_{k+1}^f = M_{k+1} P_k^a M_{k+1}^\top + Q_{k+1} )</td>
</tr>
</tbody>
</table>

#### 2.2.1 No observation

When there is not any observation to assimilate, the Kalman filter boils down to the forecast step, which reads

\[
x_{k+1}^f = M_{k+1} x_k^f,
\]  (2.13)

\[
P_{k+1}^f = M_{k+1} P_k^a M_{k+1}^\top + Q_{k+1}.
\]  (2.14)
If the dynamics is unstable (with positive Lyapunov exponents for instance), the error will grow uncontrolled. Only the assimilation of observations can help reduce the global error.

### 2.2.2 Perfect observations

Imagine we have a great confidence in the observations, so that we set $R_k = 0$. Also assume that there are as many observations as state variables and that they are independent. That is to say $H_k$ est invertible. Then:

$$
K^*_k = P^f_k H_k^\top \left( H_k P^f_k H_k^\top + R_k \right)^{-1}
= P^f_k H_k^\top \left( H_k^\top \right)^{-1} \left( P^f_k \right)^{-1} H_k^{-1}
= H_k^{-1}. \tag{2.15}
$$

As a result, we have $P^a_k = (I - K^*_k H_k) P^f_k = 0$, so that $x^a_{k+1} = M_{k+1} x^a_k = M_{k+1} H_k^{-1} y_k$. Moreover the forecast step reads $P^f_{k+1} = M_{k+1} P^a_k M_{k+1}^\top + Q_{k+1} = Q_{k+1}$. The forecast errors only depend on model error, and the system can be perfectly known through observation.

### 2.2.3 A simple example

We are back to the lifeboat problem from chapter 1. On an hourly basis, the castaway assesses the distance between the lifeboat and the shore and proceeds to an analysis. The distance to the shore, assessed at time $t_k$, $k$ hours after the shipwreck at time $t_0$, is denoted $y_k$. This estimator is assumed to be unbiased. As before, the variance is denoted $\sigma_o^2$ and is supposed to be stationary. The true coordinates of the lifeboat are $(u_k, v_k)$; the coordinates of the analysis are denoted $(u^a_k, v^a_k)$; the coordinates of the forecast are denoted $(u^f_k, v^f_k)$.

At the beginning ($t_0$), one has $(u^0_0, v^0_0) = (0, 0)$ by convention. In between times $t_k$ and $t_{k+1}$ we only know that the lifeboat has drifted, but we do not know in which direction. Our model of the lifeboat amounts to a diffusion of the lifeboat position around the origin. In this example, we have $M_{k+1} = I$, the identity matrix. Model error is significant and reads

$$
Q_k = \begin{pmatrix}
\sigma_m^2 & 0 \\
0 & \sigma_m^2
\end{pmatrix}, \tag{2.16}
$$

where $\sigma_m$ measures the uncertainty of the drift magnitude of the lifeboat between $t_k$ and $t_{k+1}$.

Finally, we assume that the analysis and forecast error covariance matrix are diagonal, i.e.

$$
P^a_k = \begin{pmatrix}
\lambda_k & 0 \\
0 & \mu_k
\end{pmatrix}, \quad P^f_k = \begin{pmatrix}
\nu_k & 0 \\
0 & \rho_k
\end{pmatrix}, \tag{2.17}
$$

because no correlation should be induced between the two coordinates of the lifeboat position. This will be verified a posteriori.

**Analysis**

Let us focus on time $t_k$. The previous forecast has yielded state $(u^f_k, v^f_k)^\top$. The castaway proceeds to measure $v^c_k$. He performs an optimal analysis with these observations. The
Kalman gain is obtained from the calculation

\[
K^*_k = \begin{pmatrix}
\nu_k & 0 \\
0 & \rho_k
\end{pmatrix}
\begin{pmatrix}
0 \\
1
\end{pmatrix}
\begin{pmatrix}
\sigma^2_o + (0, 1) \\
\nu_k & 0
\end{pmatrix}
\begin{pmatrix}
0 \\
1
\end{pmatrix}
^{-1}
\]

\[
= \frac{\rho_k}{\sigma^2_o + \rho_k}
\begin{pmatrix}
0 \\
1
\end{pmatrix}.
\]

(2.18)

The lifeboat coordinates are estimated to be

\[
\begin{pmatrix}
u_a_k \\
v_k
\end{pmatrix} = \begin{pmatrix}
u_f_k \\
v_k
\end{pmatrix} + K^*_k \left( y_k - (0, 1) \begin{pmatrix}
u_f_k \\
v_k
\end{pmatrix} \right)
\]

\[
= \begin{pmatrix}
u_f_k \\
v_k
\end{pmatrix} + K^*_k \left( y_k - v^f_k \right)
\]

\[
= \begin{pmatrix}
u_f_k + \frac{\rho_k}{\sigma^2_o + \rho_k} \left( y_k - v^f_k \right) \\
v_k
\end{pmatrix}
\]

(2.19)

From \(P^a_k = (I - K^*_k H_k) P^f_k\) we infer that

\[
\lambda_k = \nu_k \quad \text{and} \quad \frac{1}{\lambda_{k+1}} = \frac{1}{\sigma^2_o} + \frac{1}{\rho_k}.
\]

(2.20)

**Forecast**

Let us now consider the forecast step. The model does not modify the estimate of the position

\[
\begin{pmatrix}
u_{k+1} \\
v_{k+1}
\end{pmatrix} = \begin{pmatrix}
u_k \\
v_k
\end{pmatrix}
\]

(2.21)

Using \(P^f_{k+1} = P^a_k + \sigma^2_m I\), we can calculate the forecast error covariance matrix

\[
\nu_{k+1} = \lambda + \sigma^2_m, \quad \rho_{k+1} = \mu_k + \sigma^2_m.
\]

(2.22)

From the chaining of the analysis step followed by the forecast step, we have

\[
u_{k+1} = \nu_k + \sigma^2_m, \quad \frac{1}{\rho_{k+1} - \sigma^2_m} = \frac{1}{\sigma^2_o} + \frac{1}{\rho_k}.
\]

(2.24)

Since \(u_0^f = 0\), it is clear that \(u_k^f = 0\) for any \(k\): in the absence of observation, nothing is learnt on the coordinate parallel to the shore. However, the uncertainty is a linear increasing function of time since

\[
\nu_k = k \sigma^2_m.
\]

(2.25)

From Eq. (2.24) on the uncertainty of coordinate \(v\), one infers that \(\rho_k \geq \sigma^2_m\). We can look for a fixed point \(\rho_*\) of the recurrence equation satisfying this constraint. It is solution of

\[
\rho_*^2 - \sigma^2_m \rho_* - \sigma^2_o \sigma^2_m = 0
\]

(2.26)
which leads to
\[ \rho_* = \frac{\sigma_m^2}{2} \left( 1 + \sqrt{1 + 4 \frac{\sigma_o^2}{\sigma_m^2}} \right), \] (2.27)

This is the asymptotic result of a compromise between the reduction of uncertainty due to the assimilation of observations, and to the increase of the same uncertainty due to the uncontrolled drift of the lifeboat. A posteriori, we easily check that \( \rho_* \geq \sigma_m^2 \).

2.2.4 Second example: data assimilation with an oscillator

Consider the discrete model
\[ x_0 = 0, \quad x_1 = 1 \quad \text{and for } 1 \leq k \leq K : \quad x_{k+1} - 2x_k + x_{k-1} = -\omega^2 x_k. \] (2.28)

Clearly, this is a numerical implementation of the one-dimensional harmonic oscillator
\[ \ddot{x} + \Omega^2 x = 0. \] (2.29)

This is a discrete second-order equation with \( \Omega^2 \) proportional to \( \omega^2 \). Therefore, a state vector has two entries
\[ u_k = \begin{pmatrix} x_k \\ x_{k-1} \end{pmatrix} \] (2.30)

The resolvent matrix is
\[ M_k = \begin{pmatrix} 2 - \omega^2 & -1 \\ 1 & 0 \end{pmatrix}, \] (2.31)

such that \( u_k = M_k u_{k-1} \) for \( k \geq 1 \). The observation operator is \( H_k = \begin{pmatrix} 1 & 0 \end{pmatrix} \). The observation equation is
\[ y_k = H_k u_k + \xi_k, \] (2.32)

with a Gaussian white noise \( \xi_k \) of variance \( g \). The value of this variance is supposed to be known. Observations may not be available at each time step.

The Kalman filter is tested on this example. Figure 2.2 is a representation of a data assimilation experiment with a given set of observations.

2.3 The extended Kalman filter

Up to now, we have assumed that the hypothesis of linearity is valid. However, the following conditions are frequently met in geophysical data assimilation:

- the observation operator \( H_k \) might be nonlinear. It if often the case when satellite observations are considered. The observation operator that relates optical measurements (radiiances) to the state variables may involve a radiative transfer operator. When diffusion is taken into account, this model might become strongly nonlinear. Lidar and radar observations are other important examples,

- the evolution model \( M_k \) could be nonlinear. This is the case for the primitive equations (atmosphere and ocean). This is also the case in atmospheric chemistry assuming the chemical reactions are second-order.

This does not mean that we have to abandon the Kalman filter. Instead, we need to adapt it to handle the potential nonlinearity of these operators.
2.3.1 Linearising the forecast step and the analysis step

Without mystery, the forecast is given by

\[ x_{k+1}^f = M_{k+1} x_k^a. \]  (2.33)

It can be shown that this estimator is only accurate to the first order in perturbations and that it can be refined at some (important) numerical cost.

To obtain the forecast error covariance matrix, we shall use the fact that \( x_k^a \) is meant to be close enough to \( x_k \). The tangent linear matrix of \( M_k \) at \( x_k \) will be denoted \( M_k \). Let us go again through the forecast step with error

\[
\begin{align*}
\mathbf{e}_{k+1}^f &= x_{k+1}^f - x_{k+1} = M_{k+1}[x_k^a] - x_{k+1} \\
&= M_{k+1}[x_k + (x_k^a - x_k)] - x_{k+1} \\
&\approx M_{k+1}x_k + M_{k+1}(x_k^a - x_k) - x_{k+1} \\
&\approx M_{k+1}\mathbf{e}_k^a - \mathbf{w}_{k+1}. \quad (2.34)
\end{align*}
\]

As a result, we obtain

\[
P_{k+1}^f = M_{k+1}P_k^a M_{k+1}^T + Q_{k+1}. \quad (2.35)
\]

The estimator for the analysis step remains given by an analysis of the form

\[
x_k^a = x_k^f + K_k \left( y_k - H_k [x_k^f] \right). \quad (2.36)
\]
2.3 The extended Kalman filter

We infer

\[ \mathbf{e}_k^a = \mathbf{x}_k^a - \mathbf{x}_k = \mathbf{x}_k^f - \mathbf{x}_k + K_k \left( \mathbf{y}_k - H_k[\mathbf{x}_k] + H_k[\mathbf{x}_k] - H_k[\mathbf{x}_k^f] \right) \]

\[ = \mathbf{x}_k^f - \mathbf{x}_k + K_k \left( \mathbf{y}_k - H_k[\mathbf{x}_k] + H_k[\mathbf{x}_k] - H_k[\mathbf{x}_k^f - \mathbf{x}_k + \mathbf{x}_k] \right) \]

\[ \simeq \mathbf{e}_k^i + K_k \left( \mathbf{e}_k^o - H_k \mathbf{e}_k^i \right). \]  

(2.37)

and

\[ \mathbf{P}_k^a = (\mathbf{I} - K_k \mathbf{H}_k) \mathbf{P}_k^f (\mathbf{I} - K_k \mathbf{H}_k)^T + K_k \mathbf{R}_k K_k^T. \]  

(2.38)

The Kalman gain becomes

\[ K_k^* = \mathbf{P}_k^f H_k^T \left( H_k \mathbf{P}_k^f H_k^T + \mathbf{R}_k \right)^{-1}. \]  

(2.39)

and, consequently, one has

\[ \mathbf{P}_k^a = (\mathbf{I} - K_k^* \mathbf{H}_k) \mathbf{P}_k^f. \]  

(2.40)

The optimal estimator is therefore

\[ \mathbf{x}_k^a = \mathbf{x}_k^f + K_k^* \left( \mathbf{y}_k - H_k[\mathbf{x}_k^f] \right). \]  

(2.41)

Note that the tangent linear of \( H_k \) is used in \( K_k^* \), but it is not used in the innovation vector \( \mathbf{y}_k - H_k[\mathbf{x}_k^f] \).

2.3.2 Summary

These modifications that extent the Kalman filter to the nonlinear context yields the extended Kalman filter.
Extended Kalman filter

1. Initialisation
   - System state $x^f_0$ and error covariance matrix $P^f_0$.

2. For $t_k = 1, 2, ...$
   (a) Analysis
      - Gain computation
        \[ K^*_k = P^f_k H_k^T \left( H_k P^f_k H_k^T + R_k \right)^{-1} \]
      - Analysis computation
        \[ x^a_k = x^f_k + K^*_k \left( y_k - H_k[x^f_k] \right) \]
      - Error covariance matrix computation
        \[ P^a_k = (I - K^*_k H_k) P^f_k \]
   (b) Forecast
      - Forecast state computation $x^f_{k+1} = M_{k+1}[x^a_k]$ 
      - Error covariance matrix computation
        \[ P^f_{k+1} = M_{k+1} P^a_k M_{k+1}^T + Q_{k+1} \]

2.3.3 Data assimilation with a non-harmonic oscillator

The following discrete model

\[ x_0 = 0, \quad x_1 = 1 \quad \text{and for } 1 \leq k \leq K : \quad x_{k+1} - 2x_k + x_{k-1} = \omega^2 x_k - \lambda^2 x^3_k, \quad (2.42) \]

is a numerical implementation of the anharmonic one-dimensional oscillator

\[ \ddot{x} - \Omega^2 x + \Lambda^2 x^3 = 0, \quad (2.43) \]

where $\Omega^2$ is proportional to $\omega^2$ and $\Lambda^2$ is proportional to $\lambda^2$. The related potential is

\[ V(x) = -\frac{1}{2} \Omega^2 x^2 + \frac{1}{4} \Lambda^2 x^4. \quad (2.44) \]

The second term stabilises the oscillator and plays the role of a spring force, whereas the first term destabilises the point $x = 0$, leading to two potential wells. It is a second-order discrete equation, with a state vector that can be advantageously written

\[ u_k = \begin{pmatrix} x_k \\ x_{k-1} \end{pmatrix}. \quad (2.45) \]
2.3 The extended Kalman filter

The state-dependent transition matrix is

\[ M_k = \begin{pmatrix} 2 + \omega^2 - \lambda^2 x_{k-1}^2 & -1 \\ 1 & 0 \end{pmatrix}, \quad (2.46) \]

such that \( u_k = M_k u_{k-1} \) for \( k \geq 1 \). The observation operator is \( H_k = (1, 0) \). The observation equation is

\[ y_k = H_k u_k + \xi_k, \quad (2.47) \]

with a Gaussian white noise \( \xi_k \) of variance \( g \). The value of this variance is supposed to be known. The extended Kalman filter is tested with this nonlinear system. Figure 2.3 gives an example of a data assimilation experiment with a given set of observations.
In addition to the setup parameters given in the text, we choose $K = 1000$ time steps, $\omega = 0.035$ s$^{-1}$, and $\lambda = 0.003$ s$^{-1}$. The dots are the observations, of error variance $g = 7$. The observations are acquired every $\Delta = 25$ time units. The full line represents the forecast of $H_k u_k$ yielded by the extended Kalman filter.
Chapter 3

A generalised variational formalism: 4D-Var

On the one hand, we observed in the previous chapter that optimal interpolation could be generalised to sequential interpolation when

- the system has a temporal dimension,
- its dynamics is described by a model (perfect or not, deterministic or stochastic).

This has led to the Kalman filter formalism.

On the other hand, in chapter 1, we extended the optimal interpolation technique to a variational approach (3D-Var), which offers

- an efficient numerical implementation,
- an (a priori) easy and comprehensive generalisation to nonlinear problems.

Provided some conditions to be discussed later are met, we can contemplate a similar extension for the Kalman filter as seen in chapter 2 by introducing a variational formalism known as 4D-Var.

A rigorous equivalence between the analysis step of 4D-Var and that of the Kalman filter at the end of a specific data assimilation time window can only be proven when the observation operator $H$ and the dynamics $M$ are linear. These operators can depend on time (which is recalled by the notation $H_k$ and $M_k$), and we assume so in this chapter.

In addition, we will assume that the evolution model is deterministic and perfect. That is to say, it mirrors the true model of the system and model errors are negligible. The variational problem is said to be implemented with strong constraints.

3.1 Cost function and how to compute its gradient

In this chapter, we consider a partial trajectory of the system state, i.e. the $K+1$ vectors $x_k$ for each of the $K+1$ instants $t_0, t_1, \ldots, t_K$. Each of these state vectors belong to $\mathbb{R}^{N_x}$.

In addition, we assume that for each of these instants, an observation vector $y_k \in \mathbb{R}^{N_y}$ is acquired. Formally, we can generalise the 3D-Var functional (1.41) to

$$J(x) = \frac{1}{2} (x_0 - x_0^b)^\top B_0^{-1} (x_0 - x_0^b) + \frac{1}{2} \sum_{k=0}^{K} (y_k - H_k x_k)^\top R_k^{-1} (y_k - H_k x_k).$$

(3.1)
This functional is defined under the constraint that the state vectors \( \{x_k\}_{k=0, \ldots, K} \) (from which \( J \) depends) form an admissible trajectory of the system state, that is to say:

\[
\text{for } k = 0, \ldots, K - 1: \quad x_{k+1} = M_{k+1}(x_k).
\] (3.2)

It is momentarily assumed that \( M_k \triangleq M_k \) is linear. Quite often, a background information does not exist for the full trajectory (though the assumption is currently being studied by researchers). However, a background of the initial state of the partial trajectory is very often used. Let us assume that its expectation is \( x_0^b \) and its error covariance matrix is \( B_0 \).

The straightforward way to introduce the model constraint in the cost function is to introduce Lagrange multipliers \( \{\Lambda_k\}_{k=1, \ldots, K} \), where \( \Lambda_k \in \mathbb{R}^{Nx} \). The associated Lagrangian looks like

\[
L(x, \Lambda) = J(x) + \sum_{k=1}^{K} \Lambda_k^\top (x_k - M_k x_{k-1}).
\] (3.3)

This formalism is elegant. Besides, it is quite practical when the variational formalism is in the continuous time limit (time step \( t_{k+1} - t_k \to 0 \)), but with no significant added value in the discrete case, where a straightforward reasoning (without the \( \Lambda_k \)) is as efficient as the Lagrangian approach can be. Indeed, in the time discrete framework, the state variable at time \( t_k \), \( x_k \) can be explicitly written in terms of the initial value \( x_0 \), as the product of the resolvent matrices applied to the initial state, \( i.e. \)

\[
x_k = M_k M_{k-1} \ldots M_1 x_0.
\] (3.4)

Note that, with a continuous time model, the state variables are solutions of a system of ordinary differential equations that are less tractable. As a consequence, the functional (3.1) under constraints is an effective functional of \( x_0 \) that can now be minimised with respect to \( x_0 \). In order to efficiently minimise it, we wish to compute the gradient of \( J \) with respect to the initial vector \( x_0 \). We introduce:

\[
d_k = y_k - H_k [M_k M_{k-1} \ldots M_1] x_0 \quad \text{and} \quad \Delta_k = R_k^{-1} d_k.
\] (3.5)

Introduction to data assimilation
3.1 Cost function and how to compute its gradient

\[ \delta J(x_0) = \delta \left\{ \frac{1}{2} \sum_{k=0}^{K} d_k^\top R_k^{-1} d_k \right\} \]

\[ = \frac{1}{2} \sum_{k=0}^{K} \delta d_k^\top R_k^{-1} d_k + \frac{1}{2} \sum_{k=0}^{K} d_k^\top R_k^{-1} \delta d_k \]

\[ = \sum_{k=0}^{K} (\delta d_k)^\top R_k^{-1} d_k \]

\[ = - \sum_{k=0}^{K} (H_k [M_k M_{k-1} \ldots M_1] \delta x_0)^\top \Delta_k \]

\[ = - \sum_{k=0}^{K} \delta x_0^\top M_k^\top \cdots M_{k-1}^\top M_1^\top H_k^\top \Delta_k. \] (3.6)

We obtain the gradient of the function with respect to \( x_0 \)

\[ \nabla_{x_0} J = - \sum_{k=0}^{K} \left[ M_k^\top \cdots M_1^\top \frac{1}{H_k^\top \Delta_k} \right] H_k^\top \Delta_k \]

\[ = - \left( H_0^\top \Delta_0 + M_1^\top \left[ H_1^\top \Delta_1 + M_2^\top \left[ H_2^\top \Delta_2 + \ldots + [M_K^\top H_K^\top \Delta_K] \ldots \right] \right] \cdots \right) \] (3.7)

This last form of the gradient, which is a Horner factorisation, provides a means to compute it at a lower numerical cost.

M. Bocquet

Figure 3.2: 4D-Var data assimilation compared to data assimilation with a Kalman filter.

d_\text{k} is an innovation vector while \( \Delta_k \) is a normalised innovation vector. For the sake of simplicity, we leave aside the background term in \( J(x) \) because such term has already been studied in the optimal interpolation framework, and it is easy to add its contribution afterwards. We have:

\[ \delta J(x_0) = \delta \left\{ \frac{1}{2} \sum_{k=0}^{K} d_k^\top R_k^{-1} d_k \right\} \]

\[ = \frac{1}{2} \sum_{k=0}^{K} \delta d_k^\top R_k^{-1} d_k + \frac{1}{2} \sum_{k=0}^{K} d_k^\top R_k^{-1} \delta d_k \]

\[ = \sum_{k=0}^{K} (\delta d_k)^\top R_k^{-1} d_k \]

\[ = - \sum_{k=0}^{K} (H_k [M_k M_{k-1} \ldots M_1] \delta x_0)^\top \Delta_k \]

\[ = - \sum_{k=0}^{K} \delta x_0^\top M_k^\top \cdots M_{k-1}^\top M_1^\top H_k^\top \Delta_k. \] (3.6)

We obtain the gradient of the function with respect to \( x_0 \)

\[ \nabla_{x_0} J = - \sum_{k=0}^{K} \left[ M_k^\top \cdots M_1^\top \frac{1}{H_k^\top \Delta_k} \right] H_k^\top \Delta_k \]

\[ = - \left( H_0^\top \Delta_0 + M_1^\top \left[ H_1^\top \Delta_1 + M_2^\top \left[ H_2^\top \Delta_2 + \ldots + [M_K^\top H_K^\top \Delta_K] \ldots \right] \right] \cdots \right) \] (3.7)

This last form of the gradient, which is a Horner factorisation, provides a means to compute it at a lower numerical cost.

M. Bocquet
1. We first compute the $x_k$ recursively thanks to the resolvent matrix $M_k$ and the initial variable $x_0$.

2. Then one computes the normalised innovation vectors $\Delta_k = R_k^{-1} (y_k - H_k x_k)$ that we store.

3. We then define an adjoint state variable $\tilde{x}_k \in \mathbb{R}^{N_x}$. It is defined at the final instant by
$$\tilde{x}_K = H_k^\top K \Delta_K.$$ Knowing $\tilde{x}_{k+1}$, we obtain $\tilde{x}_k$ using
$$\tilde{x}_k = H_k^\top \Delta_k + M_{k+1}^\top \tilde{x}_{k+1}.$$ (3.8)

One can move back in time recursively, back to $\tilde{x}_0$.

4. Finally, we obtain $\nabla_{x_0} J = -\tilde{x}_0$.

To compute the adjoint state variable backward in time, it is necessary to compute the adjoint operator of the resolvent matrices $M_k$. This seems trivial when $M_k$ happens to be a matrix, even a large one since computing its adjoint amounts to taking its transpose. However, in practice, the system’s evolution is computed from a numerical code of several dozen of thousand of lines. Hence, we need to compute the (formally well-established) adjoint of this computer programme. This adjointisation represents a technical and mathematical hardship. Today, it is common to resort to automatic differentiation programmes (see chapter 15 in Blayo et al. (2015) by L. Hacoët) which, from the code of $M_k$, yield the code of $M_k^\top$. Examples of such programmes are: TAPENADE by INRIA (the code to be differentiated can be submitted online and its adjoint code will be returned), TAF (a famous commercial product), but also OpenAD, ADIFOR, ADIC, ADOL-C, etc. More recently, deep learning softwares such as TensorFlow (Google), PyTorch (Facebook) or Theano (University of Montreal) provide adjoint differentiation tools (coined retropropagation) for their neural networks numerical representations. Julia, a recent computer science oriented language, lends itself much more easily to automatic differentiation.

**Gradient with background term**

Accounting for the background term in the cost function is easy; the gradient becomes:

$$\nabla_{x_0} J = B_0^{-1} (x_0 - x_0^b) - \sum_{k=0}^{K} M_{k,0}^\top H_k^\top R_k^{-1} (y_k - H_k M_{k,0} x_0),$$ (3.9)

where for $k > l \geq 0$, we have noted $M_{k,l} = M_k M_{k-1} \cdots M_{l+1}$. By convention, we choose $M_{k,k} = I$. By extension, we impose $M_{k,l} = M_{l,k}^{-1} = M_{k+1}^{-1} M_{k+2}^{-1} \cdots M_l^{-1}$ for $l > k \geq 0$. Owing to this convention, we have

$$\forall k, l \quad 0 \leq k, l \leq n : \quad x_k = M_{k,l} x_l.$$ (3.10)

**3.1.1 What happens when the forward model or the observation model are nonlinear?**

In that case, we can follow the method of section 1.3.3 used so far within the 3D-Var approach. As before, we denote $M$ the tangent linear of $\mathcal{M}$ with respect to $x$. One can
possibly indicate the point where the tangent linear is computed: \( M(x_0) \) for instance when evaluated at \( x_0 \).

The computation of the gradient is not as simple as in the linear case. Mainly, the calculation of \( \delta d_k \) is modified into (Leibnitz’s rule)

\[
\delta d_k = -H_k(x_k)M_k(x_{k-1})M_{k-1}(x_{k-2}) \cdots M_1(x_0) \delta x_0.
\] (3.11)

Then the gradient reads

\[
\nabla_{x_0} J = -K \sum_{k=0}^{K} \left[ M_k^T(x_0)M_k^T(x_1) \cdots M_k^T(x_{k-2})M_k^T(x_{k-1}) \right] H_k^T(x_k) \Delta_k.
\] (3.12)

When considering complex and nonlinear evolution or observation operators, it turns out that we will need not only to know how to derive an adjoint but also the differential (the tangent linear) of the code in the first place. This is again called automatic differentiation.

### 3.2 Solution of the variational problem

Our first objective is to minimise the cost function \( J(x) \). Later, we will show that this variational approach can be equivalent to the Kalman filter under certain conditions.

#### 3.2.1 Solution with respect to \( x_0 \)

We can obtain a solution to the minimisation problem of \( J \) with respect to \( x_0 \) by using our expression of the gradient and solving \( \nabla_{x_0} J = 0 \). To do so, one can define the Hessian of \( J \) over the time interval \([t_0, t_K]\), which will be denoted \( H_{K,0} \). If the cost function is quadratic (which is what we assumed so far), the Hessian matrix does not depend on \( x_0 \) and is formally easily obtained from the gradient:

\[
H_{K,0} = B_0^{-1} + \sum_{k=0}^{K} M_k^T M_k R_k^{-1} H_k M_k^T.
\] (3.13)

With these notations, the solution \( x_0^a \) is easily obtained from (3.9)

\[
x_0^a = x_0^b - H_{K,0}^{-1} \nabla_{x_0} J(x_0^b).
\] (3.14)

This is the Newton solution to the minimisation problem of a quadratic functional.

If the model is nonlinear and the cost-function is non-quadratic, Eq. (3.13) would turn out to only be an approximation of the exact Hessian since second-order derivatives of the model would appear in the exact Hessian. In that case, the Newton method is rather called the Gauss-Newton method.

The analysis error of this estimate is given by the following error covariance matrix

\[
(P_0^a)^{-1} = H_{K,0} = B_0^{-1} + \sum_{k=0}^{K} M_k^T H_k R_k^{-1} H_k M_k^T.
\] (3.15)

Hence, we obtain

\[
(P_0^a)^{-1} = H_{K,0} = B_0^{-1} + \sum_{k=0}^{K} M_k^T H_k R_k^{-1} H_k M_k^T.
\] (3.16)
Solution with respect to $x_k$

When the evolution model is perfect, the (true) values of the system state $x_k$ are all connected thanks to the resolvent matrices $M_{k,l}$. Performing the analysis on the initial condition $x_0$ is therefore equivalent to performing the analysis on any $x_j$, $j \in [0,K]$. To make it clearer, one can rewrite the cost function $(3.1)$ with respect to $x_j$ ($K \geq j \geq 0$) and not $x_0$ anymore following

$$J(x) = \frac{1}{2} \left( M_{0,j} \left( x_j - x_j^b \right) \right) ^\top B_0^{-1} M_{0,j} \left( x_j - x_j^b \right) + \frac{1}{2} \sum_{k=0}^{K} \left( y_k - H_k M_{k,j} x_j \right) R_k^{-1} \left( y_k - H_k M_{k,j} x_j \right),$$

(3.17)

using the convention $(3.10)$. We observe that $M_{0,j} ^\top B_0^{-1} M_{0,j} = \left( M_{j,0} B_0 M_{j,0} ^\top \right)^{-1} \triangleq B_j^{-1}$. The calculation of the gradient with respect to $x_j$ follows

$$-\nabla_{x_j} J = B_j^{-1} \left( x_j^b - x_j \right) + \sum_{k=0}^{K} M_{k,j} ^\top H_k ^\top R_k^{-1} \left( y_k - H_k M_{k,j} x_j \right).$$

(3.18)

This leads to the analysis defined by $\nabla_{x_j} J(x) = 0$,

$$x_j^a = x_j^b - H_{K,j}^{-1} \nabla_{x_j} J(x_j^b),$$

(3.19)

where the Hessian is

$$H_{K,j} = B_k^{-1} + \sum_{k=0}^{K} M_{k,j} ^\top H_k ^\top R_k^{-1} H_k M_{k,j}.$$  

(3.20)

The related analysis error covariance matrix is

$$P_{k}^a = E \left[ (x_j^a - x_j) \left( x_j^a - x_j \right) ^\top \right].$$  

(3.21)

By analogy with the result (1.46), we infer

$$\left( P_{j}^a \right)^{-1} = H_{K,j} B_j^{-1} + \sum_{k=0}^{K} M_{k,j} ^\top H_k ^\top R_k^{-1} H_k M_{k,j}.$$  

(3.22)

This means that the confidence matrix at $x_j^a$ is the sum of the confidence of the background and of the confidence attached to the observations propagated by the inverse evolution model back to time $t_0$.

3.3 Properties

3.3.1 Propagation of the analysis error

It is easy to check that

$$M_{j,l} ^\top H_{K,j} M_{j,l} = H_{K,l}.$$  

(3.23)

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Because of $H_{K,k} = (P_k^a)^{-1}$, it leads to

$$P_l^a = M_{l,j} P_j^a M_{l,j}^T. \quad (3.24)$$

This describes how the analysis error is propagated within the (strong constraint) 4D-Var formalism. Obviously, it is different from the error propagation of the Kalman filter. It originated from the fact that the error analysis in the Kalman filter is causal, meaning that the error depends on the observations before the analysis time, when the 4D-Var analysis error is not causal: the error analysis might depend on future observations.

This relates to one of the weakest point of 4D-Var: it is not simple to estimate the posterior uncertainty.

### 3.3.2 Transferability of optimality

We now have the main tools to demonstrate one of the most appealing property of the minimisation of the functional. We consider a data assimilation window $[t_0, t_K]$. It is split into two periods $[t_0, t_m]$ and $[t_m, t_K]$ with $m$ an integer between 0 and $K$. One can minimise the associated cost function following two paths:

1. The full cost function is minimised over the full interval $[t_0, t_K]$.
2. The minimisation is carried out in two steps. We first minimise the cost function on the interval $[t_0, t_m]$. The result is a solution $x_m^a$ defined at $t_m$, with an error covariance matrix $P_m^a$. Then, a second minimisation is carried out on interval $[t_m, t_K]$, but using the outcome of the first step, $x_m^a$ and $P_m^a$, in a background term in the second cost function.

We would like to show that the two approaches are equivalent.

From what precedes, the result of path (1) is the outcome of the optimisation of functional (3.1) with respect to $x_0$

$$J(x_0) = \frac{1}{2} \left( x_0 - x_0^b \right)^\top B_0^{-1} \left( x_0 - x_0^b \right) + \frac{1}{2} \sum_{k=0}^{K} (y_k - H_k M_{k,0} x_0)^\top R_k^{-1} (y_k - H_k M_{k,0} x_0). \quad (3.25)$$

We can decompose the same cost function into

$$J_{m,0}(x_0) = \frac{1}{2} \left( x_0 - x_0^b \right)^\top B_0^{-1} \left( x_0 - x_0^b \right) + \frac{1}{2} \sum_{k=0}^{m} (y_k - H_k M_{k,0} x_0)^\top R_k^{-1} (y_k - H_k M_{k,0} x_0), \quad (3.26)$$

and

$$J_{K,m}(x_0) = \frac{1}{2} \sum_{k=m+1}^{K} (y_k - H_k M_{k,0} x_0)^\top R_k^{-1} (y_k - H_k M_{k,0} x_0), \quad (3.27)$$

so that we have

$$J(x_0) = J_{m,0}(x_0) + J_{K,m}(x_0). \quad (3.28)$$

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We remark that the functional $J_{m,0}(x_0)$ is the cost function attached to the first step of path (2). Let $x^a_0$ and $P^a_0$ be the result of this analysis. Because of the assumptions of linearity of the operators, $J_{m,0}(x_0)$ is quadratic and coincides with its second-order Taylor expansion at any point, in particular at $x^a_0$. We have

$$J_{m,0}(x_0) = J_{m,0}(x^a_0) + \nabla_{x_0} J_{m,0}(x^a_0) \cdot (x_0 - x^a) + \frac{1}{2} (x_0 - x^a)^T H_{m,0} (x_0 - x^a). \quad (3.29)$$

By definition, $\nabla_{x_0} J_{m,0}(x^a_0) = 0$. As a consequence, minimising $J_{m,0}$ is equivalent to minimising the following square

$$C(x_0) = \frac{1}{2} (x_0 - x^a_0)^T (P^a_0)^{-1} (x_0 - x^a_0), \quad (3.30)$$

with

$$x^a_0 = x^b_0 - H^{-1}_{m,0} \nabla_{x_0} J(x^a_0)$$

$$(P^a_0)^{-1} = H_{m,0} = B_0^{-1} + \sum_{k=0}^m M_{k,0}^T H_k^{-1} R_k^{-1} H_k M_{k,0}, \quad (3.31)$$

since $J_{m,0}(x_0) = J_{m,0}(x^a_0) + C(x_0)$. Hence

$$J(x_0) = J_{m,0}(x^a_0) + C(x_0) + J_{K,m}(x_0). \quad (3.32)$$

But $C(x_0)$ is exactly the background term of the second step of path (2). The result of the global minimisation of $J(x_0)$ over $[t_0, t_K]$ with respect to $x_0$ coincides with the result of the minimisation of the two-step procedure with a background term in the second step to inform about the result of the first step. Note that we could have minimised over $x_K$ as well, $x_m$, or any other $x_k$. This property is called transferability of optimality. However, be aware that there is no identity between

- the outcome of the analysis through path (1) when focused on state vector $x^a_m$ defined at $t_m$, and
- the outcome of the intermediary analysis at time $t_m$ obtained from the first step of path (2).

Indeed, the analysis of an intermediary step through path (1) uses all observations including the ones posterior to the intermediary step, while the analysis of the same intermediary step through path (2), after its first step, ignores future observations.

### 3.3.3 Equivalence between 4D-Var and the Kalman filter

Here, we justify the use of cost function (3.1) by showing that the final result $x^a_K$ at $t_K$ is the same as the sequential analysis of the Kalman filter at $t_K$. We focus on the data assimilation time window $[t_0, t_K]$. For the two methods to be on equal footing, we assume that they use the same background statistics $x^b_0$ and $P^b_0$ at the beginning of the window, i.e. at $t_0$.

We previously showed the transferability of optimality of the variational formalism by splitting the time interval. We can subdivide these two intervals and again invoke the
3.4 Minimisation algorithms for the cost functions

Let us now discuss one practical but essential aspect of the variational methods: the minimisation of the cost functions. Suppose we want to minimise the cost function $J(x)$. If $J$ is quadratic and strictly convex, it has a unique global minimum in $\mathbb{R}^{N_x}$. However, more generally, $J$ can exhibit several local minima. Note that if the problem is physically well posed, one expects that a global minimum exists. Determining all of these minima is a very difficult task. Here, we focus on finding the minimum, or one of the minima. Let us denote $g(x)$ the gradient of $J(x)$, and $H(x)$ is the related Hessian matrix. A way to test the optimisation algorithms is to apply them to the basic case of a quadratic cost function,

$$J_Q(x) = \frac{1}{2} x^\top Q x + b^\top x,$$

where $Q$ is a positive definite matrix. Hence, $J_Q$ is strictly convex.

3.4.1 Descent algorithms

The minimisation algorithms start at a point $x_0 \in \mathbb{R}^{N_x}$ and build a sequence of points $x_k$ which is meant to converge to a local minimum. $x_0$ typically is in the basin of attraction of the local minimum. At step $k$ of the algorithm, we determine a direction $d_k \in \mathbb{R}^{N_x}$ which is characteristic of the method. This direction of exploration is used to define the next point of the sequence

$$x_{k+1} = x_k + \lambda_k d_k,$$

where $\lambda_k$ is a positive real number, $\lambda_k \in \mathbb{R}^+$. An optimal $\lambda_k$ is obtained from a one-dimensional minimisation (line search), along the affine direction defined by $x_k$ and $d_k$. The optimal $\lambda_k$ results from the minimisation of

$$\varphi_k(\lambda) \triangleq J(x_k + \lambda d_k).$$

Since it is a minimisation problem, we wish to have

$$\varphi'_k(0) \triangleq \partial_\lambda \varphi_k(0) = g(x_k)^\top d_k < 0.$$

Hence, for $d_k$ to be a descent direction, we must have $g(x_k)^\top d_k < 0$. 

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Steepest descent, conjugate gradient and Newton-Raphson methods

We can choose for $d_k$ the direction of steepest descent, yielding the \textit{steepest descent method}, $d_k = -g(x_k)$, which is pointed at by minus the gradient of the cost function. We can hope for a better algorithm when the $d_k$ better explores the minimisation space (we have no choice for the $d_k$ in the steepest descent method).

Before considering the full nonlinear case, let us focus on the quadratic case Eq. (3.33). We can impose the $\{d_j\}_{0 \leq j \leq k}$ to be $Q$-conjugate, i.e. they are mutually orthogonal with respect to the quadratic form defined by $Q$, i.e., $d_j^T Q d_l = 0$ whenever $j \neq l$. This construction is analogous to the Gram-Schmidt orthogonalisation. It ensures that the solution can be obtained in less than $N_x + 1$ steps (recall $N_x$ is the dimension of the system state) if the machine precision is infinite. Moreover, it is clear that in the quadratic case, we can obtain an explicit solution for $\lambda_k$ (given below). This method is known as the \textit{conjugate gradient}.

This method also applies to the nonlinear case. In this case, the conjugation is only approximate and the solution usually requires more than $N_x + 1$ steps. The several implementations (Hestenes-Stiefel, Polak-Ribi`ere and Fletcher-Reeves) of this generalisation are distinguished by the explicit formula for $\lambda_k$. The crucial point is that none of these methods make use of the Hessian of the cost function.

In spite of the elegance and sophistication of the conjugate gradient and variants, it may be more efficient to use a Newton-Raphson method which may offer a faster convergence rate. Let us explain its main idea. In a neighbourhood of $x_k$, we wish to approximate the cost function $J$ by a quadratic cost function.

$$J(x) = J(x_k) + g(x_k)^T (x - x_k) + \frac{1}{2} (x - x_k)^T H(x_k) (x - x_k) + o(\|x - x_k\|^2). \quad (3.37)$$

An estimation of the local minimum is given by the minimum of the local quadratic approximation $q(x)$, denoted $x_{k+1}$, that we obtain by solving $\nabla q(x) \simeq 0$, i.e.

$$0 \simeq \nabla q(x) = g(x_k) + H(x_k)(x - x_k). \quad (3.38)$$

If the Hessian is invertible, a better idea of the minimum point $\bar{x}$ is given by

$$x_{k+1} = x_k - H(x_k)^{-1} g(x_k). \quad (3.39)$$

That is to say

$$d_k = -H(x_k)^{-1} g(x_k). \quad (3.40)$$

This corresponds to making a choice not only for $d_k \in \mathbb{R}^{N_x}$, but also for $\lambda_k \in \mathbb{R}^+$ which is taken to be 1. It can be shown that the method is converging provided the starting point $x_0$ is close enough to a local minimum. Applied to a quadratic cost function, this method is qualified as \textit{second-order} because

$$x_{k+1} - \bar{x} = O(\|x_k - \bar{x}\|^2), \quad (3.41)$$

assuming $\bar{x}$ is the solution to the minimisation problem.

This method is meant to be very efficient. It has nevertheless the major drawback that it requires to compute the Hessian and its inverse. For a nonlinear problem, the Hessian must (in principle) be recomputed at each iteration since it is point-dependent. For large dimensional problems, of particular interest to us, this could be a prohibitive computation.
3.4 Minimisation algorithms for the cost functions

3.4.2 Quasi-Newton algorithm

An alternative minimisation technique consists in using at step $k$ a matrix $H_k$ that could easily be computed, and behaves similarly to the inverse of the Hessian in the subspace generated by the gradient $g(x_k)$, that we shall also denote $g_k$, so that we could write $d_k = -H_k g_k$. Please bear in mind the abrupt change of notation: $H_k$ now designates an approximation of the inverse of the Hessian (this is customary).

The idea is to build up $H_k$ along with the iterations. The directions of $\mathbb{R}^{N_x}$ that are explored when we compute $x_{k+1}$ knowing $x_k$ are $\Delta x_k = x_{k+1} - x_k$, and $\Delta g_k = g_{k+1} - g_k$, since we only have access to the $x_k$ and to the gradients $\Delta g_k$. But, a simple exact integration shows that $\Delta x_k$ and $\Delta g_k$ are related via

$$\left\{ \int_0^1 d\lambda \nabla^2 J(x + \lambda d_k) \right\} \Delta x_k = \Delta g_k.$$  (3.42)

The left-hand side $\int_0^1 d\lambda \nabla^2 J(x + \lambda d_k)$ is the mean of the Hessian along the trajectory $\Delta x_k$. The conclusion of this argument is that we can extract from $\Delta g_k$ and $\Delta x_k$ some information about the Hessian. Since this information is of low rank, the inverse of the Hessian needs to be built progressively, iteration after iteration.

Building on this heuristic argument, we wish that the sequence $H_k$ satisfies

$$\forall k \geq 0, \quad H_k (g_{k+1} - g_k) = x_{k+1} - x_k.$$  (3.43)

This condition makes the algorithm a quasi-Newton method. We also wish $H_k$ to be positive semi-definite. Indeed, we have

$$\varphi_k'(0) = g_k^T d_k = -g_k^T H_k g_k,$$  (3.44)

so that the positivity of $H_k$ is sufficient to ensure the descent criterion $\varphi_k'(0) \leq 0$.

In spite of their sophistication, the quasi-Newton methods are very popular, because they combine convergence speed while limiting the algorithmic complexity (no explicit computation of the Hessian). Besides, on the shelf quasi-Newton softwares are available and widely distributed. The generic quasi-Newton method is displayed in algorithm 3.4.2.

The sequence $H_k$ must satisfy the quasi-Newton condition, but this condition does not suffice to fully define it. The complete specification of the sequence distinguishes between variants of the quasi-Newton method.

**Application to a quadratic function**

Let us have a look at the way a quasi-Newton method works in the linear case, i.e. when the cost function is quadratic $J_Q(x) = \frac{1}{2} x^T Q x + b^T x$. The generic algorithm of a quasi-Newton minimisation in the linear case, is given in algorithm 3.4.

The value of $\lambda_k$ is the outcome of the minimisation of

$$J(x_k + \lambda d_k) = \frac{1}{2} (x_k + \lambda d_k)^T Q (x_k + \lambda d_k) + b^T (x_k + \lambda d_k),$$  (3.45)

of obvious exact solution:

$$\lambda_k = -\frac{g_k^T d_k}{d_k^T Q d_k},$$

with $g_k = Q x_k + b$. 

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Generic quasi-Newton algorithm

1. We start at $x_0$ at time $k = 0$
   and we choose an initial symmetric matrix for $H_0$.

2. For $k = 0, 1, 2, \ldots$
   • If $g_k = 0$, the algorithm ends, otherwise
     \[ d_k = -H_k g_k \]
   • We compute (line search)
     \[ \lambda_k = \arg\min \ J(x_k + \lambda d_k) \quad \text{for} \quad \lambda \geq 0 \]
     then
     \[ x_{k+1} = x_k + \lambda_k d_k \]
   • Build $H_{k+1}$ using $H_k$ and
     \[ \Delta x_k = x_{k+1} - x_k \quad \text{and} \quad \Delta g_k = g_{k+1} - g_k \]

Figure 3.3: The generic quasi-Newton scheme.

Q-Conjugation

A remarkable property of the generic quasi-Newton algorithm when applied to the quadratic functional is that the directions $d_k$ are mutually $Q$-conjugate.

3.4.3 A quasi-Newton method: the BFGS algorithm

In practice, one of the most efficient quasi-Newton implementation has been proposed in 1970 by Broyden, Fletcher, Goldfarb and Shanno, hence the name of the method: BFGS.

The sequence of the matrices $H_k$ is defined by the following matrix recursive formula

\[ H_{k+1} = H_k + \frac{1}{\Delta g_k, \Delta x_k} \begin{pmatrix} \Delta g_k, & \Delta g_k \end{pmatrix} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ \frac{1}{\Delta g_k, \Delta x_k} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ -H_k \Delta g_k (\Delta x_k)^T + (H_k \Delta g_k (\Delta x_k)^T) \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \begin{pmatrix} \Delta g_k, & \Delta x_k \end{pmatrix} \]

\[ (3.46) \]

\[ H_{k+1} = H_k + \left(1 - \frac{\Delta x_k \Delta g_k^T}{\Delta g_k, \Delta x_k} \right) H_k \left(1 - \frac{\Delta g_k (\Delta x_k)^T}{\Delta g_k, \Delta x_k} \right) + \frac{\Delta x_k (\Delta x_k)^T}{\Delta g_k, \Delta x_k}. \]

\[ (3.47) \]

The method is rank-2 because

\[ \text{rank} \left( H_{k+1} - H_k \right) \leq 2. \]

\[ (3.48) \]

It is easy to check that the method is quasi-Newton, i.e. it satisfies Eq. (3.43). Moreover, we check that if $H_k$ is positive semi-definite, then $H_{k+1}$ is also positive semi-definite.

Introduction to data assimilation
3.5 Weak-constraint 4D-Var

The variational analogue of the Kalman filter that incorporates model error is given by the weak-constraint 4D-Var of cost function

\[
J(x) = \frac{1}{2} (x_0 - x_0^b)^\top B_0^{-1} (x_0 - x_0^b) + \frac{1}{2} \sum_{k=0}^{K} (y_k - H_k x_k)^\top R_k^{-1} (y_k - H_k x_k) \\
+ \frac{1}{2} \sum_{k=1}^{K} (x_k - M_k x_{k-1})^\top Q_k^{-1} (x_k - M_k x_{k-1}).
\]

(3.50)

Obviously, this method is numerically very costly as one needs to optimise this cost function over all \(x_0, x_1, \ldots, x_K\). A proper modelling of the not so-well known \(Q_k\) is also needed.
Besides, nothing guarantees that it can properly account for any kind of model error. That is why, in spite of being contemplated very early in the data assimilation field, it is still barely used. However, it has more recently drawn attention because it is amenable to a parallelised implementation as opposed to the strong constraint 4D-Var.

The formalism can also be simplified. For instance, one can consider a forcing weak-constraint 4D-Var where model error is given by a systematic bias which does not vary over the time window. The associated cost function reads

\[ J(x) = \frac{1}{2} (x_0 - x_0^b)^\top B_0^{-1} (x_0 - x_0^b) + \frac{1}{2} \sum_{k=0}^{K} (y_k - H_k x_k)^\top R_k^{-1} (y_k - H_k x_k) \\
+ \frac{1}{2} (\eta - \eta^b)^\top Q^{-1} (\eta - \eta^b), \quad (3.51) \]

under the constraints:

\[ \text{for } k = 0, \ldots, K - 1 : \quad x_{k+1} = M_{k+1}(x_k) + \eta. \quad (3.52) \]

Hence, compared to the strong-constraint 4D-Var, only the model error vector \( \eta \in \mathbb{R}^{N_x} \) has been added to the control variables. \( J(x) \) is actually a function of the \( x_0 \) and \( \eta \). This approximation of the weak-constraint 4D-var is in operation at the ECMWF for the stratosphere where the meteorological forecast model (IFS) is known to have systematic biases.
Part II

Advanced methods of data assimilation
Chapter 4

Nonlinear data assimilation

We will illustrate the performance of the algorithms introduced in this chapter with an anharmonic oscillator and some additional model error. That is why we use again the nonlinear oscillator of chapter 2 to which we add some noise corresponding to some model error. This model error, an additive noise, weakens the predictability of the system:

\[
x_0 = 0, \quad x_1 = 1 \quad \text{and for } 1 \leq k \leq K \quad x_{k+1} - 2x_k + x_{k-1} = \omega^2 x_k + \lambda^2 x_k^3 + \xi_k. 
\]

(4.1)

We have chosen the following parameters: \( \omega = 0.035, \lambda = 3.10^{-5} \). The variance of model error is 0.0025 while the variance of the observation error is 49. The initial condition is \( x_0 = 0 \) and \( x_1 = 1 \). The system is iterated 10000 times.

4.1 The limitations of the extended Kalman filter

Let us look at the typical behaviour of an extended Kalman filter applied to a strongly nonlinear system. Let us experiment with the stochastic anharmonic oscillator described above. At first, we choose the observation to be performed every 50 timesteps. This yields the typical run shown in figure 4.1, where the extended Kalman filter succeeds in tracking the system's state.

Secondly, we choose to assimilate an observation every 58 iterations, at a somewhat lower frequency. The behaviour of the extended Kalman filter is illustrated in figure 4.2. We see that the filter ultimately diverges because it does not have enough information on the system state's trajectory and because the error propagation is approximated by the tangent linear model in between two analysis steps.

To reduce the impact of the tangent linear approximation, we could resort to higher order extended Kalman filters, which better handle the nonlinearity of the model. However, these require the use of not only the tangent linear but also an Hessian, which would be very costly to implement (in terms of CPU and memory). This is a stringent limitation of these methods for high dimensional systems.

To understand and improve the performance of the filters with strongly nonlinear dynamical systems, it would be interesting to define a data assimilation method able to account for all statistical moments of the state distribution. Indeed a nonlinear dynamics propagates the statistical moments of the system state in a non-trivial fashion as opposed to linear dynamics.
4.2 The ultimate data assimilation method?

So far, we have tried to solve the following problem: given an a priori but partial and imperfect knowledge of the system and a fresh observation set, what is, in the mean square sense, the optimal system state estimate and what is the variance of the analysis error?

Yet, the question could be more general. We could reformulate the problem in terms of probabilities. We do not limit ourselves to the expectation and covariances of the variable distributions but we work on all statistical moments (Tarantola and Valette, 1982).

Provided we know the probability density of the background as well as that of the observations, what is the probability density of a system state? A opposed to the approaches of the previous chapters, there is not any such a priori criterion such as the least squares from which we could define an estimator for the system state and its error. Instead, we will need to choose a criterion that defines optimality, which yields a specific estimator.

4.2.1 Assimilation of an observation

Let us consider a very general observation equation

\[ y = H(x, v), \]

where the observation operator \( H \) could be nonlinear. We assume that we know a priori the probability density function (pdf) of the distribution of \( x \), i.e. the pdf of the background, which is denoted \( p_X(x) \). Besides, the pdf of the noise, denoted \( p_V(v) \) is also assumed to be known. We wish to know the probability density function of \( x \), knowing the value of the observation vector \( y \), i.e. the conditional pdf of \( x \), knowing \( y \), denoted \( p_{X|Y}(x|y) \). To that goal, let us consider the joint pdf of the state and the observation vector, \( p_{X,Y}(x,y) \).
Figure 4.2: Data assimilation with a stochastic anharmonic oscillator, using an extended Kalman filter. The observation frequency is one observation every 58 timesteps which is insufficient and the filter ultimately diverges.

We can easily decompose it in two ways:

\[ p_{X,Y}(x,y) = p_{X|Y}(x|y)p_Y(y) = p_{Y|X}(y|x)p_X(x). \]  

(4.3)

This yields the infamous Bayes’ rule (that was later rediscovered by Pierre-Simon Laplace in a more general form than initially proposed by Thomas Bayes), we obtain:

\[ p_{X|Y}(x|y) = p_{Y|X}(y|x) \frac{p_X(x)}{p_Y(y)}. \]  

(4.4)

which determines \( p_Y \). Computing this latter term corresponds to a normalisation of the pdf with respect to \( x \):

\[ p_Y(y) = \int p_{Y|X}(y|x)p_X(x) \, dx. \]  

(4.5)

In practice it is often unnecessary to compute this normalisation since it does not depend on \( x \).

### 4.2.2 Estimation theory and BLUE analysis

To convince ourselves of the richness of this probabilistic standpoint, let us check that it is consistent with the optimal interpolation (BLUE). To this end, the observation equation is simplified into

\[ y = Hx + v, \]  

(4.6)

where the observation operator \( H \) is assumed linear here. We also assume that \( x \) is distributed according to a Gaussian distribution \( N(\bar{x}, P) \), that is to say according to the

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(normalised) pdf

\[ p_X(x) = \frac{1}{(2\pi)^{N_x/2}|P|^{1/2}} \exp\left(-\frac{1}{2}(x - \bar{x})^\top P^{-1}(x - \bar{x})\right), \tag{4.7} \]

with \(|P|\) denoting the determinant of the square matrix \(P\). This constitutes the background pdf of \(x\). Since \(v\) is distributed according to \(N(0, R)\), independent from that of \(x\), the conditional probability of \(y\), knowing \(x\), is

\[ p_{Y|X}(y|x) = \frac{1}{(2\pi)^{N_y/2}|R|^{1/2}} \exp\left(-\frac{1}{2}(y - Hx)^\top R^{-1}(y - Hx)\right). \tag{4.8} \]

The calculation of \(p_Y\) (normalisation of the product of the two previous densities) consists in convolving two Gaussian pdfs. As a consequence, the resulting density is also Gaussian in \(y\). It is easy to compute its expectation \(\bar{y} = E[y]\) and its variance \(P_y\). Indeed,

\[
\begin{align*}
\bar{y} &= E[Hx + v] = H\bar{x}, \\
P_y &= E[(y - \bar{y})(y - \bar{y})^\top] = E[(H(x - \bar{x}) + v)(H(x - \bar{x}) + v)^\top] \\
&= HE[(x - \bar{x})(x - \bar{x})^\top]H^\top + E[vv^\top] = HPH^\top + R. \tag{4.9}
\end{align*}
\]

This entirely characterises the distribution of \(y\)

\[ p_Y(y) = \frac{1}{(2\pi)^{N_y/2}|HPH^\top + R|^{1/2}} \exp\left(-\frac{1}{2}(y - H\bar{x})^\top(HPH^\top + R)^{-1}(y - H\bar{x})\right). \tag{4.10} \]

Using (4.10), (4.8) and Bayes' rule (4.4), we obtain the targeted conditional density:

\[ p_{X|Y}(x|y) = \frac{|HPH^\top + R|^{1/2}}{(2\pi)^{N_x/2}|P|^{1/2}|R|^{1/2}} \times \exp\left(-\frac{1}{2}(y - Hx)^\top R^{-1}(y - Hx) - \frac{1}{2}(x - \bar{x})^\top P^{-1}(x - \bar{x})
\]

\[ + \frac{1}{2}(y - H\bar{x})^\top HPH^\top + R)^{-1}(y - H\bar{x})\right). \tag{4.11} \]

Factorising the argument of the exponential according to \(x\), we obtain

\[ p_{X|Y}(x|y) = \frac{|HPH^\top + R|^{1/2}}{(2\pi)^{N_x/2}|P|^{1/2}|R|^{1/2}} \exp\left(-\frac{1}{2}(x - x^*)^\top P_*^{-1}(x - x^*)\right), \tag{4.12} \]

with

\[ P_*^{-1} = P^{-1} + H^\top R^{-1}H \quad \text{and} \quad x^* = P_*(H^\top R^{-1}y + P^{-1}\bar{x}). \tag{4.13} \]

### 4.2.3 Choosing an estimator

Then one needs to choose an estimator. That is, one needs to make rigorous the subjective concept of most likely value of an analysis of the system state \(x\), knowing the observations \(y\). This estimator is to be defined from the conditional pdf \(p_{X|Y}\). We would
like the estimator to be unbiased. In the present case, since the distributions are Gaussian, all reasonable choices lead to the same estimator. However, with more general distributions, we would have to distinguish between the minimum variance estimator and the maximum a posteriori which corresponds to the maximum of the pdf $p_{X|Y}$. But other choices are possible.

In the present case, the two coincide with the estimator $\hat{x} = E[x|y]$, which is easily read from the pdf. Now, using the optimal gain $K^*$, we obtain

$$\hat{x} = (P^{-1} + H^T R^{-1} H)^{-1} (H^T R^{-1} y + P^{-1} x)$$

$$= K^* y + (P^{-1} + H^T R^{-1} H)^{-1} P^{-1} x$$

$$= x + K^* (y - H x)$$

(4.14)

where we used relationships (1.26) and (1.24). This result should be compared with (1.18). The result of this probabilistic calculation hence coincides with the result of BLUE, assuming $x_b = x$.

### 4.3 Sequential assimilation and probabilistic interpretation

Let us consider a sequential data assimilation scheme. We wish to study the scheme using probability theory following the ideas of the previous section. This is the Bayesian filtering problem. Let us denote $Z_k$, the set of all past observation vectors from $t_0$ to $t_k$, i.e.

$$Z_k = \{z_k, z_{k-1}, \cdots, z_0\}.$$  

(4.15)

We wish to compute the conditional pdf of $x_k$ knowing $Z_k$, denoted $p(x_k|Z_k)$. For the sake of simplicity, we give up the pdf index. Therefore, they will all be denoted $p$, and only their argument can help discriminate them.

#### 4.3.1 Forecast step

In the forecast step, one wishes to compute the pdf $p(x_{k+1}|Z_k)$, knowing the pdf $p(x_k|Z_k)$. Without any particular restriction, we have

$$p(x_{k+1}|Z_k) = \int dx_k p(x_{k+1}|x_k) p(x_k|Z_k).$$

(4.16)

A very general stochastic modelling of the dynamical system would be

$$x_{k+1} = F(x_k, w_k),$$

(4.17)

where $w_k$ is meant to be a white noise (uncorrelated in time) of density $p_W$. Thus we have

$$p(x_{k+1}|x_k) = \int dw_k p_W(w_k) \delta (x_{k+1} - F(x_k, w_k)),$$

(4.18)

where $\delta$ is the Dirac distribution.

Hence, the densities change between $t_k$ and $t_{k-1}$ according to

$$p(x_{k+1}|Z_k) = \int dx_k dw_k p_W(w_k) p(x_k|Z_k) \delta (x_{k+1} - F(x_k, w_k)).$$

(4.19)
When the noise is additive, i.e. if (4.17) becomes
\[ x_{k+1} = M(x_k) + w_k, \] (4.20)
the convolution of densities simplifies into
\[ p(x_{k+1}|Z_k) = \int dx_k p_W(x_{k+1} - M(x_k)) p(x_k|Z_k). \] (4.21)

When model error is zero, we can further simplify since
\[ p_W(x_{k+1} - M(x_k)) = \delta(x_{k+1} - M(x_k)). \] (4.22)

In that case:
\[ p(x_{k+1}|Z_k) = p\left(\frac{M_k^{-1}x_{k+1}|Z_k)}{(M_k')^{-1}M_k^{-1}(x_{k+1})}\right), \] (4.23)
where $|O|$ is the determinant of $O$.

### 4.3.2 Analysis step

Once a new observation vector is acquired, an analysis is performed. A general observation equation is
\[ z_k = H(x_k, v_k), \] (4.24)
where $v_k$ is a noise of arbitrary distribution. Using Bayes’ rule, we obtain:
\[ p(x_k|Z_k) = p(x_k|z_k, Z_k_{k-1}) = p(z_k|x_k, Z_k-1) \frac{p(x_k|Z_k_{k-1})}{p(z_k|Z_k_{k-1})}. \] (4.25)

It is slightly more elaborated than formula (4.4). $p(x_k|Z_{k-1})$ is supposed to be known since it is the outcome pdf of the forecast from $t_{k-1}$ to $t_k$. We remark that we have
\[ p(z_k|Z_{k-1}) = \int dx_k p(z_k|x_k, Z_k_{k-1}) p(x_k|Z_k_{k-1}), \] (4.26)
which looks like the normalisation of the targeted pdf. In summary
\[ p(x_k|Z_k) = \frac{p(z_k|x_k, Z_k_{k-1}) p(x_k|Z_k_{k-1})}{\int dx_k p(z_k|x_k, Z_k_{k-1}) p(x_k|Z_k_{k-1})}. \] (4.27)

If we assume the noise $v_k$ to be white in time, $p(z_k|x_k, Z_k_{k-1}) = p(z_k|x_k)$, since the knowledge of $p(z_k|x_k)$ is equivalent to that of the white noise at time $t_k$, which is independent from $Z_{k-1}$. In that case, the previous formula simplifies into
\[ p(x_k|Z_k) = \frac{p(z_k|x_k) p(x_k|Z_k_{k-1})}{\int dx_k p(z_k|x_k) p(x_k|Z_k_{k-1})}. \] (4.28)

Provided we can afford the computations, these forecast and analysis steps are sufficient to cycle the system’s state pdf and perform sequential data assimilation.

Introduction to data assimilation
4.3.3 Estimation theory and Kalman filter

Let us again assume the conditions of validity of the Kalman filter, i.e. the linearity of operators. We would like to show that, essentially under these conditions, the estimation theory yields the Kalman filter results. Hence, this is a particular implementation of the previous forecast and analysis steps.

**Forecast step**  The model equation (transition from $t_k$ to $t_{k+1}$) is

$$x_{k+1} = M_k x_k + w_k,$$  \hspace{1cm} (4.29)

where $w_k$ is a Gaussian white noise of covariance matrix $Q_k$. By construction, we have in one hand

$$p_W(x_{k+1} - M_k x_k) = \frac{1}{(2\pi)^{N_x/2}|Q_k|^{1/2}} \exp \left( -\frac{1}{2} (x_{k+1} - M_k x_k) \top (Q_k)^{-1} (x_{k+1} - M_k x_k) \right),$$  \hspace{1cm} (4.30)

and

$$p(x_k | Z_k) = \frac{1}{(2\pi)^{N_x/2}|P_{a_k}^{a}|^{1/2}} \exp \left( -\frac{1}{2} (x_k - x_{a_k}^a) \top (P_{a_k}^{a})^{-1} (x_k - x_{a_k}^a) \right),$$  \hspace{1cm} (4.31)

in the other hand, with $x_{a_k}^a$ and $P_{a_k}^{a}$ outcome of the analysis at time $t_k$. Integral (4.21) remains to be calculated. Again, we have to convolve two Gaussian distributions. The calculation is formally equivalent to (4.7), (4.8) and (4.10). We obtain

$$p(x_{k+1} | Z_k) = \frac{1}{(2\pi)^{N_x/2}|P_{k+1}^{f}|^{1/2}} \exp \left( -\frac{1}{2} (x_{k+1} - x_{k+1}^f) \top (P_{k+1}^{f})^{-1} (x_{k+1} - x_{k+1}^f) \right),$$  \hspace{1cm} (4.32)

with

$$x_{k+1}^f = M_k x_k^a \quad \text{and} \quad P_{k+1}^{f} = M_k^\top P_{a_k}^{a} M_k + Q_k.$$  \hspace{1cm} (4.33)

**Analysis step:** There is nothing to prove for the analysis step, since this part coincides with the derivation of the BLUE analysis by estimation theory, which was described in section 4.2.2.

Hence, we recover that estimation theory yields the Kalman filter in the linear case and with Gaussian a priori statistics. A possible estimator is that of the maximum a posterior (MAP).

4.4 Variational data assimilation and probabilistic interpretation

Thinking in terms of probability, we wish that the nonlinear data assimilation system

$$\begin{align*}
x_{k+1} &= M_{k+1}(x_k) + w_k \\
z_k &= H_k(x_k) + v_k,
\end{align*}$$  \hspace{1cm} (4.34)

has a 4D-Var formulation over the interval $[t_0, t_K]$. The models $M_k$ and $H_k$ could be nonlinear. Model error noise $w_k$ as well as observation error $v_k$ are supposed to be
Gaussian white noise, unbiased and with error covariance matrices $Q_k$ and $R_k$ respectively. Their probability density functions are

$$p_W(w_k) = N(0, Q_k) \quad \text{and} \quad p_V(v_k) = N(0, R_k).$$

(4.35)

Differently from the targeted pdf of sequential assimilation, we rather wish to compute $p(X_k|Z_k)$, a Bayesian smoothing problem, with

$$X_k = \{x_k, x_{k-1}, \ldots, x_0\}.$$  

(4.36)

This requires for us to adapt the probabilistic calculation of the sequential approach. First of all, the Bayes’ rule gives

$$p(X_k|Z_k) = p(Z_k|X_k) \frac{p(X_k)}{p(Z_k)}.$$  

(4.37)

Since we are only interested in the dependence in $X_k$ of $p(X_k|Z_k)$, we will target the following pdf $p(X_k|Z_k) \propto p(Z_k|X_k) p(X_k)$. It is clear that

$$p(X_k) = p(x_k, x_{k-1}) = p(x_k|x_{k-1}) p(X_{k-1}) = p(x_k|x_{k-1}) p(x_k|Z_{k-1}),$$  

(4.38)

where we used the fact that the sequence of random variables $X_k$ has the Markov property. Moreover, we can simplify $p(Z_k|X_k)$:

$$p(Z_k|X_k) = p(z_k, z_{k-1}|X_k) = p(z_k|X_k) p(z_{k-1}|X_k)$$

$$\quad = p(z_k|x_k, x_{k-1}) p(z_{k-1}|x_k, x_{k-1})$$

$$\quad = p(z_k|x_k) p(z_{k-1}|x_{k-1}).$$  

(4.39)

We infer that

$$p(X_k|Z_k) \propto p(Z_k|X_k) p(X_k) = p(z_k|x_k) p(z_{k-1}|x_{k-1}) p(x_k|x_{k-1}) p(X_{k-1})$$

$$\quad \propto p(z_k|x_k) p(x_k|x_{k-1}) p(Z_{k-1}|x_{k-1}) p(X_{k-1})$$

$$\quad \propto p_V(z_k - H(x_k)) p_W(x_k - M_k(x_{k-1})) p(Z_{k-1}|X_{k-1}) p(X_{k-1}).$$  

(4.40)

This recursive relationship is sufficient to determine the density $p(X_k|Z_k)$ taking into account the hypotheses on the noise:

$$p(X_k|Z_k) \propto \left[ \prod_{j=0}^{k} p_V(z_j - H_j(x_j)) \prod_{j=0}^{k-1} p_W(x_{j+1} - M_{j+1}(x_j)) \right] p(x_0)$$

$$\quad \propto \exp \left[ -\frac{1}{2} \sum_{j=0}^{k} (z_j - H_j(x_j))^\top R_j^{-1} (z_j - H_j(x_j)) \
\quad -\frac{1}{2} \sum_{j=0}^{k-1} (x_{j+1} - M_{j+1}(x_j))^\top Q_j^{-1} (x_{j+1} - M_{j+1}(x_j)) \right] p(x_0).$$  

(4.41)

The probability density $p(x_0)$ defines the system state background at time $t_0$. If we choose the MAP as the estimator, we need to maximise the log-density (the argument of the exponential). The negative log-density $-\ln p(X_k|Z_k)$ is merely the 4D-Var cost function (to be minimised!). This justifies the use of this functional in the 4D-Var formalism even if the model operators are nonlinear. This extends the results of chapter 3. Besides, this is another proof of the equivalence between 4D-Var and the Kalman filter in a linear context.
4.5 Particle filters (Monte Carlo)

Can we actually conceive a numerical algorithm that converges to the solution of the Bayesian filtering problem? Such numerical approach would likely belong to the Monte Carlo methods. This means that a probability density function is represented by a discrete sample of the targeted pdf. Rather than trying to compute the exact solution of the Bayesian filtering equations, the transformations of such filtering (Bayes’ rule for the analysis, model propagation for the forecast) are applied to the members of the sample. The statistical properties of the sample, such as the moments, are meant to be those of the targeted pdf. Of course this sampling strategy can only be exact in the asymptotic limit, that is in the limit where number of members (or particles) is going to infinity.

The most popular and simple algorithm of Monte Carlo type that solves the Bayesian filtering equations, is called the **bootstrap particle filter** (Gordon et al., 1993). Its description follows.

**Sampling**  Let us consider a sample of particles \{x_1, x_2, \ldots, x_M\}. The related probability density function at time \(t_k\) is \(p_k(x): p_k(x) \simeq \sum_{i=1}^{M} w_i^k \delta(x - x_i^k)\), which is meant to be an approximation of the exact density that the samples emulates. \(w_i^k\) is a scalar number which weights the importance of particle \(i\) within the ensemble. At this stage, we assume that the weights \(w_i^k\) are uniform \(w_i^k = 1/M\).

**Forecast**  At the forecast step, the particles are propagated by the model without approximation, \(i.e. p_{k+1}(x) \simeq \sum_{i=1}^{M} w_i^k \delta(x - x_{i, k+1}^k)\), with \(x_{i, k+1}^k = M_{k+1}(x_k)\).

**Analysis**  The analysis step of the particle filter is extremely simple and elegant. The rigorous implementation of Bayes’ rule makes us attach to each particle a specific statistical weight that corresponds to the likelihood of the particle with respect to the data. The weight of each particle is changed according to

\[
 w_{k+1}^{i} \propto w_{k}^{i} \cdot p(y_{k+1}|x_{k+1}^{i}).
\]  (4.42)

**Resampling**  Unfortunately, these statistical weights have a potentially large amplitude of fluctuation. Worse, as sequential filtering goes on, one particle, \(i.e. one trajectory of the model will stand out among the others. Its weight will largely dominate the others \(w_i \lesssim 1\) while the other weights vanish. Then, the particle filter becomes very inefficient as an estimating tool since it loses any variability. This phenomenon is called **degeneracy** of the particle filter. An example of such degeneracy is given in Fig. 4.3, where the statistical properties of the biggest weight is studied on a meteorological toy model of 40 and 80 variables. In a degenerate case, the maximum weight will often reach 1 or close to 1, whereas in a balanced case, values very close to 1 will be less frequent.

One way to mitigate this impeding phenomenon is to resample the particles by redrawing a sample with uniform weights from the degenerate distribution. After resampling, all particles have the same weight, \(w_i^k = 1/M\).

The principle of the bootstrap particle filter is illustrated in Fig. 4.4.

The particle filter is very efficient on highly nonlinear system but of low dimension. Unfortunately, it is not suited for high-dimensional systems, as soon as the dimension gets
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Figure 4.3: On the left: statistical distribution of the weights of the particle bootstrap filter in a balanced case. The physical system is a Lorenz 1995 model with 40 variables (Lorenz and Emmanuel, 1998). On the right: the same particle filter is applied to a Lorenz 1995 low-order model, but with 80 variables. The weights clearly degenerate with a peak close to 1.

over, say about 10. Avoiding degeneracy requires a great many particles. Roughly, this number increases exponentially with the system state space dimension!

Figure 4.4: Scheme of the bootstrap particle filter.

For the forecast step, it is also crucial to introduce stochastic perturbations of the states. Indeed the ensemble will impoverish with the many resampling to undergo. To enrich the sample, it is necessary to stochastically perturb the states of the system.

Finally, even if it seems inefficient to apply such particle filters, tests have been successfully run in oceanography (Van Leeuwen, 2003) that are promising in the long run.

Introduction to data assimilation
Chapter 5

The ensemble Kalman filter

In chapter 3, we found that the Kalman filter had two major drawbacks. Firstly, it is numerically hardly affordable in high-dimensional systems. One has to manipulate the error covariance matrix, which requires $N_x(N_x + 1)/2$ scalars to be stored. Clearly this is impossible for high-dimensional systems for which the storage of a few state vectors is already challenging. Moreover, during the forecast step from $t_k$ to $t_{k+1}$, one has to compute a forecast error covariance matrix of the form (see chapter 3)

$$P_{f_{k+1}} = M_{k+1} P_{a_{k}} M_{k+1}^\top + Q_k.$$  

Such computation would require to use the model $2N_x$ times (a left matrix multiplication by $M_{k+1}$, followed by a right matrix multiplication by $M_{k+1}^\top$). For a high-dimension system, this is likely to be much too costly, even if leveraging on parallel computing.

Another drawback of the Kalman filter is that its extension to nonlinear models, namely the extended Kalman filter, is an approximation. It makes use of the tangent linear of the model. When the tangent linear approximation is breached, for instance when the timestep between two consecutive updates is long enough, the extended Kalman filter may yield a too coarse approximation of the forecast error covariance matrix, which may induce the divergence of the filter.

5.1 The reduced rank square root filter

The reduced rank square root filter, denoted RRSQRT, is a solution to the main problem of the dimensionality. The issue of the computation and propagation of the error covariance matrices is astutely circumvented. The error covariance matrices are represented by their principal axes (those directions with the largest eigenvalues), that is to say by a limited selection of modes. The update and the forecast step will therefore bear on a limited number of modes, rather than on matrices (i.e. a collection of $N_x$ modes for a state space of dimension $N_x$).

The initial system state is $x_f^0$, with an error covariance matrix $P_f^0$. We assume a decomposition in terms of the principal modes, $P_f^0 : P^0_0 \simeq S^0_f (S^0_f)^\top$, where $S^0_f$ is a matrix of size $N_x \times N_m$ with $N_m N_x$-vector columns that coincide with the $N_m$ first dominant modes of $P^0_0$. For instance, this could be obtained through the diagonalisation of the
symmetric semi-definite positive $P_0^f$:

$$P_0^f = \sum_{n=1}^{N_N} \sigma_n s_n s_n^\top \approx \sum_{i=1}^{N_m} \sigma_i s_i s_i^\top,$$

(5.2)

where the eigenvalues are arranged by decreasing order $\sigma_1 \geq \sigma_2 \geq \ldots \sigma_{N_N} \geq 0$. We could choose $S_0^f = \left[ \sqrt{\sigma_1} s_1 \quad \sqrt{\sigma_2} s_2 \quad \ldots \quad \sqrt{\sigma_{N_N}} s_{N_N} \right]$ so that indeed $P_0^f \approx S_0^f (S_0^f)^\top$.

Next, we assume that such decomposition persists at later time: $P_k^f \approx S_k^f (S_k^f)^\top$, or getting rid of the time index, $P_f^f \approx S_f S_f^\top$. The background representation has been simplified. Rather than thinking in terms of $P_f^f$, we think about its dominant modes $S_f$, as well as the transformed matrix of size $N_M \times N_M$ in observation space, $Y_f = HS_f$. $H$ is either the observation operator when it is linear, or its tangent linear. These matrices appear when computing the Kalman gain in the analysis step,

$$K^* = P_f^f H^\top \left( H P_f^f H^\top + R \right)^{-1}$$

$$= S_f S_f^\top H^\top \left( H S_f S_f^\top H^\top + R \right)^{-1}$$

$$= S_f (HS_f)^\top \left( (HS_f) (HS_f)^\top + R \right)^{-1}.$$

(5.3)

Hence, the Kalman gain, computed at the analysis step, is simply expressed with the help of the $Y_f$ matrix:

$$K^* = S_f Y_f^\top \left( Y_f Y_f^\top + R \right)^{-1}.$$

(5.4)

The analysis estimate $\mathbf{x}^a$ can be obtained from this gain and the standard update formula.

Then, what happens to the formula that gives the analysis error covariance matrix $P^a$? We have

$$P^a = (I_x - K^* H) P_f^f$$

$$= \left( I_x - S_f Y_f^\top (Y_f Y_f^\top + R)^{-1} H \right) S_f S_f^\top$$

$$= S_f \left( I_m - Y_f^\top (Y_f Y_f^\top + R)^{-1} Y_f \right) S_f^\top.$$

(5.5)

We now look for a square root matrix $S^a$ such that $S^a (S^a)^\top = P^a$. One such matrix is

$$S^a = S_f \left( I_m - Y_f^\top (Y_f Y_f^\top + R)^{-1} Y_f \right)^{1/2}.$$

(5.6)

Of size $N_M \times N_m$, $S^a$ represents a collection of $m$ state vectors, that is to say a posterior ensemble. This avoids a brutal calculation of the error covariance matrices. The computation of the square root matrix might look numerically costly. This is not the case though since $Y_f$ is of reduced size and $I_m - Y_f^\top (Y_f Y_f^\top + R)^{-1} Y_f$ is of dimension $N_m \times N_m$. In addition, the conditioning of the square root matrix is better than that of the initial error covariance matrix (it is the square root of the original conditioning). This ensures an improved numerical precision.

After the analysis step, we seek to reduce the dimension of the system. We wish to shrink the number of modes from $N_m$ to $N_m - N_q$. We first diagonalise $S_a^\top S_a = V \Lambda V^\top$. 

Introduction to data assimilation
We consider the first $N_m - N_q$ eigenmodes with the largest eigenvalues. We keep the first $N_m - N_q$ eigenvectors, which correspond to the first $N_m - N_q$ columns of $\mathbf{V}$ if the diagonal entries of $\mathbf{A}$ are stored in decreasing order. These $N_m - N_q$ vectors are stored in the $(N_m - N_q) \times N_x$ matrix $\mathbf{V}$. Then $\mathbf{S}^a$ is reduced to $\tilde{\mathbf{S}}^a \equiv \mathbf{S}^a \mathbf{V}$.

At the forecast step, the analysis is forecasted through $x_{k+1}^f = M_{k+1}(x_k^a)$. The square root matrix $\mathbf{S}_k^a$ is propagated with the tangent model. Then, the matrix is enlarged by adding $N_q$ modes that are meant to introduce model error variability. The matrix of these augmented modes is of the form

$$
\mathbf{S}_k^f = [M_{k+1}\tilde{\mathbf{S}}_k^a, \mathbf{T}_k],
$$

(5.7)

where $\mathbf{M}$ is the tangent linear model of $\mathcal{M}$. It has $N_m$ modes, so that the assimilation procedure can be cycled.

This type of filter has been put forward in air quality where the number of chemical species considerably increases the state space dimension (Segers, 2002). A similar filter, known as SEEK, has also been used in oceanography for the same reasons (Pham et al., 1998).

These filters clearly overcome the main drawback of the Kalman filter, assuming the dynamics of the system can indeed be represented with a limited $N_m \ll N_x$ number of modes. However, by still making use of the tangent linear model, they only approximate the nonlinear propagation of uncertainty. Moreover, it requires to develop the tangent linear of the model. In that respect, one can propose a better reduced Kalman filter, known as the ensemble Kalman filter.

### 5.2 The stochastic ensemble Kalman filter

The ensemble Kalman filter was proposed by Geir Evensen in 1994, and later amended in 1998 (Evensen, 1994; Burgers et al., 1998; Houtekamer and Mitchell, 1998; Evensen, 2009). Its semi-empirical justification could be disconcerting and not as obvious as that of the RRSQRT. Nevertheless, the ensemble Kalman filter has proven very efficient on a large number of academic and operational data assimilation problems.

The ensemble Kalman filter (EnKF in the following) is a reduced-order Kalman filter, just like the RRSQRT filter. It only handles the error statistics up to second order. Therefore, the EnKF is not a particle filter of the family that we discussed in chapter 4. It is a Gaussian filter. It has been shown that in the limit of a large number of particles, the EnKF does not solve the Bayesian filtering problem as exposed in chapter 4, as opposed to the particle filter, except when the models are linear and when the initial error distribution is Gaussian.

Yet, just as the particle filter and the RRSQRT filter, the EnKF is based on the concept of particles, a collection of state vectors, the members of the ensemble. Rather than propagating huge covariance matrices, the errors are emulated by scattered particles, a collection of state vectors whose dispersion (variability) is meant to be representative of the uncertainty of the system’s state. Just like the particle filter, but unlike the RRSQRT, the members are to be propagated by the model without any linearisation.
Reduced rank square root (RRSQRT) filter

1. Initialisation
   - System state $x^f_0$ and error covariance matrix $P^{f}_0$.
   - Decomposition: $Q_k = T_k T_k^T$ and $P^{f}_0 = S^{f}_0 S^{f}_0^T$.

2. For $t_k = 1, 2, \cdots$
   (a) Analysis
      - Gain computation:
        $$ Y_k = H_k S^f_k, \quad K^*_k = S^f_k Y_k^T \left( Y_k Y_k^T + R_k \right)^{-1} $$
      - Computation of the analysis
        $$ x^a_k = x^f_k + K^*_k \left( y_k - H_k(x^f_k) \right) $$
      - Computing the (square root) matrix of the modes
        $$ S^a_k = S^f_k \left( I - Y_k^T (Y_k Y_k^T + R_k)^{-1} Y_k \right)^{1/2} $$
        - $S^a_k$ has $m$ modes.
   (b) Reduction
      - Diagonalisation $V A V^T = S^a_k^T S^a_k$
      - Reduction of the ensemble $\tilde{S}^a_k = S^a_k \tilde{V}$
      - $\tilde{S}^a_k$ has $m - q$ modes.
   (c) Forecast
      - Computation of the forecast estimate $x^f_{k+1} = M_{k+1}(x^a_k)$
      - Computing the matrix of the modes
        $$ S^{f}_{k+1} = [M_{k+1} S^{a}_k, T_k] $$
        - $S^{f}_{k+1}$ has $m$ modes.
5.2 The stochastic ensemble Kalman filter

5.2.1 The analysis step

The EnKF seeks to mimic the analysis step of the Kalman filter but with an ensemble of limited size rather than with error covariance matrices. The goal is to perform for each member of the ensemble an analysis of the form

\[ x_a^i = x_f^i + K^* \left( y_i - H(x_f^i) \right). \]  

(5.8)

where \( i = 1, \ldots, N_e \) is the member index in the ensemble, \( x_f^i \) is state vector \( i \) forecasted at the analysis time. For the moment, we shall assume that \( y_i \triangleq y \), the observation vector, does not depend on \( i \). \( K^* \) should be the Kalman gain, that we would like to compute from the ensemble statistics

\[ K^* = P^f H^T \left( H P^f H^T + R \right)^{-1}, \]  

(5.9)

where \( R \) is the given observational error covariance matrix. The forecast error covariance matrix is estimated from the ensemble

\[ x_f = \frac{1}{N_e} \sum_{i=1}^{N_e} x_f^i, \quad P^f = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} \left( x_f^i - x_f \right) \left( x_f^i - x_f \right)^T. \]  

(5.10)

Thanks to Eq. (5.8), we obtain a posterior ensemble \( \{ x_a^i \}_{i=1}^{N_e} \) from which we can compute the posterior statistics. Firstly, the analysis is computed as the mean of the posterior ensemble

\[ x_a = \frac{1}{N_e} \sum_{i=1}^{N_e} x_a^i. \]  

(5.11)

Secondly, the analysis error covariance matrix is computed from the posterior ensemble

\[ P_a = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} \left( x_a^i - x_a \right) \left( x_a^i - x_a \right)^T. \]  

(5.12)

Note, however, that the computation of \( P_a \) is optional since it is not required by the sequential algorithm. If \( y_i \triangleq y \), the ensemble anomalies, \( e_a^i = x_a^i - x_a \), \( i.e. \) the deviations of the ensemble members from the mean are

\[ e_a^i = e_f^i + K^* \left( 0 - He_f^i \right) = (I - K^*H)e_f^i, \]  

(5.13)

which yields

\[ P_a = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} \left( x_a^i - x_a \right) \left( x_a^i - x_a \right)^T = (I_x - K^*H)P^f(I_x - K^*H)^T. \]  

(5.14)

However, to mimic the BLUE analysis of the Kalman filter, we should have obtained instead

\[ P_a = (I_x - K^*H)P^f(I_x - K^*H)^T + K^*R K^*^T. \]  

(5.15)

Therefore, the error is systematically underestimated since the second positive term is ignored, which is likely to lead to the divergence of the EnKF.

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A solution around this problem is to perturb the observation vector for each member,
\[ y_i = y + u_i, \]
where \( u_i \) is drawn from the Gaussian distribution \( u_i \sim N(0, R) \). Firstly, we would like to ensure that \( \sum_{i=1}^{N_e} u_i = 0 \) to avoid biases. Besides, we define the empirical error covariance matrix
\[
R_u = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} u_i u_i^\top,
\]
which should coincide with \( R \) in the asymptotic limit \( N_e \rightarrow \infty \).

The anomalies are modified accordingly,
\[
e_i^a = e_i^f + K_u^* (e_i^o - He_i^f),
\]
where the gain \( K_u^* \) is the same as the Kalman gain \( K^* \) but with \( R \) replaced by the empirical \( R_u \). This error yields the correct analysis error covariance matrix:
\[
P^a = (I_x - K_u^* H)P^f (I_x - K_u^* H)^\top + K_u^* R K_u^* = (I_x - K_u^* H)P^f.
\]

### 5.2.2 The forecast step

In the forecast step, the updated ensemble obtained in the analysis is propagated by the model to the next timestep,
\[
\text{for } i = 1, \ldots, N_e \quad x_i^f = M(x_i^a).
\]
The forecast estimate is the mean of the forecast ensemble, \textit{i.e.}
\[
\bar{x}^f = \frac{1}{N_e} \sum_{i=1}^{N_e} x_i^f,
\]
while the forecast error covariance matrix is estimated via
\[
P^f = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} (x_i^f - \bar{x}^f) (x_i^f - \bar{x}^f)^\top.
\]

It is important to observe that we have avoided the use of the tangent linear operator, or of any linearisation. This makes a significant difference with the RRSQRT filter. This difference should particularly matter in stronger nonlinear regimes.

### 5.2.3 Assets of the EnKF

Like the RRSQRT filter, the particle filter or the 3D-Var, the EnKF does not require the adjoint model. More interestingly it can help avoid the computation of the adjoint of the observation operator. Let us consider the formula of the Kalman gain Eq. (5.9) that

---

1More to the point, it is \( H(x_i^f) \) that needs to be perturbed rather than \( y \). Since both are formally equivalent, \textit{perturbing the observations} is an expression that remains very common.
The stochastic ensemble Kalman filter

1. Initialisation
   • Initial system state $x_{f0}$ and initial error covariance matrix $P_{f0}$.

2. For $t_k = 1, 2, \cdots$
   (a) Observation
      • Draw a statistically consistent observation set: for $i = 1, \ldots, N_e$:
        $$y_i = y + u_i \sum_{i=1}^{N_e} u_i = 0$$
      • Compute the empirical error covariance matrix
        $$R_u = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} u_i u_i^\top$$
   (b) Analysis
      • Compute the gain $K_u^* = P_f H^\top (HP_f H^\top + R_u)^{-1}$
      • Ensemble of analyses for $i = 1, \ldots, N_e$
        $$x_{a}^i = x_{f}^i + K_u^* (y_i - H(x_{f}^i))$$
        and their mean
        $$\bar{x}^a = \frac{1}{N_e} \sum_{i=1}^{N_e} x_{a}^i$$
      • Compute the analysis error covariance matrix
        $$P_a = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} (x_{a}^i - \bar{x}^a) (x_{a}^i - \bar{x}^a)^\top$$
   (c) Forecast
      • Compute the ensemble forecast for $i = 1, \ldots, N_e$
        $$x_{f}^i = M(x_{a}^i)$$
        and their mean
        $$\bar{x}^f = \frac{1}{N_e} \sum_{i=1}^{N_e} x_{f}^i$$
      • Compute the forecast error covariance matrix
        $$P_f = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} (x_{f}^i - \bar{x}^f) (x_{f}^i - \bar{x}^f)^\top$$
clearly makes use of the adjoint of the observation operator. Instead of computing the
matrix products $P^f H^\top$ and $HP^f H^\top$, we compute

\[
\hat{y}^f = \frac{1}{N_e} \sum_{i=1}^{N_e} H\left(x_i^f\right),
\]

\[
P^f H^\top = \frac{1}{N_e-1} \sum_{i=1}^{N_e} \left(x_i^f - \bar{x}^f\right) \left[H\left(x_i^f - \bar{x}^f\right)\right]^\top,
\]

\[
\approx \frac{1}{N_e-1} \sum_{i=1}^{N_e} \left(x_i^f - \bar{x}^f\right) \left[H\left(x_i^f - \bar{x}^f\right) - \hat{y}^f\right]^\top,
\]

\[
HP^f H^\top = \frac{1}{N_e-1} \left[H\left(x_i^f - \bar{x}^f\right)\right] \left[H\left(x_i^f - \bar{x}^f\right)\right]^\top,
\]

\[
\approx \frac{1}{N_e-1} \sum_{i=1}^{N_e} \left[H\left(x_i^f - \bar{x}^f\right) - \hat{y}^f\right] \left[H\left(x_i^f - \bar{x}^f\right) - \hat{y}^f\right]^\top.
\] (5.22)

This avoids the use of not only the adjoint model, but also the tangent linear of the
observation operator.

Except with the computation of the Kalman gain, all the operations on the ensemble
members are independent. This implies that their parallelisation can be trivially carried
out. This is one of the main reasons for the success and popularity of the EnKF.

Note that there are other variants of the EnKF. The habit of the author of these notes
is to call this variant with perturbation of the observations the **stochastic EnKF**.

### 5.2.4 Examples

Let us again consider the nonlinear oscillator of chapter 2. The performance of the stochastic
EnKF on this test case is shown in Fig. 5.1.

Let us now apply the stochastic EnKF on the somewhat more difficult perturbed anharmonic oscillator introduced in chapter 4. It shows the divergence of the extended Kalman filter in the absence of frequent observations (less than one observation every 58 timesteps). Under the same setting, we implement the EnKF with a frequency of observation every 50 timesteps, then one observation every 100 timesteps. The results are shown in Fig. 5.2 and Fig. 5.3 respectively. No divergence is observed.

Hence, the EnKF accommodates a lower observation frequency than the extended
Kalman filter.

### 5.3 The deterministic ensemble Kalman filter

The stochastic EnKF enables to individually track each member of the ensemble, with
the analysis step for them to interact. This nice property comes with a price: we need
to independently perturb the observation vector of each member. Although this seems
elegant, it also introduces numerical noise when drawing the perturbation $u_i$. This can
affect the performance of the stochastic EnKF, especially when the number of observations
for a single analysis is limited.
5.3 The deterministic ensemble Kalman filter

An alternative idea to perform a statistically consistent EnKF analysis would be to follow the square root approach of the RRSQRT filter. This yields the so-called ensemble square root Kalman filter (EnSRF). Because with this scheme the observations are not perturbed, it qualifies as a deterministic EnKF.

There are several variants of the EnSRF. Be aware that they are, to a large extent, theoretically equivalent. In this course, we shall focus on the so-called ensemble-transform variant of the deterministic EnKF, usually abbreviated ETKF (Bishop et al., 2001; Hunt et al., 2007). Rather than performing the linear algebra in state or observation space, the algebra is mostly performed in the ensemble space.

5.3.1 Algebra in the ensemble space

Let us define this ensemble space. We can write the forecast error covariance matrix as

$$P^f = \frac{1}{N_e - 1} \sum_{i=1}^{N_e} (x^f_i - \bar{x}^f) (x^f_i - \bar{x}^f)^\top = X_f X_f^\top. \tag{5.23}$$

where $X_f$ is a $N_x \times N_m$ matrix whose columns are the normalised anomalies

$$[X_f]_i = \frac{x^f_i - \bar{x}^f}{\sqrt{N_e - 1}}. \tag{5.24}$$

The matrix of the anomalies, $X_f$, plays the role of the matrix of the reduced modes $S^f$ of the RRSQRT filter.

We shall assume that the analysis belongs to the affine space generated by

$$x = \bar{x}^f + X_f w \tag{5.25}$$
Figure 5.2: Implementation of the stochastic EnKF on the perturbed anharmonic oscillator. The system is observed once every 50 integration timesteps. This graph can be compared to that of Fig. 4.1.

where $w$ is a vector of coefficients in the vector space $\mathbb{R}^{N_e}$. We call it the ensemble space. Note that the decomposition of $x$ is not unique, because the vector of coefficients $w + \lambda 1$ with $1 = (1, \ldots, 1)^T \in \mathbb{R}^{N_e}$ yields the same state vector $x$ since $X_f 1 = 0$.

Again we shall use the notation $Y_f$ to represent $HX_f$ if the observation operator is linear. If it is nonlinear, we consider $Y_f$ to be the matrix of the observation anomalies

$$[Y_f]_i = \frac{H(x^f_i) - \bar{y}^f}{\sqrt{N_e - 1}} \quad \text{with} \quad \bar{y}^f = \frac{1}{N_e} \sum_{i=1}^{N_e} H(x^f_i),$$

which stands as a fine approximation following the ideas of section 5.2.3.

### 5.3.2 Analysis in ensemble space

As opposed to the stochastic EnKF, we wish to perform a single analysis rather than performing an analysis for each member of the ensemble. For the analysis estimate, we can adopt the mean analysis of the stochastic EnKF, i.e.

$$\bar{x}^a = \bar{x}^f + K^* \left( y - H(\bar{x}^f) \right)$$

where we used $K^* = P^f H^T (H P^f H^T + R)^{-1}$ rather than $K_u$, as defined by Eq. (5.17), since the observations are not perturbed in a deterministic approach.

We shall reformulate this analysis but working in the ensemble space. We look for the optimal coefficient vector $w^a$ that stands for $x^a \equiv \bar{x}^a$ defined above. To do so we write

$$x^a = \bar{x}^f + X_f w^a.$$
5.3 The deterministic ensemble Kalman filter

Figure 5.3: Implementation of the stochastic EnKF on the perturbed anharmonic oscillator. The system is observed once every 100 integration timesteps. The filter remains stable and manages to track the truth. This graph can be compared to that of Fig. 4.2.

Inserting this decomposition into Eq. (5.27), and denoting $\delta = y - H(\overline{x}^f)$, we obtain

$$\overline{x}^f + X_f w^a = \overline{x}^f + X_f X_f^T H^T \left( H X_f X_f^T H^T + R \right)^{-1} \delta,$$

which suggests

$$w^a = X_f^T H^T \left( H X_f X_f^T H^T + R \right)^{-1} \delta = Y_f^T \left( Y_f Y_f^T + R \right)^{-1} \delta.$$  \hspace{1cm}(5.30)

The gain $K^*$ is computed in the observation space. Remember that using the Sherman-Morrison-Woodbury lemma, it was possible to compute the gain in state space. We showed in chapter 1 that $P^b H^T (HP^b H^T + R)^{-1} = (P^{-1} + H^T R^{-1} H)^{-1} H^T R^{-1}$. Now, with the identifications $P^b = I_e$ and $H = Y_f$, we obtain

$$w^a = \left( I_e + Y_f^T R^{-1} Y_f \right)^{-1} Y_f^T R^{-1} \delta.$$  \hspace{1cm}(5.31)

The gain matrix is now computed in the ensemble space rather than the observation space or the state space! This makes a difference as far numerical complexity is concerned.

5.3.3 Generating the posterior ensemble

Writing the analysis in ensemble space is merely a reformulation of the stochastic EnKF, hardly more. The genuine difference between the deterministic EnKF and the stochastic EnKF comes in the generation of the posterior ensemble.

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To generate a posterior ensemble that would be representative of the posterior uncertainty, we would like to factorise \( P^a = X_aX_a^\top \). We can repeat the derivation of the RRSQRT filter, Eq. (5.6),

\[
P^a = (I_x - K^kH)P^f \\
\simeq \left( I_x - X_fY_f^\top (Y_fY_f^\top + R)^{-1}H \right)X_fX_f^\top \\
\simeq X_f \left( I_e - Y_f^\top (Y_fY_f^\top + R)^{-1}Y_f \right)X_f^\top.
\]

That is why we choose

\[
X_a = X_f \left( I_e - Y_f^\top (Y_fY_f^\top + R)^{-1}Y_f \right)^{1/2}.
\]

Remarkably, this expression can be simplified into

\[
X_a = X_f \left( I_e - Y_f^\top (Y_fY_f^\top + R)^{-1}Y_f \right)^{1/2} \\
= X_f \left( I_e - (I_e + Y_f^\top R^{-1}Y_f)^{-1}Y_f^\top R^{-1}Y_f \right)^{1/2} \\
= X_f \left[ (I_e + Y_f^\top R^{-1}Y_f)^{-1} \left( I_e + Y_f^\top R^{-1}Y_f - Y_f^\top R^{-1}Y_f \right) \right]^{1/2} \\
= X_f \left( I_e + Y_f^\top R^{-1}Y_f \right)^{-1/2},
\]

where we have used the Sherman-Morrison-Woodbury lemma again between the first line and the second line.

This posterior ensemble of anomalies is all that we need to cycle the deterministic EnKF. Defining \( T = (I_e + Y_f^\top R^{-1}Y_f)^{-1/2} \), we can build the posterior ensemble as

\[
\text{for } i = 1, \ldots, N_e \quad x^a_i = \bar{x}^a + \sqrt{N_e^{-1}I} [T]_i = \bar{x}^f + X_f \left( w^a + \sqrt{N_e^{-1}I} [T]_i \right).
\]

Note that the updated ensemble is centred on \( \bar{x}^a \), since

\[
\frac{1}{N_e} \sum_{i=1}^{N_e} x^a_i = \bar{x}^a + \sqrt{N_e^{-1}I}X_f1 = \bar{x}^a + \sqrt{N_e^{-1}I}X_f1 = \bar{x}^a,
\]

where \( 1 = [1, \ldots, 1]^\top \). Not all EnSRF have a centred ensemble. It has been shown that a centred posterior ensemble often (if not always) yields better performance.

### 5.4 Localisation and inflation

The EnKF approach seems magical. We traded the extended Kalman filter for a considerably computationally cheaper filter meant to achieve similar performances if not better. But there is no free lunch and this comes with significant drawbacks to cope with. Fundamentally, one cannot hope to represent the full error covariance matrix of a complex...
The deterministic ensemble Kalman filter
(Ensemble transform Kalman variant)

1. Initialisation
   • Ensemble of state vectors \( \mathbf{E}_0 = \{ x_0, \ldots, x_{N_e} \} \).

2. For \( t_k = 1, 2, \ldots \)
   (a) Analysis
      • Compute forecast mean, the ensemble anomalies and the observation anomalies:
        \[
        \bar{\mathbf{x}}_k^f = \frac{1}{N_e} \mathbf{E}_k^f \\
        \mathbf{X}_k^f = \frac{\mathbf{E}_k^f - \bar{\mathbf{x}}_k^f 1^\top}{\sqrt{N_e - 1}} \\
        \mathbf{Y}_k^f = \frac{(H_k(\mathbf{E}_k^f) - H_k(\bar{\mathbf{x}}_k^f) 1^\top)}{\sqrt{N_e - 1}}
        \]
      • Computation of ensemble transform matrix
        \[
        \mathbf{\Omega} = \left( \mathbf{I}_e + \mathbf{Y}_k^f 1^\top \mathbf{R}_k^{-1} \mathbf{Y}_k ight)^{-1}
        \]
      • Analysis estimate in ensemble space
        \[
        \mathbf{w}_k^a = \mathbf{\Omega} (\mathbf{Y}_k^f)^\top \mathbf{R}_k^{-1} (\mathbf{y}_k - H_k(\bar{\mathbf{x}}_k^f))
        \]
      • Generating the posterior ensemble
        \[
        \mathbf{E}_k^a = \bar{\mathbf{x}}_k^f 1^\top + \mathbf{X}_k^f \left( \mathbf{w}_k^a 1^\top + \sqrt{N_e - 1} \mathbf{\Omega}^{1/2} \right)
        \]
    (b) Forecast
      • Forecast ensemble
        \[
        \mathbf{E}_{k+1}^f = M_{k+1} (\mathbf{E}_k^a)
        \]
high-dimensional system with a few modes \( N_e \ll N_x \), usually from a few dozen to a few hundred. This implies large sampling errors, meaning that the error covariance matrix is only sampled by a limited number of modes. Even though the unstable degrees of freedom of dynamical systems that we wish to control with a filter are usually far fewer than the dimension of the system, they often still represent thousands, possibly millions of independent degrees of freedom. Nonetheless, forecasting an ensemble of such size is usually not affordable.

The consequence of this issue always is the divergence of the filter which often makes the EnKF a useless tool without the implementation of efficient fixes. In other words, in order to make the EnKF a viable algorithm, one needs to cope with the sampling errors. Fortunately, there are clever tricks to overcome this major issue. They are known as the localisation and the inflation.

5.4.1 Localisation

For most geophysical systems distant observables are barely correlated. In other words, two distant parts of the system are almost independent at least for short time scales. We could exploit this and spatially localise the analysis, i.e. restrict its spatial scope. There are two main ways to achieve this goal.

The first one is called domain localisation, or local analysis. Instead of performing a global analysis valid at any location in the domain, we perform a local analysis to update the local state variables using local observations. Typically one would update a state variable at a location \( M \) by assimilating only the observations within a fixed range of \( M \). If this range, known as the localisation length, is too long, then the analysis becomes global as before, and may fail. On the other hand, if the localisation length is too small, some short to medium-range correlation might be neglected resulting in a possibly viable but a poor data assimilation system. One would then update all the state variables performing those local analyses. That may seem a formidable computational task but all these analyses can be carried out in parallel. Besides, since the number of local observations is a fraction of the total, the computation of the local gain is much faster than that of the global gain. All in all, such an approach is viable even on very large systems. Figure 5.4 is a schematic representation of a local update around \( M \) in the domain localisation approach where only the surrounding observations of a grid-cell to update are assimilated. One significant drawback of domain localisation is its inability to process non-local observations such as space-borne radiance measurements, since they involve state vector components across several local domains. One has to resort to drastic approximation to assimilate them.

The second approach called covariance localisation focuses on the forecast error covariance matrix. It is based on the remark that the forecast error covariance matrix \( P^f \) is of low rank, at most \( N_e - 1 \). If the empirical \( P^f \) is probably a good approximation of the true error covariance matrix at short distances, the insufficient rank will induce long-range correlations. These correlations are spurious: they are likely to be unphysical and are not present in the true forecast error covariance matrix. A brilliant idea is to regularise \( P^f \) by smoothing out these long-range spurious correlations and increasing the rank of \( P^f \). To do so, one can regularise \( P^f \) by multiplying it point-wise with a short-range predefined correlation matrix \( \rho \in \mathbb{R}^{N_x} \). The point-wise multiplication is called a Schur product and
denoted with a ◦. It is defined by

$$\left[ P^f \circ \rho \right]_{ij} = \left[ P^f \right]_{ij} \left[ \rho \right]_{ij}. \quad (5.37)$$

With reasonable conditions on $\rho$, one can mathematically ensure that $P^f \circ \rho$ is full-rank definite-positive. As can be read from Eq. (5.37), the spurious correlations have been levelled off by $\rho$, if $\rho$ is short-ranged. This regularised $P^f \circ \rho$ will be used in the EnKF analysis as well as in the generation of the posterior ensemble.

Of course, as a side-effect, it might also remove some physical and true long-distance correlations, thus inducing some physical imbalance in the analysis.

Covariance localisation and domain localisation seem quite different. Indeed they are mathematically distinct, though they are both based on two sides of the same paradigm. However, it can be shown that they should yield very similar results if the innovation at each time step is not too strong (Sakov and Bertino, 2011).

Figure 5.5 illustrates covariance localisation. We consider a one-dimensional state space of $N_x = 200$ variables, and an error covariance matrix defined by

$$[P]_{n,m} = e^{-\frac{|n-m|}{L}}, \quad (5.38)$$

where $n, m = 1, \ldots, N_x$ and $L = 10$. An ensemble of $N_e = 20$ members is generated from the exact Gaussian distribution of error covariance matrix $P$. The empirical error covariance matrix is computed from this ensemble, and a Schur localisation is applied to the empirical covariance matrix to form a regularised covariance matrix. For the sake of the demonstration, we take the localisation matrix to be the correlation matrix of the true covariance matrix.
5.4.2 Inflation

Even when the analysis is made local, the error covariance matrices are still evaluated with a finite-size ensemble. This often leads to sampling errors. With a proper localisation scheme, they might be small. However small they may be, they will accumulate and they will carry over to the next cycles of the sequential EnKF scheme. As a consequence, there is always a risk that the filter may ultimately diverge. One way around is to inflate the error covariance matrix by an empirical factor $\lambda^2$ slightly greater than one. It can be applied before or after the analysis. For instance, after the analysis, inflating means

$$P^a \rightarrow \lambda^2 P^a. \quad (5.39)$$

Another way to achieve this is to inflate the ensemble

$$x^a_i \rightarrow \bar{x}^a + \lambda (x^a_i - \bar{x}^a). \quad (5.40)$$

This type of inflation is called \textit{multiplicative inflation}.

In a perfect model context, the multiplicative inflation is meant to compensate for sampling errors and the approximate assumption of Gaussianity (Bocquet, 2011). In that case, it helps cure an intrinsic source of error of the EnKF scheme.

Yet, inflation can also compensate for the underestimation of the errors due to the presence of model error, a source of error external to the EnKF scheme. In this context, additive type of inflation can also be used, either by

$$P^f \rightarrow P^f + Q, \quad (5.41)$$

or by adding noise to the ensemble members

$$x^f_i \rightarrow x^f_i + \epsilon_i \quad \text{with} \quad E[\epsilon_i \epsilon_i^T] = Q. \quad (5.42)$$

5.5 Numerical tests on the Lorenz-96 model

The Lorenz-96 low-order model is a one-dimensional toy-model introduced by the famous meteorologist Edward Lorenz in 1996 (Lorenz and Emmanuel, 1998). It represents a mid-latitude zonal circle of the global atmosphere. It has $N_x = 40$ variables $\{x_n\}_{n=1,\ldots,N_x}$. Its dynamics are given by the following set of ordinary differential equations,

$$\frac{dx_n}{dt} = (x_{n+1} - x_{n-2})x_{n-1} - x_n + F, \quad (5.43)$$
5.5 Numerical tests on the Lorenz-96 model

for \( n = 1, \ldots, N_x \). The domain on which is defined the 40 variables is circle-like, so that \( x_{-1} \equiv x_{39}, x_0 \equiv x_{40}, \) and \( x_{41} \equiv x_1 \). Hence it is assumed periodic. \( F \) is chosen to be 8. The term \( (x_{n+1} - x_{n-2})x_{n-1} \) is a nonlinear convective term. It preserves the energy \( \sum_{n=1}^{N_x} x_n^2 \). \(-x_n\) is a dampening term while \( F \) is an energy injection/extraction term (depending on the sign of the variables). This makes the dynamics of this model chaotic. The model has 13 positive Lyapunov exponents and one is equal to zero. This implies that, out of the 40 degrees of freedom of this model, 14 (asymptotic) directions correspond to growing or neutral modes. In practice and for short-term forecast, around 13 directions (that change with time) in the 40 model state space make a small perturbation grows under the action of the model. The model is usually integrated using the 4th-order Runge-Kutta scheme using an integration time-step of \( \delta t = 0.05 \).

We shall apply the deterministic EnKF to this model. The time interval between observational update will be \( \Delta t = 0.05 \), meant to be representative of a data assimilation cycle of global meteorological models (about 6 hours). With such value for \( \Delta t \), the data assimilation system is considered to be weakly nonlinear, leading to statistics of the errors weakly diverging from Gaussianity.

Figure 5.6 displays a trajectory of the model state. The model is characterised by about eight nonlinear waves that interact. As can be seen, it seems difficult to predict the behaviour of these waves except that they have the tendency to drift westward. These waves are meant to represent large-scale Rossby waves.

![Figure 5.6: Trajectory of a state of the Lorenz-96 model. The Ox coordinate represents time in units of \( \Delta t = 0.05 \). The Oy coordinate represents the 40 state variables.](image)

A twin experiment is conducted. The truth is represented by a free model run, meant to be tracked by the data assimilation system. The system is assumed to be fully observed \( (N_y = 40) \) every \( \Delta t \), so that \( H_k = I_{40} \), with the observation error covariance matrix \( R_k = I_{40} \). The related synthetic observations are generated from the truth, and perturbed according to the same observation error prior. The performance of a scheme is measured by the temporal mean of a root mean square difference between a state estimate \( (x^a) \) and the truth \( (x^t) \). Typically, one averages over time the analysis root mean square error

\[
\sqrt{\frac{1}{N_x} \sum_{n=1}^{N_x} (x^a_n - x^t_n)^2}.
\] (5.44)

All data assimilation runs will extend over \( 10^5 \) cycles, after a burn-in period of \( 5 \times 10^3 \) cycles, which guarantees satisfying convergence of the error statistics (due to ergodicity of the dynamics).

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We vary the ensemble size from $N_e = 5$ to $N_e = 50$ and compare the performance in terms of the root mean square error of

- the deterministic EnKF without inflation and without localisation,
- the deterministic EnKF with optimally-tuned inflation and without localisation,
- the deterministic EnKF without inflation and with optimally-tuned localisation,
- and the deterministic EnKF with optimally-tuned inflation and with optimally-tuned localisation.

The average root mean square errors of the analyses over the long run are displayed in Fig. 5.7.

By optimally-tuned, it is meant that the parameters in the localisation and inflation are chosen in order to minimise the root mean square error. If one agrees that the application of the EnKF to this low-order model captures several of the difficulties of realistic data assimilation, it becomes clear from these numerical results that localisation and inflation are indeed mandatory ingredients of a satisfying implementation of the EnKF.

Figure 5.7: Average analysis root mean square error for a deterministic EnKF (ETKF) without localisation and without inflation (□); without localisation and with optimally-tuned inflation (○); with optimally-tuned localisation and no inflation (◊); with optimally-tuned localisation and with optimally-tuned inflation (△).
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