# AtmoData Scientific documentation

Version 1 for AtmoData 1.1

CEREA – ENPC / EDF R&D Hervé Njomgang, Vivien Mallet and Luc Musson-Genon

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## Introduction

AtmoData ([6]) is a library for data processing and parameterizations in atmospheric chemistry and physics. This document describes the parameterizations and the miscellaneous physical functions available in the library.

Apart from this scientific documentation, additional resources are available on AtmoData website at http://www.enpc.fr/cerea/atmodata/: a more complete introduction, the lastest version this scientific documentation, a user's guide and the reference documentation.

In case you want to refer to AtmoData in a publication, a conference, a poster or so, please include the reference [6]: Vivien Mallet and Bruno Sportisse, *Data processing and parameterizations in atmospheric chemistry and physics: the AtmoData library*, Submitted to Atmospheric Environment, 2005.

## Notations

Variables							
Attenuation							
Cloud fraction							
Clouds basis height							
Critical relative humidity							
Molecular diffusivity of the species $s$							
Emission for the species $s$							
Emission factor for the species $s$							
Emission rate for the species $s$							
Heat flux							
Momentum flux							
Reactivity of species $s$							
Relative humidity							
Henry constant for species $s$							
Photolysis rate							
Vertical diffusion							
Mixing length							
Monin-Obukhov length							
Land use coverage							
Liquid water content							
Liquid water path							
Cloudiness							
Pressure							
Photosynthetically active radiation							
Planetary boundary layer							
Planetary boundary layer height							
Partial pressure for species $\#i$							
Saturation pressure for species $\#i$							
Specific humidity							
Surface sensible heat flux							
Richardson number							
Critical Richardson number							
Solar radiation							
Temperature							
Cloud transmissivity							
Virtual temperature							
Meridional wind							

- $\mathcal{W}$  Wind module
- U Zonal wind
- $u_*$  Surface friction velocity
- $V_d$  Deposition velocity
- $w_*$  Convective velocity scale
- Z Zenith angle
- $Z_r$  Roughness height
- $\theta$  Potential temperature
- $\lambda \qquad \qquad {\rm Asymptotic\ mixing\ length}$
- $\rho$  Mass density
- au Cloud optical depth

## Coordinates

$F_i$	Value of field F at node $i$ along longitude
$F_{i_{interf}}$	Value of field F at interface $i$ along longitude
$F_j$	Value of field F at node $j$ along latitude
$F_{j_{interf}}$	Value of field F at interface $j$ along latitude
$F_k$	Value of field F at node $k$ along altitude
$F_{k_{interf}}$	Value of field F at interface $k$ along altitude
z ,	Altitude
$\sigma$	Ratio of pressure on surface pressure

Note:  $z_k$  is dedicated to the altitudes of *nodes*; so  $z_k \neq 0$  and  $z_{0_{interf}} = 0$ .

## Constants

$c_p$	Specific heat capacity	$\frac{7}{2}r$
	at constant pressure	
$c_p$	Specific heat capacity	$\frac{5}{2}r$
	at constant volume	
g	Standard gravity	$9.80665 \text{ m} \cdot \text{s}^{-1}$
$M_{air}$	Molar mass of air	$28.96 \text{ g} \cdot \text{mol}^{-1}$
$M_{vapor}$	Molar mass of vapor	$18 \mathrm{g} \cdot \mathrm{mol}^{-1}$
$P_0$	Standard pressure	101,325 Pa
r	Molar gas constant	$287.0 \text{ J} \cdot \text{kg}^{-1} \cdot \text{K}^{-1}$
	for dry air	
$r_{vap}$	Molar gas constant	$461.5 \mathrm{J\cdot kg^{-1}\cdot K^{-1}}$
	for vapor	
R	Gas constant	$8.314~\mathrm{J}\cdot\mathrm{K}^{-1}\cdot\mathrm{mol}^{-1}$
$T_0$	Standard temperature	273.15 K
$\kappa$	Von Kármán constant	0.4

## 1 Meteorology

## 1.1 Conversion of coordinates

#### 1.1.1 ECMWF

In the ECMWF model, the vertical coordinates are hybrid  $\sigma$ -P levels. The pressure at level k is computed as follows:

$$P_k = \alpha_k P_{surf} + \beta_k P_{ref} \tag{1}$$

where  $P_{ref}$  is a reference pressure, and  $\alpha_k$  and  $\beta_k$  are hybrid coefficients.

To convert the pressure into meters, one uses the following formula:

$$z_{n+1} = z_n + \frac{rT}{P} \ln \frac{P_n}{P_{n+1}}$$

$$\tag{2}$$

where index n denotes the level and  $z_0 = 0$ m. This formula can be easily derived from the hydrostatic approximation and the perfect gas law, assuming that the atmosphere is *isotherm* within two layers.

> ComputePressure, ComputeHeight, ComputeInterfHeight, ComputeMiddleHeight

### 1.1.2 MM5

In MM5, the pressure at level k is given by

$$P_k = \sigma_k P_{surf} \tag{3}$$

Then formula (2) can also be used to convert the pressure to altitudes.

ComputeHeight, ComputeInterfHeight, ComputeMiddleHeight

For horizontal coordinates, MM5 uses three projections:

- Lambert Azimuthal Equal Area,
- Lambert Conformal Conic, and
- Mercator.

AtmoData provides classes to perform coordinate transformations from these projections to latitude/longitude coordinates, and vice versa. Note that AtmoData does not manipulate directly the three MM5 coordinates but it handles MM5 indices instead, which is more convenient.

> LonlatToMM5LccInd, LonlatToMM5MercInd, LonlatToMM5StereInd, MM5LccIndToLonlat, MM5MercIndToLonlat, MM5StereIndToLonlat

#### 1.2 Basic meteorology

#### **1.2.1** Potential temperature

The potential temperature is computed with:

$$\theta = T \left(\frac{P}{P_{\text{ref}}}\right)^{-\frac{r}{c_p}} \tag{4}$$

Another formula (equivalent):

$$\theta = T \left(\frac{P}{P_{\rm ref}}\right)^{\frac{1-\gamma}{\gamma}} \tag{5}$$

where  $\gamma = \frac{c_p}{c_v}$ ,  $c_p = c_v + r$  for a perfect gas and, usually,  $P_{\text{ref}} = 100,000$ Pa.

## ${\tt ComputePotentialTemperature}$

## 1.2.2 Virtual temperature

$$T_v = T\left(1 - q + \frac{r_{vap}}{r}q\right)$$
(6)

where q is the specific humidity (cf. 1.2.3). So

$$T_v = T \left( 1 + 0.608q \right) \tag{7}$$

## 1.2.3 Specific humidity

Definition:

$$q = \frac{\rho_{vapor}}{\rho_{vapor} + \rho_{air}}$$
(8)

It is not computed by AtmoData, but computed by meteorological models.

#### 1.2.4 Relative humidity

First are computed the partial pressure of vapor  $P_{vapor}$  and the saturation pressure of vapor  $P_{vapor}^{sat}$ :

$$P_{vapor} = \frac{q}{q + \frac{M_{vapor}}{M_{air}}(1-q)}P$$
(9)

$$P_{vapor}^{sat} = 611.2 \exp\left(\frac{17.67(T - 273.15)}{T - 29.65}\right)$$
(10)

from which is derived relative humidity:

$$h = \frac{P_{vapor}}{P_{vapor}^{sat}} \tag{11}$$



## **1.2.5** Winds

Note: this section deals with a tranformation specific to the chemistry-transport model Polair $3D^{\dagger 1}$ .

The zonal and meridional winds U and V from the meteorological models are transformed so as to be input of the chemistry-transport model Polair3D:

$$\begin{cases} U_{Polair3D} = \frac{U}{\cos\varphi} \\ V_{Polair3D} = V\cos\varphi \end{cases}$$
(12)

where  $\varphi$  represents the latitude.

## TransformMeridionalWind, TransformZonalWind

## 1.2.6 Module of Winds

The wind module is, of course, the Euclidean norm of the vector (U, V):

$$\mathcal{W} = \sqrt{U^2 + V^2} \tag{13}$$

## 1.2.7 Richardson number

The Richardson number Ri is computed as follows. We set:

$$D\mathcal{W}_k = \sqrt{\left(\frac{\Delta U_k}{\Delta z_k}\right)^2 + \left(\frac{\Delta V_k}{\Delta z_k}\right)^2}$$

<sup>&</sup>lt;sup>†1</sup>http://www.enpc.fr/cerea/polair3d/

with  $\Delta X_k = X_k - X_{k-1}$ , and then

$$Ri_{k} = \frac{g \ \Delta \theta_{k}}{\max(D \mathcal{W}_{k}^{2}, WT^{2}) \ \theta_{k-1} \ \Delta z_{k}}$$
(14)

where WT is the wind threshold, set to 0.001 by default.

ComputeRichardson

## 1.2.8 Planetary boundary layer height

The planetary boundary layer height, PBLH, is usually retrieved from the outputs of a meteorological model. Otherwise, two diagnoses are proposed. One returns the height at which the Richardson number reaches a critical value (usually set to 0.21). The other diagnosis follows [12]. Details are provided in section 2.2.2.

ComputeBoundaryLayerHeight\_Richardson, ComputeBoundaryLayerHeight\_TM

### 1.3 Clouds

#### 1.3.1 Critical relative humidity

The critical relative humidity is the threshold value above which there is formation of clouds.

There are several parameterizations in AtmoData for CRH, taking in account the dependence upon the altitude and the special case of the planetary boundary layer.

### • Parameterization with $\sigma$

This is a well known parameterization, widely used (e.g. in CMAQ, [1]).

It approximates CRH with:

$$CRH_k = 1 - \lambda_0 \sigma_k (1 - \sigma_k) (1 + \lambda_1 (\sigma_k - 0.5))$$
(15)

where  $\sigma_k$  is the ratio of pressure at level k on the surface pressure, and the  $\lambda_i$  are empirical coefficients. Their default values are  $\lambda_0 = 2$  and  $\lambda_1 = \sqrt{3}$ .

A more general parameterization is derived from the previous one:

$$CRH_k = 1 - \lambda_0 \sigma_k^{a_0} (1 - \sigma_k)^{a_1} (1 + \lambda_1 (\sigma_k - 0.5))$$
(16)

Default values:  $\lambda_0 = 1.1$ ,  $\lambda_1 = \sqrt{1.3}$ ,  $a_0 = 0$ , and  $a_1 = 1.1$ .

#### • In the planetary boundary layer

Within the planetary boundary layer, one may set the critical relative humidity to a constant (e.g., [1, pages 48-49]):

$$z_k < PBLH \implies CRH_k = CRH_{PBL}$$
 (17)

where PBLH is the PBL height, and  $|CRH_{PBL} = 0.98|$  by default.

• Pressure parameterization

$$CRH_{k} = \begin{cases} CRH_{0} &, P_{ref,0} < P_{k} \\ CRH_{1} &, P_{ref,1} < P_{k} < P_{ref,0} \\ CRH_{2} &, P_{k} < P_{ref,1} \end{cases}$$
(18)

By default,  $P_{ref,0} = 70,000$  Pa,  $P_{ref,1} = 40,000$  Pa,  $CRH_0 = 0.75$ ,  $CRH_1 = 0.95$ ,  $CRH_2 = 0.95$ .

ComputeCriticalRelativeHumidity, ComputeCriticalRelativeHumidity\_extended

#### 1.3.2 Cloud Fraction

The cloud fraction CF is computed using CRH, by one of the following formulae (one may refer to [1, pages 48-49]):

$$CF_{k} = \begin{cases} 0.34 \left(\frac{h_{k} - CRH}{1 - CRH}\right) &, z_{k} < PBLH \\ \left(\frac{h_{k} - CRH}{1 - CRH}\right)^{2} &, z_{k} > PBLH \end{cases}$$
(19)

In an alternative and simplified parameterization, one uses the second formula for all vertical levels.

#### 1.3.3 Cloudiness

The parameterization is similar to that of CMAQ model ([1]). One considers three layers and thus low, medium and high cloudinesses. The layer limits are determined by the two given pressures  $P_{lim,0}$  and  $P_{lim,1}$ . In each layer, the clouds are contained within the levels  $k_{base}$  (included) and  $k_{top}$  (excluded), with  $k_{base} < k_{top}$ .  $k_{base}$  is the first layer (starting from the ground) where CF reaches 50% of its maximum value along the vertical column and

in the low, medium or high cloud-layer, and  $k_{top}$  is the first above  $k_{base}$  where CF is less than this threshold. Finally, the low, medium and high cloudinesses are computed with:

$$\begin{split} & \mathcal{N} = \frac{\sum_{k_{base}}^{k_{top}-1} CF_k \Delta z_k}{\sum_{k_{base}}^{k_{top}-1} \Delta z_k} \end{split} \tag{20}$$

#### 1.3.4 Clouds basis height

The height CH of the basis of clouds is

- 1. at the first level k where the relative humidity exceeds the critical relative humidity, or
- 2. given by the base of the clouds as it appears when the cloudiness is computed ( $k_{base}$  in section 1.3.3).

ComputeCloudHeight

## 1.4 Attenuation

The attenuation  $\mathcal{A}$  measures the decrease in the photolysis rates of chemical reactions when solar radiation is partially absorbed by clouds. The photolysis rates should be multiplied by  $\mathcal{A}$ .

## 1.4.1 "RADM" parameterization

AtmoData implements the parameterization from [2, 5]. The photolysis rates J become:

$$J_{below} = \mathcal{A}_{below} J_{clean}$$

below the clouds, and

$$J_{above} = \mathcal{A}_{above} J_{clear}$$

above the clouds, to take in account the enhancement in photolysis rates due to the reflection of radiations. The attenuation coefficients are computed with:

$$\mathcal{A}_{below} = 1 - cfrac(1 - 1.6Tr\cos Z)$$
(21)

$$\mathcal{A}_{above} = 1 + cfrac(1 + (1 - Tr)\cos Z)$$
(22)

where cfrac is the cloud coverage fraction (computed, with the sum of medium and high cloudinesses, as  $\min(1, \mathcal{N}_{medium} + \mathcal{N}_{high}))$ , Z is the zenith angle, and Tr is the cloud transmissivity, calculated with:

$$Tr = \frac{5 - e^{-\tau}}{4 + 3(1 - f)\tau}$$
(23)

where f represents the scattering phase function asymmetry factor, set to 0.86, and  $\tau$  the cloud optical depth, following an empirical formula from [11]:

$$\log(\tau) = 0.2633 + 1.7095 \ln(\log LWP) \tag{24}$$

where LWP is the liquid water path, which is a function of the liquid water content:  $LWP = LWC\Delta z$ , and  $\Delta z$  is the cloud thickness.

The attenuation coefficients are computed from the top to the bottom. Until  $\tau_k < 0.5$ ,  $Tr_k$  is set to 1 and  $\mathcal{A}_k$  is computed with the equation (21). Below the clouds and if  $\tau_k > 0.5$ , equations (23) and (22) are used. Inside clouds, the attenuation coefficients are linearly interpolated..

#### $ComputeAttenuation_LWC$

#### 1.4.2 ESQUIF parameterization

This parameterization was developped during the ESQUIF campaign ([3]) with measurements of the photolysis rates for  $NO_2$ :

$$\mathcal{A} = (1 - a \ \mathcal{N}_{high})(1 - b \ \mathcal{N}_{medium})e^{-cB}$$
(25)

with:

$$B = \frac{\int_{z=0}^{z<1500m} \min(0., h(z) - 0.7)}{\int_{z=0}^{z<1500m} 1 - 0.7}$$

and h is piecewise linear.

Note that the attenuation coefficients do not depend on the altitude.

## 2 Vertical diffusion

The vertical diffusion coefficient  $K_z$  is computed in AtmoData using two parameterizations respectively introduced in [4] and [12] (the latter is especially designed for models with a low vertical resolution). Notice that  $K_z$  is provided at the interfaces of the vertical grid.

#### 2.1 Louis' parameterization

For each vertical cells, the stability function F is first calculated:

$$F(Ri_k) = \begin{cases} \left(1 + 3BRi_k\sqrt{1 + DRi_k}\right)^{-1} & , \quad Ri_k \ge 0\\ 1 - 3BRi_k\left(1 + \frac{3BC \ l_{k_{interf}}^2}{(z_{k_{interf}} + z_0)^2}\sqrt{\frac{|Ri_k|}{27}}\right)^{-1} & , \quad Ri_k < 0 \end{cases}$$
(26)

where Ri is the Richardson number (refer to section 1.2.7), and  $l_{k_{interf}}$  is the mixing length at level  $k_{interf}$ :

$$l_{k_{interf}} = \frac{\kappa z_{k_{interf}}}{1 + \frac{\kappa z_{k_{interf}}}{\lambda}}$$
(27)

 $\kappa$  is the Von Kármán constant set to 0.40,  $\lambda$  is the asymptotic mixing length, set to 100m. The default values of the other parameters are:

$$B = C = D = 5$$
 and  $z_0 = 1$ 

In addition, a threshold value  $Ri_t$  ( $Ri_k > Ri_t \implies Ri_k = Ri_t$ ) for the Richardson number is computed following [8, 9]:

$$Ri_t = a \left(\frac{\Delta z_k}{\Delta z_0}\right)^b \tag{28}$$

with a = 0.115, b = 0.175 and  $\Delta z_0 = 0.01$ , by default.

Finally

$$K_{z,k_{interf}} = l_{k_{interf}}{}^{2}F(Ri_{k}) \left\| \frac{\partial(U_{k},V_{k})}{\partial z} \right\|_{2}$$
(29)

where  $\frac{\partial(U_k, V_k)}{\partial z}$  is approximated with first-order finite differences.

ComputeLouisKz

## 2.2 Troen's & Mahrt's parameterization

## **2.2.1** $K_z$

 $K_z$  is computed with the following formula

$$K_{z,k_{interf}} = u_* \kappa z_{k_{interf}} \Phi_{m,k_{interf}}^{-1} \left( 1 - \frac{z_{k_{interf}}}{PBLH} \right)^p$$
(30)

 $u_*$  is the surface friction velocity and  $\Phi_{mk}$  the nondimensional shear. In *stable* conditions, it is computed as

$$\Phi_{m,k_{interf}} = 1 + 4.7 \frac{z_{k_{interf}}}{LMO} \tag{31}$$

LMO being the Monin-Obukhov length. In unstable conditions:

• in the surface layer  $(z < \varepsilon PBLH$ , with  $\varepsilon = 0.1)$ ,

$$\Phi_{m,k_{interf}} = \left(1 - 7\frac{z_{k_{interf}}}{LMO}\right)^{-1/3} \tag{32}$$

• above the surface layer,  $\Phi_m$  is constant

$$\Phi_m = \frac{u_*}{w_s} \tag{33}$$

 $w_s$  is the velocity scale given by

$$w_s = \left(u_*{}^3 + 7\varepsilon\kappa w_*{}^3\right)^{1/3} \tag{34}$$

where  $w_*$  is the convective velocity scale:

$$w_* = \left(\frac{gQ_0}{T_{2m}}h\right)^{1/3}$$
(35)

 $T_{2m}$  being the temperature at 2 m above the ground and  $Q_0$  the surface sensible heat flux.

The exponent p is set to 2, but its value can be modified to adjust the vertical profile of  $K_z$  (see the discussion in [12]).

#### **2.2.2** *PBLH*

Troen & Mahrt propose to compute PBLH so that it satisfies:

$$PBLH = Ri_c \; \frac{T_{srf} \mathcal{W}^2(PBLH)}{g(\theta(PBLH) - \theta_s)} \tag{36}$$

 $Ri_c$  is the critical Richardson number (set to 0.21 by default),  $\theta$  is the potential temperature,  $\theta_s$  is the potential temperature of the air near the surface.  $\theta_s$  is estimated as the potential temperature  $\theta_0$  in the first layer plus a scaled virtual temperature excess  $\theta_e$ :

$$\theta_s = \theta_0 + \theta_e$$

with

$$\theta_e = C \frac{Q_0}{w_s}$$

C is a coefficient, usually set to 6.5, and the velocity scale  $w_s$  is given by equations (34) and (35). In AtmoData, one searches for the lowest level  $k_s$  (starting from the first layer) where

$$\theta_{k_s} \ge \theta_s + Ri_c \; \frac{T_{srf} \mathcal{W}_{k_s}^2}{g z_{k_s}} \tag{37}$$

Then one performs the following interpolation:

$$PBLH = z_{k_s-1} + \frac{-\Delta_{k_s-1}}{\Delta_{k_s} - \Delta_{k_s-1}} (z_{k_s} - z_{k_s-1})$$
(38)

with

$$\Delta_k = \theta_k - \left(\theta_s + Ri_c \; \frac{T_{srf} \mathcal{W}_k^2}{gz_k}\right)$$

An additional parameterization is based on the Richardson number. PBLH is determined by the first level where the Richardson number Ribecomes greater than its critical value  $Ri_c$ . If

$$k_c = \min\{k | Ri_k > Ri_c\}$$

then

$$PBLH = z_{k_c-1} + \frac{Ri_c - Ri_{k_c-1}}{Ri_{k_c} - Ri_{k_c-1}} \left( z_{k_c} - z_{k_c-1} \right)$$

## 3 Deposition velocities

The deposition velocities are assumed to be in the form:

$$V_d = \frac{1}{R_a + R_b + R_c} \tag{39}$$

 $V_d$  is in m · s<sup>-1</sup>.  $R_a$  is the aerodynamic resistance,  $R_b$  is the quasi-laminar sublayer resistance, and  $R_c$  is the canopy resistance. In general the efforts are dedicated to the calculation of  $R_c$ .

## 3.1 R<sub>a</sub>

The value of  $R_a$  can be either diagnosed according to wind and stability, or parameterized with the momentum flux or the heat flux (function of the Richardson number).

## 3.1.1 Diagnosis with the drag coefficient

The (rough) formula for  $R_a$  is

$$R_a = \frac{1}{C_D \max(0.1, \mathcal{W})} \tag{40}$$

 $\mathcal{W}$  is the wind module and  $C_D$  is the drag coefficient of momentum. It is computed with

$$C_D = r_0 \psi \tag{41}$$

and

$$r_0 = \frac{\kappa^2}{\ln \frac{25}{Z_r}}$$

 $Z_r$  is the roughness height.  $\psi$  is the stability function:

$$\psi = \begin{cases} 1 - 9.4 \frac{Ri}{1 + r_1 \sqrt{-Ri}} &, Ri < 0\\ (1 + 4.7Ri)^2 &, Ri > 0 \end{cases}$$
(42)

with

$$r_1 = 9.4 \times 7.4 r_0 \sqrt{\frac{25}{Z_r}}$$

The Richardson number is taken at the surface.

$$ComputeRa\_drag$$

## 3.1.2 Using the momentum flux

First, to estimate the momentum flux, we introduce:

$$f_m = \begin{cases} 1 - 2b_1 \frac{R_i}{1 + 3c\sqrt{-Ri}} &, Ri < 0\\ \left(1 + 2b_1 \frac{R_i}{\sqrt{1 + d_1Ri}}\right)^{-1} &, Ri > 0 \end{cases}$$
(43)

where  $b_1 = d_1 = 5$ . The value of c is more complicated:

$$c = b_1 c_1 A_u A_t \sqrt{1 - \frac{Z_t}{z_0 + Z_t}} \left( \left(\frac{z_0 + Z_t}{Z_t}\right)^{1/3} - 1 \right)^{3/2}$$

where  $c_1 = 5$  and  $Z_t$  is the roughness length for heat assumed to be equal to  $\frac{Z_r}{10}$ .

$$A_u = \frac{\kappa}{\ln\left(\frac{z_0 + Z_r}{Z_r}\right)}$$

and  $A_t$  has the same expression where  $Z_r$  is replaced with  $Z_t$ . Then

$$R_a = \left(A_u^2 f_m \mathcal{W}\right)^{-1}$$
ComputeRa\_momentum
(44)

#### 3.1.3 Using the heat flux

The heat flux is

$$f_{h} = \begin{cases} 1 - 3b_{1} \frac{R_{i}}{1 + 3c\sqrt{-Ri}} &, Ri < 0\\ \left(1 + 3b_{1} \frac{R_{i}}{1 + d_{1}Ri}\right)^{-1} &, Ri > 0 \end{cases}$$
(45)

The coefficients  $b_1$ , c and  $d_1$  are the same as for the momentum flux.

Then

$$R_a = (A_u A_t f_h \mathcal{W})^{-1}$$
ComputeRa\_heat
$$(46)$$

## 3.2 R<sub>b</sub>

 $R_b$  can also be diagnosed as  $R_a$ , or is computed as function of the friction velocity.

#### 3.2.1 Diagnosis with the drag coefficient

The formula is

$$R_b = \frac{2}{\kappa} \left(\frac{\nu \sqrt{M_i/18}}{D_e P r}\right)^{2/3} \frac{1}{C_D \max(0.1, \mathcal{W})}$$
(47)

where  $\nu$  is the kinematic viscosity of air, set to 0.15,  $M_i$  is the molar mass of the species for which  $R_b$  is calculated,  $D_e$  is the molecular diffusivity of water, set to 0.25, and Pr is the Prandtl number, set to 0.74.

$$ComputeRb_drag$$

## 3.2.2 Using the friction velocity

In this case

$$R_b = \frac{1}{\kappa u_*} \left(\frac{Sc(s)}{Pr}\right)^{2/3}$$
(48)

where Sc(s) is the Schmidt number for species s, which is the ratio of the kinematic viscosity of air upon the molecular diffusivity  $D_s$  of the species.

$$Sc(s) = \frac{\nu}{D_s}$$

The formula 48 is multiplied by two in presence of vegetation.

$$ComputeRb_friction$$

## 3.3 R<sub>c</sub>

Two parameterizations are proposed to compute  $R_c$  for a given species s.  $H_s$  is the Henry constant of this species,  $f_s$  its reactivity and  $D_s$  its molecular diffusivity.

#### 3.3.1 Wesely's parameterization

Warning: in this section, temperatures are in degrees Celsius.

This parameterization is described in [14].  $R_c$  is divided into four parallel resistances:

$$\frac{1}{R_c} = \frac{1}{R_{stom} + R_m} + \frac{1}{R_{lu}} + \frac{1}{R_{dc} + R_{cl}} + \frac{1}{R_{ac} + R_{gs}}$$
(49)

•  $R_{stom}$  is the stomatal resistance, computed with

$$R_{stom} = Ri \left( 1 + \frac{200}{SR + 0.1} \right)^2 \frac{400}{T_{srf}(40 - T_{srf})} \frac{D_s}{D_e}$$
(50)

if  $0^{\circ}C < T_{srf} < 40^{\circ}C$ . Otherwise it is set to  $\infty$ .  $T_{srf}$  is the surface temperature, and SR the solar radiation.

•  $R_m$  is the mesophyll resistance

$$R_m = \frac{1}{H_s/3000 + 100f_s} \tag{51}$$

•  $R_{lu}$  is the cuticle resistance

$$R_{lu} = \frac{R_{LUC}}{10^{-5}H_s + f_s} \tag{52}$$

where  $R_{LUC}$  is the leaf cuticle resistance for the current land cover (*LUC* means "land use coverage"). Note that typical values of  $R_{LUC}$  are provided in [14].

•  $R_{dc}$  is the buoyant convection (at sunlight) resistance

$$R_{dc} = \frac{100}{1 + \frac{1000}{SR + 10}} \tag{53}$$

• Finally  $R_{cl}$ , (the resistance of the exposed surfaces in the lowest structures above the ground) and  $R_{gs}$  (resistance to uptake at the ground) are computed as function of the values for O<sub>3</sub> and SO<sub>2</sub> (notably provided in [14]):

$$\frac{1}{R_{cl,gs}^{s}} = \frac{\alpha_s}{R_{cl,gs}^{O_3}} + \frac{\beta_s}{R_{cl,gs}^{SO_2}}$$
(54)

•  $R_{ac}$  (for transfers that depend only on the canopy height) is given in tables (in [14]).

ComputeWesely

#### 3.3.2 Zhang's parameterization

Warning: in this section, temperatures are in degrees Celsius.

A detailed description of this parameterization may be found in [15]. The whole resistance is

$$\frac{1}{R_c} = \frac{1 - W_{st}}{R_{stom} + R_m} + \frac{1}{R_{cut}} + \frac{1}{R_{ac} + R_g}$$
(55)

•  $W_{st}$  is the fraction of stomatal blocking, which is added to the model in presence of dew or rain

$$\begin{cases} 0 & , SR \le 200 \text{ W.m}^{-2} \\ \frac{SR-200}{800} & , 200 < SR \le 600 \text{ W.m}^{-2} \\ 0.5 & , 600 < SR \end{cases}$$
(56)

•  $R_{stom}$  is the stomatal resistance. His parameterization is quite complicated

$$R_{stom} = (G_s(PAR)f_{cr}(T)f_{cr}(\mathcal{D})f_{cr}(\psi) \ fracD_sD_e)^{-1}$$
(57)

First,  $G_s$  is the unstressed canopy stomatal conductance, depending on the photosynthetically active radiations (PAR). It may be computed as proposed in [16], through the formula

$$G_s(PAR) = \frac{F_{sun}}{r_{st}(PAR_{sun})} + \frac{F_{shade}}{r_{st}(PAR_{shade})}$$
(58)

 $F_{sun}$  and  $F_{shade}$  are the total sunlit and shaded leaf area indexes given by:

$$F_{sun} = \left(1 - e^{\frac{LAI}{2\cos Z}}\right) 2\cos Z$$
$$F_{shade} = LAI - F_{sun}$$

Z is the zenith angle, and LAI the leaf area index (given in the tables). If  $Z > 90^{\circ}$ ,  $F_{sun}$  and  $F_{shade}$  are set to 0.

 $r_{st}$  is the unstressed leaf stomatal resistance, given by equation 59, depending on  $PAR_{sun}$  and  $PAR_{shade}$ , the PAR received by sunlit and shaded leaves (equations 60 and 61):

$$r_{st} = r_{stmin} \left( 1 + \frac{b_{rs}}{PAR} \right) \tag{59}$$

 $r_{stmin}$  depends on the LUC category, and  $b_{rs}$  is an empirical constant. Both are given in tables.

When LAI < 2.5 or SR < 200 W.m<sup>-2</sup>

$$PAR_{shade} = R_{diff} e^{\frac{1}{2}LAI^{0.7}} + 0.07R_{dir}(1.1 - 0.1LAI)e^{-\cos Z}$$
$$PAR_{sun} = R_{dir}\frac{\cos\alpha}{\cos Z} + PAR_{shade}$$
(60)

otherwise

$$PAR_{shade} = R_{diff} e^{\frac{1}{2}LAI^{0.8}} + 0.07R_{dir}(1.1 - 0.1LAI)e^{-\cos Z}$$
$$PAR_{sun} = R_{dir}^{0.8} \frac{\cos \alpha}{\cos Z} + PAR_{shade}$$
(61)

 $R_{diff}$  and  $R_{dir}$  are the downward visible radiation fluxes above the canopy from diffuse and direct-beam radiations, respectively. They are computed following [13].  $\alpha$  is the angle between the leaves and the sun, and has a mean value of 60° for a canopy assumed to have a spherical leaf angle distribution.

In equation 57, the  $f_{cr}$  factors are the conductance-reducing effects of temperature T, water-vapor-pressure deficit  $\mathcal{D}$  and water stress  $\psi$ , given by:

$$f_{cr}(T) = \frac{T - T_{min}}{T_{opt} - T_{min}} \left(\frac{T_{max} - T}{T_{max} - T_{opt}}\right)^{b_t}$$
(62)

with  $b_t = \frac{T_{max} - T_{opt}}{T_{opt} - T_{min}}$ .  $T_{min}$ ,  $T_{max}$  and  $T_{opt}$  depend on the LUC category, and are provided in tables.

$$f_{cr}(\mathcal{D}) = 1 - b_{vpd}\mathcal{D} \tag{63}$$

with  $\mathcal{D} = \frac{P_{vapor}^{sat} - P_{vapor}}{1000}$ .  $b_{vpd}$  is the water-vapor-pressure-deficit constant, given in tables.  $P_{vapor}^{sat}$  and  $P_{vapor}$ , the saturation and ambient water vapor pressure, are computed with:

$$P_{vapor} = \frac{q_{srf} P_{srf}}{q_{srf} + 0.622(1 - q_{srf})}$$
$$P_{vapor}^{sat} = 610.78 \ e^{\frac{17.2694T_{srf}}{T_{srf} + 237.29}}$$

And

$$f_{cr}(\psi) = \frac{\psi - \psi_{c_2}}{\psi_{c_1} - \psi_{c_2}}$$
(64)

with  $\psi = -0.72 - 0.0013SR$  in presence of tall vegetation. Otherwise  $\psi = 102(-0.043T_{srf})$ . If  $\psi > \psi_{c_1}$  (no water potential stress),  $f_{cr}(\psi)$  is set to 1.

Note that if the temperature is out of the limits  $T_{min}$  and  $T_{max}$  (usually at night),  $R_{st}$  is set to  $\infty$ .

- The mesophyll resistance  $R_m$  is given in the tables (as function of the species refer to [16]).
- $R_{ac}$ , the aerodynamic resistance in canopy is computed with

$$R_{ac} = \frac{R_{ac0} LA I^{1/4}}{{u_*}^2} \tag{65}$$

where the reference value  $R_{ac0}$  is to be found in tables, and depends on the LUC category.

• Finally the ground resistance  $R_g$  and the cuticular resistance  $R_{cut}$  are computed with equation 54.

Meanwhile, if  $T_{srf} < -1^{\circ}$ C, they are multiplied by an empirical coefficient:  $e^{-0.2(1+T_{srf})}$ .

ComputeZhang

## 4 Surface emissions

The surface emissions are divided in two parts: the anthropogenic emissions, that come from human activities (transport, industry, etc.), and the biogenic emissions, which are emitted by the vegetation.

## 4.1 Anthropogenic emissions

They are interpolated from  $EMEP^{\dagger 2}$  data to a grid with fixed steps along latitude and longitude. The EMEP emissions are adjusted so that, within a given EMEP cell, the ratios of urban emissions and forest emissions upon remaining emissions are equal to 1.6 and 1.2 respectively. They are also multiplied by monthly, daily (especially to adjust the saturday and sunday emissions) and hourly factors. The chemical species are handled as advocated in [7].

> ReadEmep, EmepToPolair, SpeciationAggregation, GridCorrespondences, ComputeVerticalDistribution

 $<sup>^{\</sup>dagger 2} \rm Co-operative$  Programme for Monitoring and Evaluation of the Long-range Transmission of Air Pollutants in Europe <code>-www.emep.int/</code>

#### 4.2**Biogenic** emissions

They are computed for three species: isoprene, terpenes and NO. As proposed in [10], the biogenic emissions are computed as follows:

$$E_{s,i,j} = \sum_{c=1}^{N_c} \gamma_{s,i,j}(c) \ d(c) \ EF_s(c) \ LUC_{i,j}(c)$$
(66)

where  $N_c$  is the number of LUC classes, d(c) is the foliar biomass density of the class c,  $EF_s(c)$  is the emission factor of the species s for the class c, and  $LUC_{i,j}(c)$  is the proportion (in [0,1]) of the class c in the cell.  $\gamma_{s,i,j}(c)$  is a correction factor depending on s, c, the temperature and the PAR in the  $\operatorname{cell}(i,j).$ 

For isoprene,  $\gamma = c_l * c_t$ , with

$$c_l = \frac{\alpha c_{l1} PAR}{\sqrt{1 + \alpha^2 PAR^2}}$$

where  $\alpha = 0.0027$  and  $c_{l1} = 1.066$ . And

$$c_{t} = \frac{e^{\frac{c_{t1}(T_{srf} - T_{r})}{RT_{r}T_{srf}}}}{1 + e^{\frac{c_{t2}(T_{srf} - T_{m})}{RT_{r}T_{srf}}}}$$

where  $T_r = 303$  K,  $T_m = 314$  K,  $c_{t1} = 95,000$  and  $c_{t2} = 230,000$ .  $T_{srf}$  is the surface temperature and R the gas constant.

For terpenes,  $\gamma = e^{0.09(T_{srf} - T_r)}$ . For NO  $\gamma = e^{0.071T_r^{NO}}$  where

$$T_r^{NO} = \begin{cases} 0.67(T_{srf} - 273.15) + 8.8 &, EF_{NO,c} = 0.9 \\ \\ 0.84(T_{srf} - 273.15) + 3.6 &, \text{ otherwise} \end{cases}$$

One can get the emission rates, too:

$$ER_{si,j} = \sum_{c=1}^{N_c} d(c) \ EF_s(c) \ LUC_{i,j}(c)$$
(67)

ComputeBiogenicEmissions, ComputeBiogenicRates

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